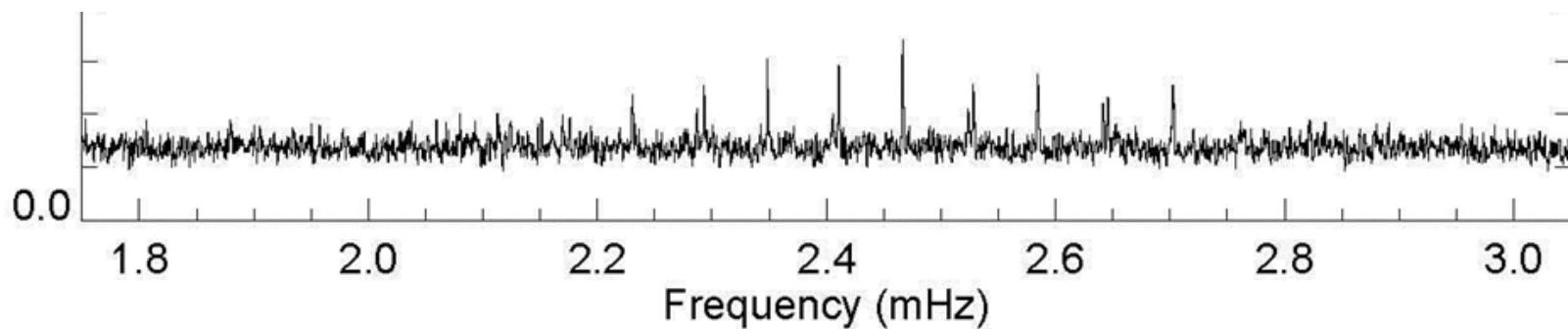
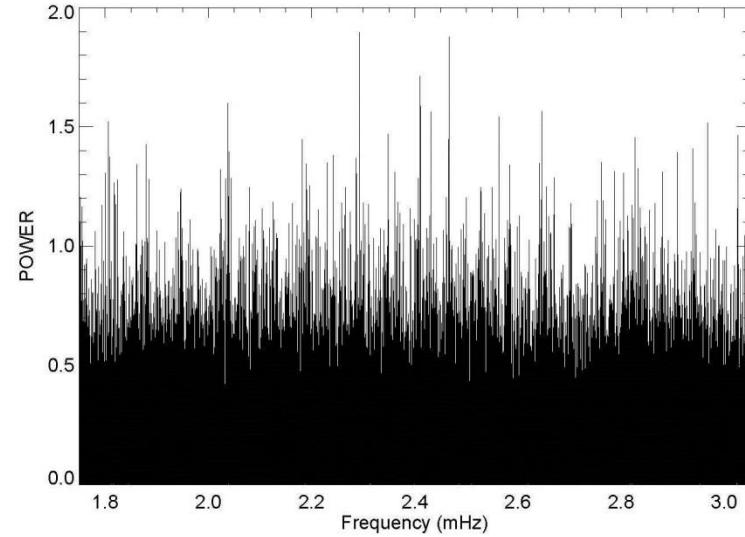
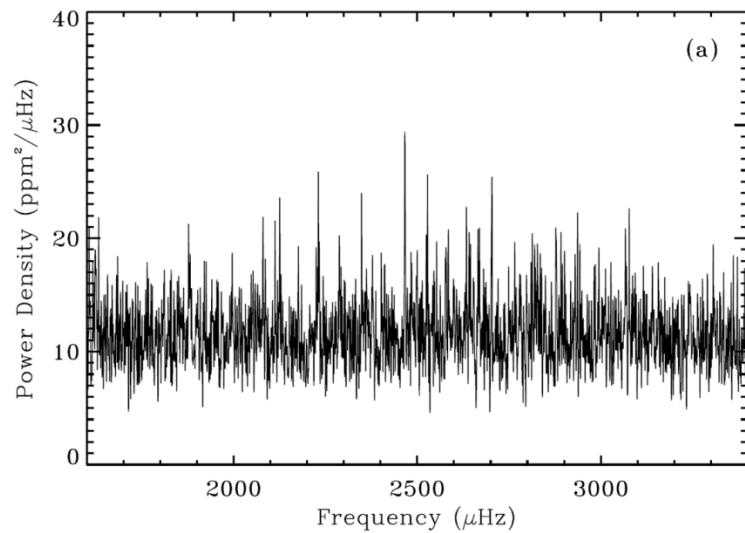


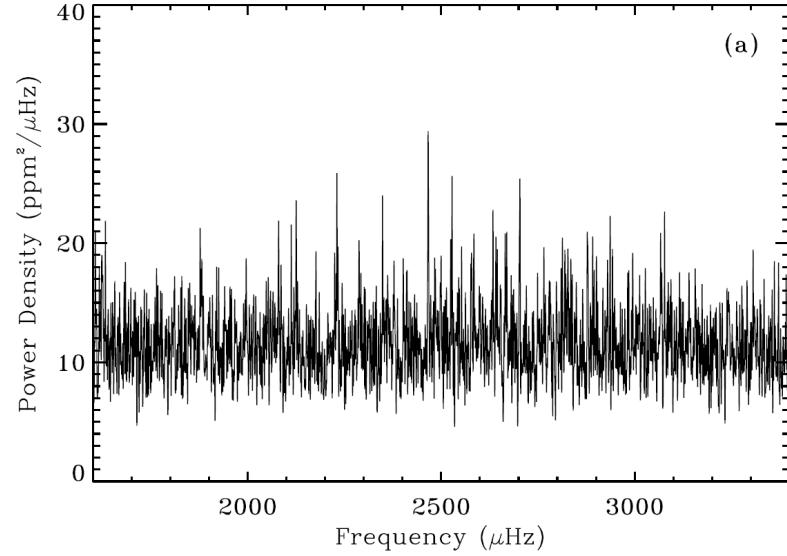
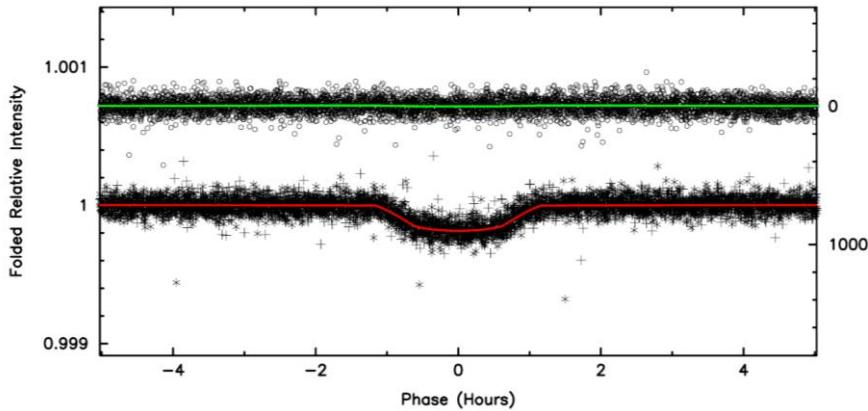
# Analysis of low SNR data

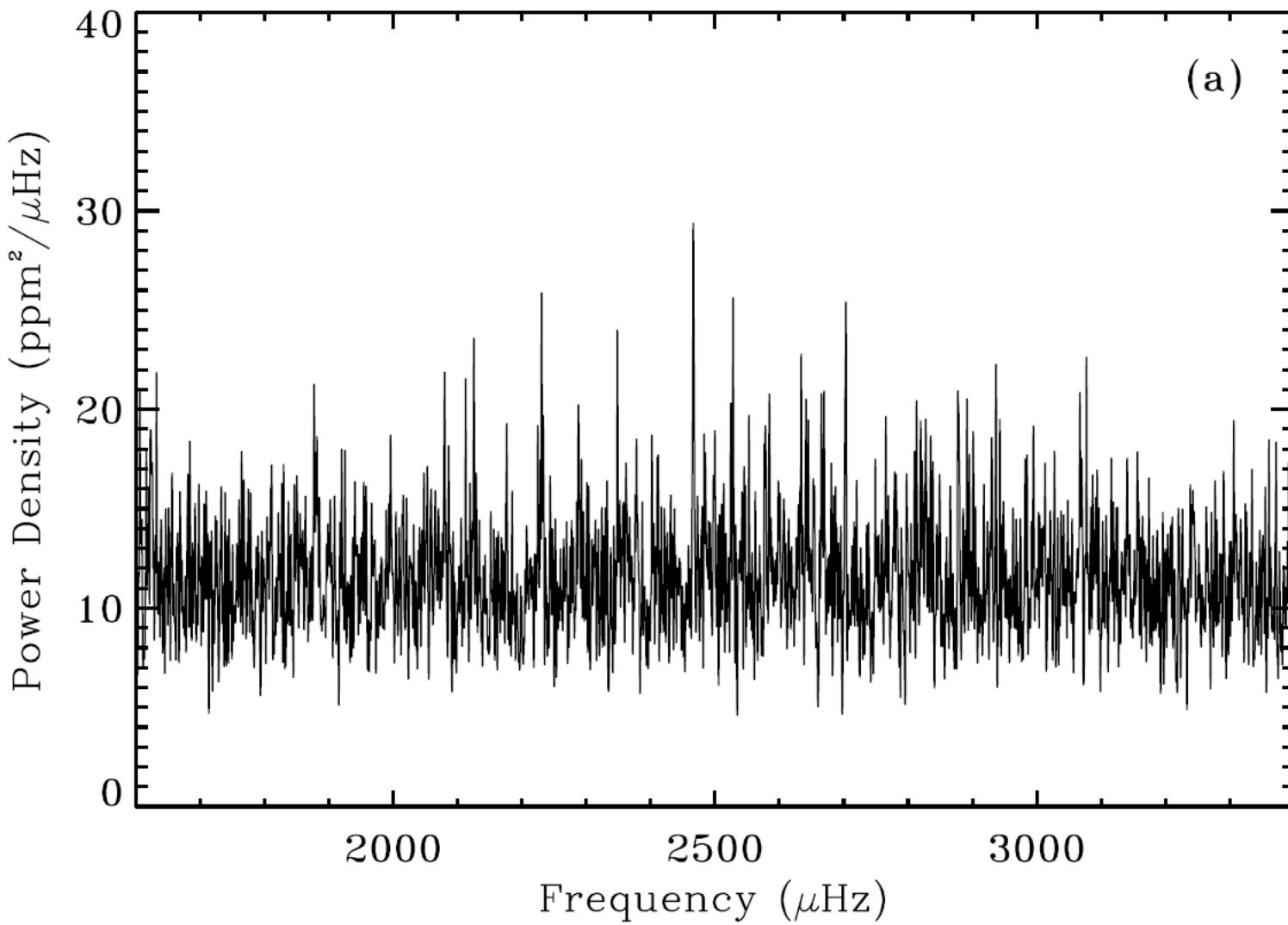
Hans Kjeldsen, Stellar Astrophysics Centre, Aarhus University, Denmark

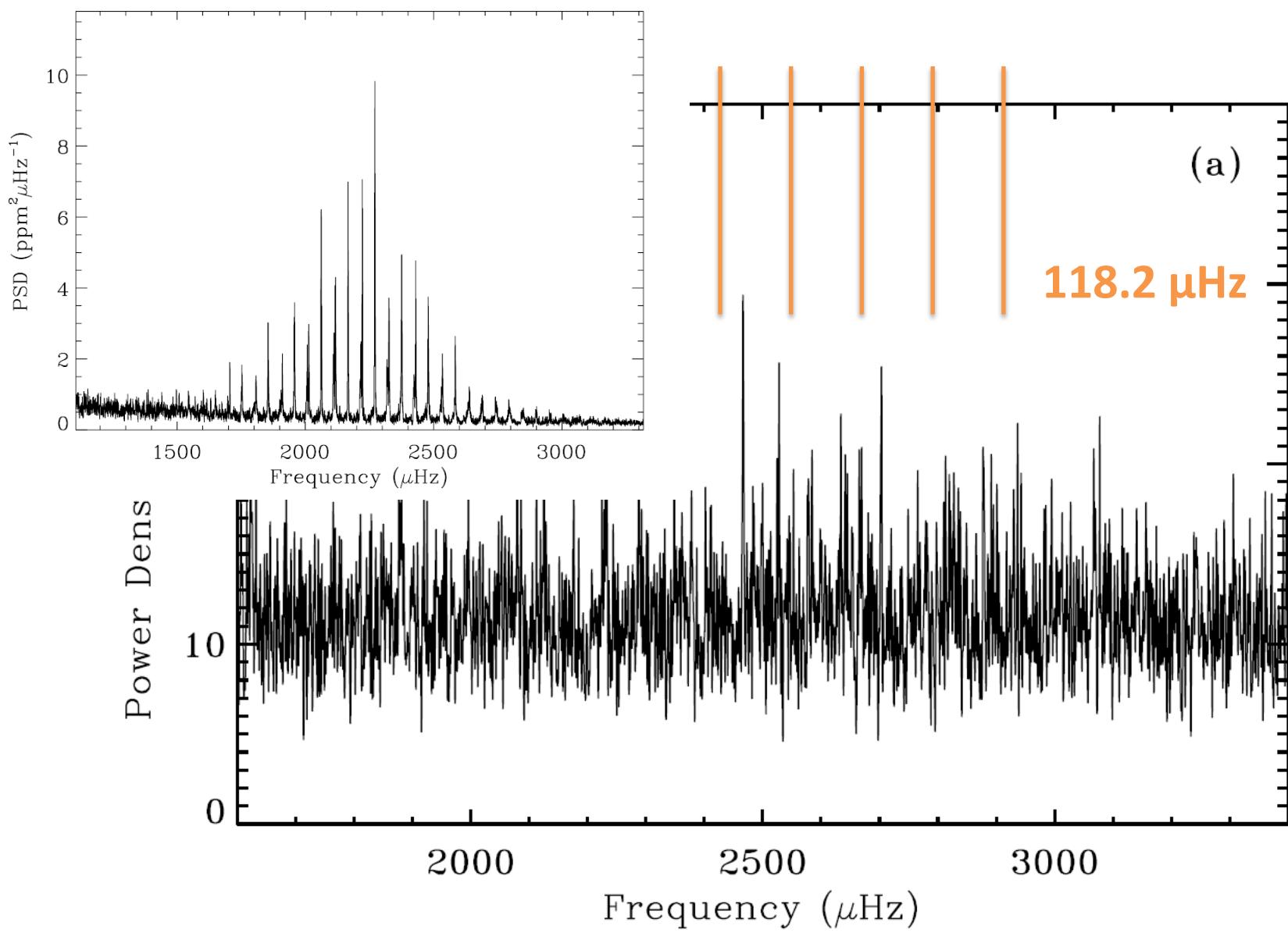


## KEPLER'S FIRST ROCKY PLANET: KEPLER-10b\*

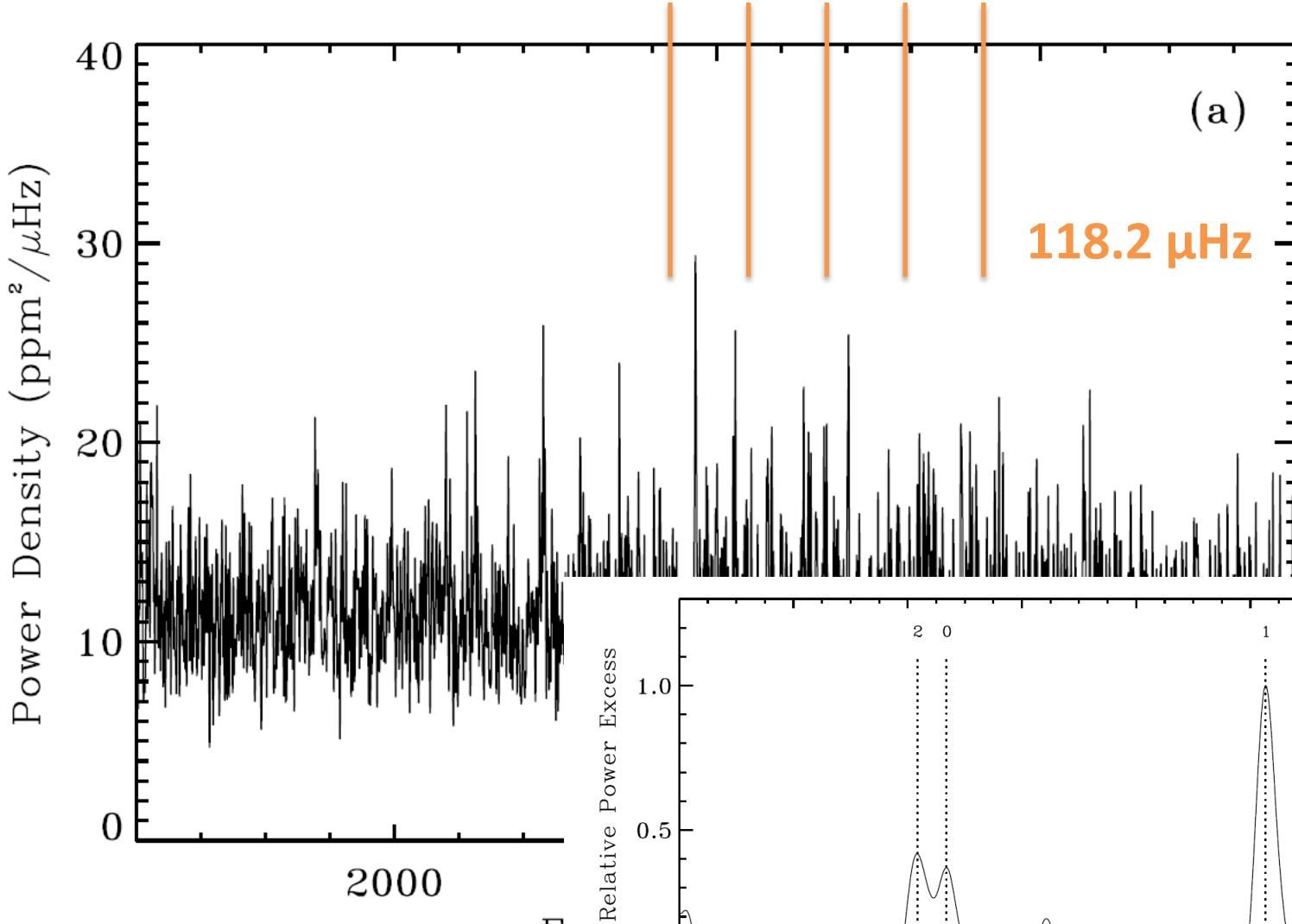
NATALIE M. BATALHA<sup>1</sup>, WILLIAM J. BORUCKI<sup>2</sup>, STEPHEN T. BRYSON<sup>2</sup>, LARS A. BUCHHAVE<sup>3</sup>, DOUGLAS A. CALDWELL<sup>4</sup>,  
JØRGEN CHRISTENSEN-DALSGAARD<sup>5,6</sup>, DAVID CIARDI<sup>7</sup>, EDWARD W. DUNHAM<sup>8</sup>, FRANCOIS FRESSIN<sup>3</sup>, THOMAS N. GAUTIER III<sup>9</sup>,  
RONALD L. GILLILAND<sup>10</sup>, MICHAEL R. HAAS<sup>2</sup>, STEVE B. HOWELL<sup>11</sup>, JON M. JENKINS<sup>4</sup>, HANS KJELDSSEN<sup>5</sup>, DAVID G. KOCH<sup>2</sup>,  
DAVID W. LATHAM<sup>3</sup>, JACK J. LISSAUER<sup>2</sup>, GEOFFREY W. MARCY<sup>12</sup>, JASON F. ROWE<sup>2</sup>, DIMITAR D. SASSELOV<sup>3</sup>, SARA SEAGER<sup>13</sup>,  
JASON H. STEFFEN<sup>14</sup>, GUILLERMO TORRES<sup>3</sup>, GIBOR S. BASRI<sup>12</sup>, TIMOTHY M. BROWN<sup>15</sup>, DAVID CHARBONNEAU<sup>3</sup>,  
JESSIE CHRISTIANSEN<sup>2</sup>, BRUCE CLARKE<sup>4</sup>, WILLIAM D. COCHRAN<sup>16</sup>, ANDREA DUPREE<sup>3</sup>, DANIEL C. FABRYCKY<sup>3</sup>, DEBRA FISCHER<sup>17</sup>,  
ERIC B. FORD<sup>18</sup>, JONATHAN FORTNEY<sup>19</sup>, FORREST R. GIROUARD<sup>20</sup>, MATTHEW J. HOLMAN<sup>3</sup>, JOHN JOHNSON<sup>21</sup>, HOWARD ISAACSON<sup>12</sup>,  
TODD C. KLAUS<sup>20</sup>, PAVEL MACHALEK<sup>4</sup>, ALTHEA V. MOOREHEAD<sup>18</sup>, ROBERT C. MOREHEAD<sup>18</sup>, DARIN RAGOZZINE<sup>3</sup>,  
PETER TENENBAUM<sup>4</sup>, JOSEPH TWICKEN<sup>4</sup>, SAMUEL QUINN<sup>3</sup>, JEFFREY VANCLEVE<sup>4</sup>, LUCIANNE M. WALKOWICZ<sup>12</sup>,  
WILLIAM F. WELSH<sup>22</sup>, EDNA DEVORE<sup>4</sup>, AND ALAN GOULD<sup>23</sup>







**Large frequency separation**

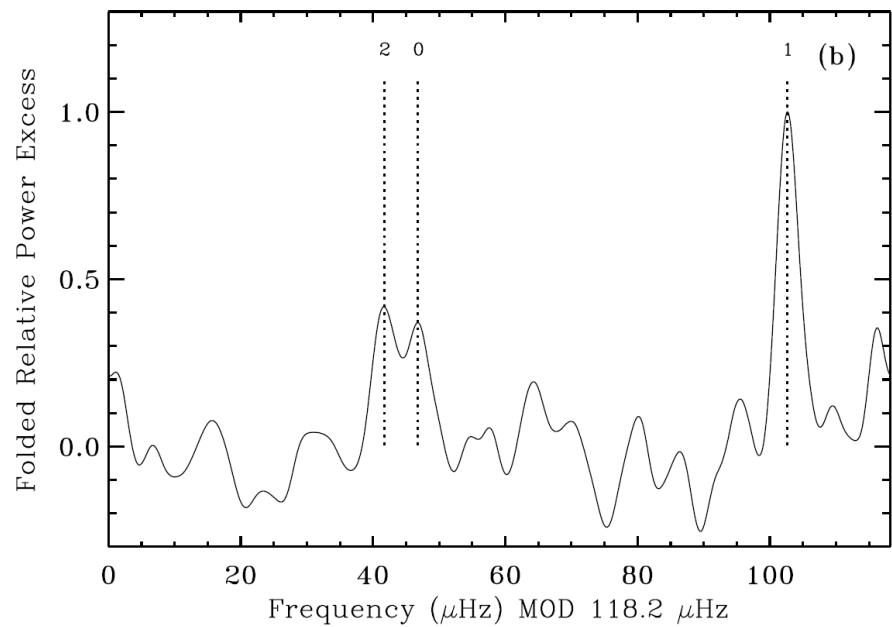


**Small frequency separation**

Frequency ( $\mu\text{Hz}$ ) MOD  $118.2 \mu\text{Hz}$

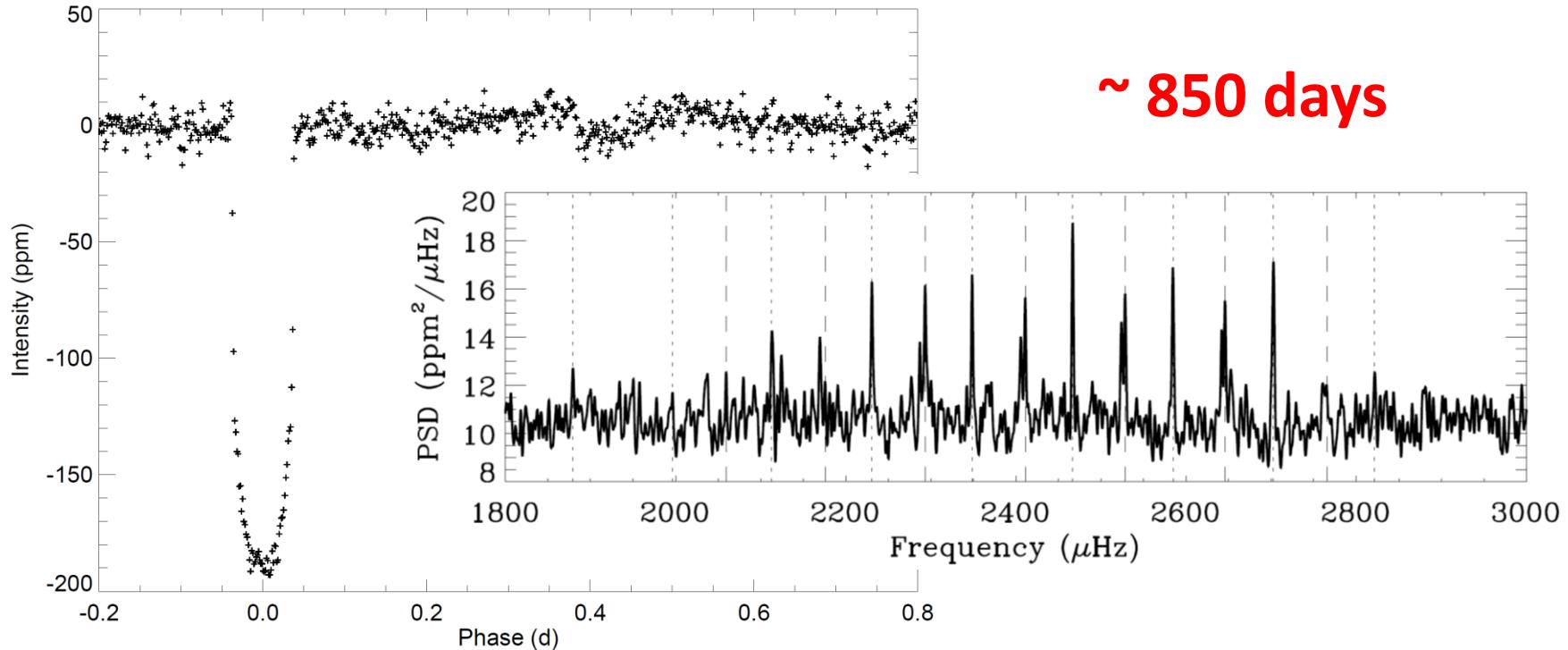
Mass (Msun)	$0.995 \pm 0.060$	( 6%)
Radius (Rsun)	$1.056 \pm 0.021$	( 2 %)
Age (Gyr)	$11.9 \pm 4.5$	(38%)

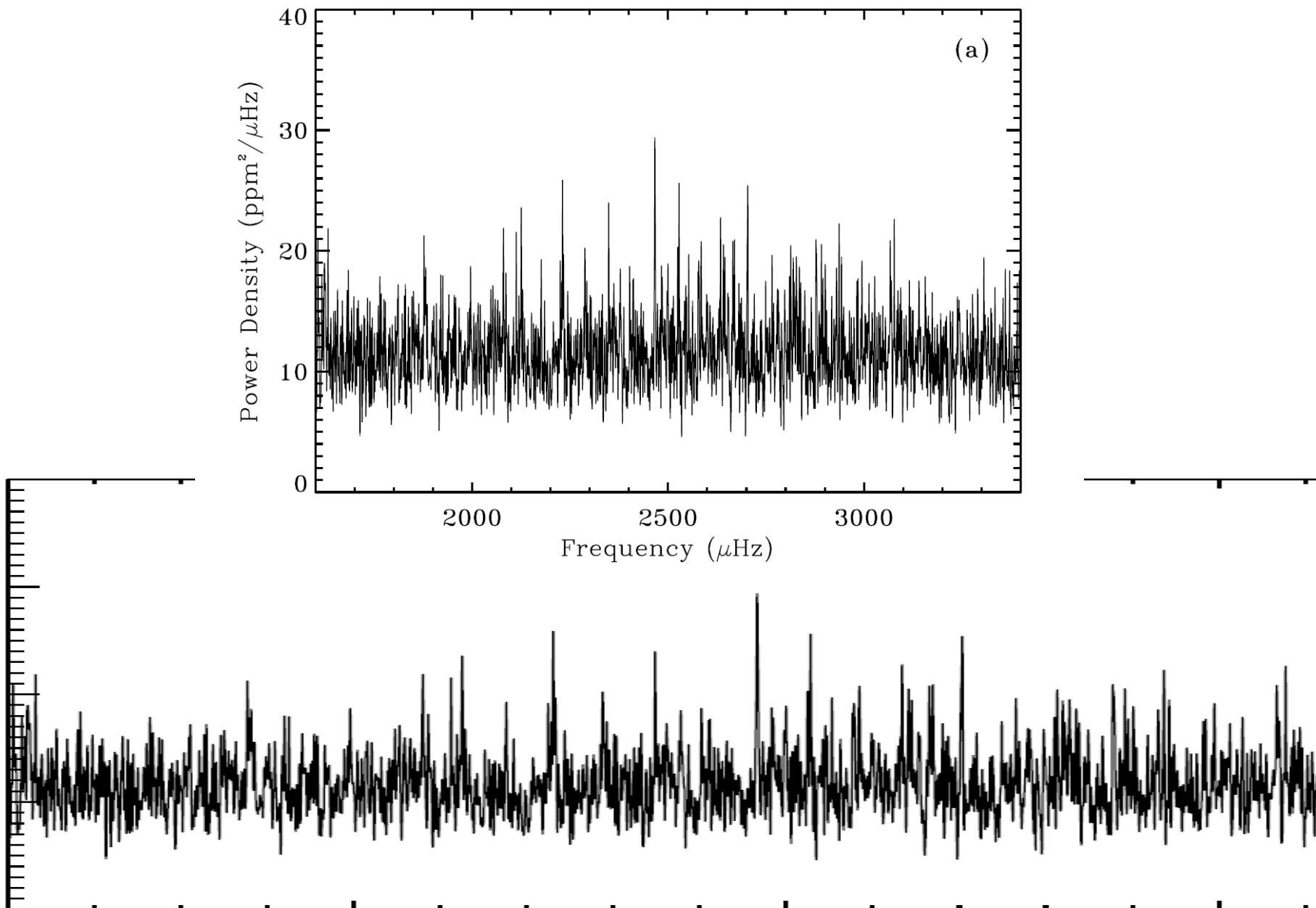
- Batalha et al. 2011

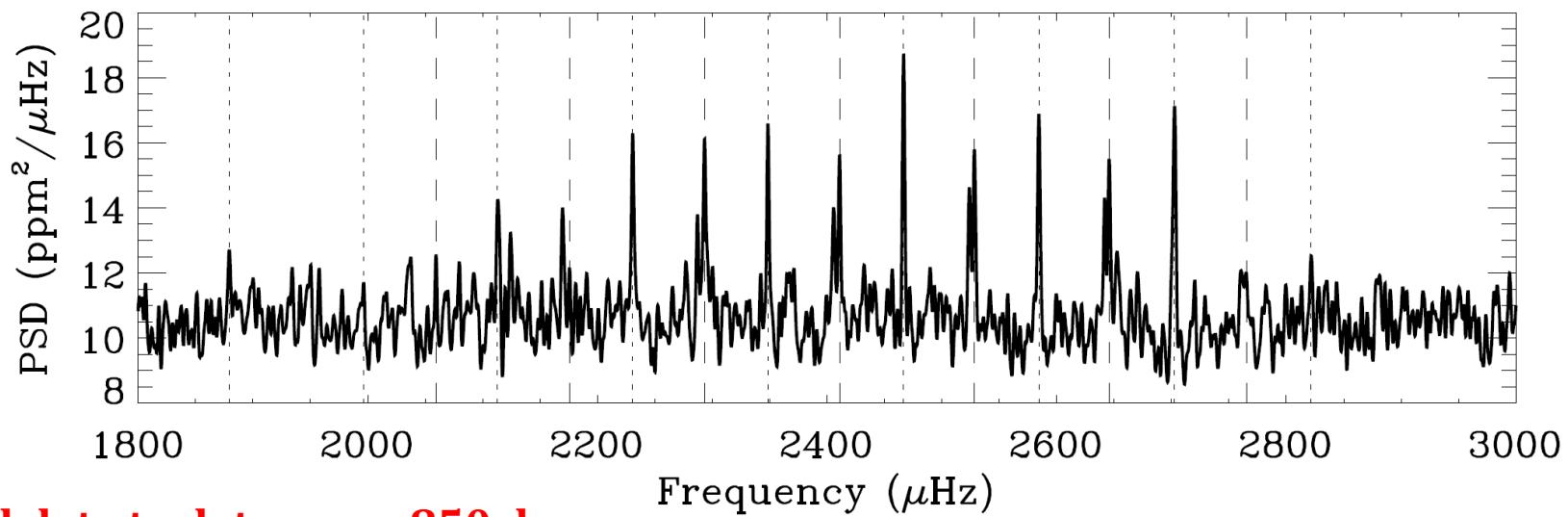


# Analysis of more than two years of data....

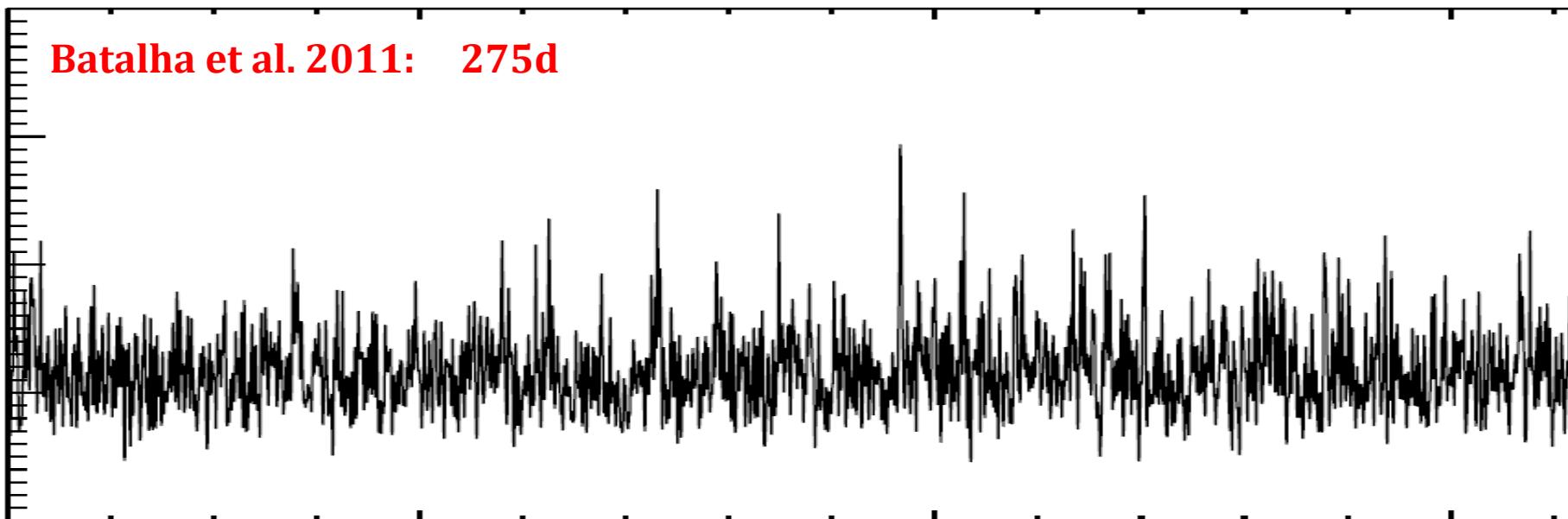
Alexandra Fogtmann-Schulz, Brian Hinrup, Vincent Van Eylen, Jørgen Christensen-Dalsgaard, Hans Kjeldsen, Victor Silva Aguirre and Brandon Tingley  
Stellar Astrophysics Centre, Department of Physics and Astronomy, Aarhus University, Denmark.



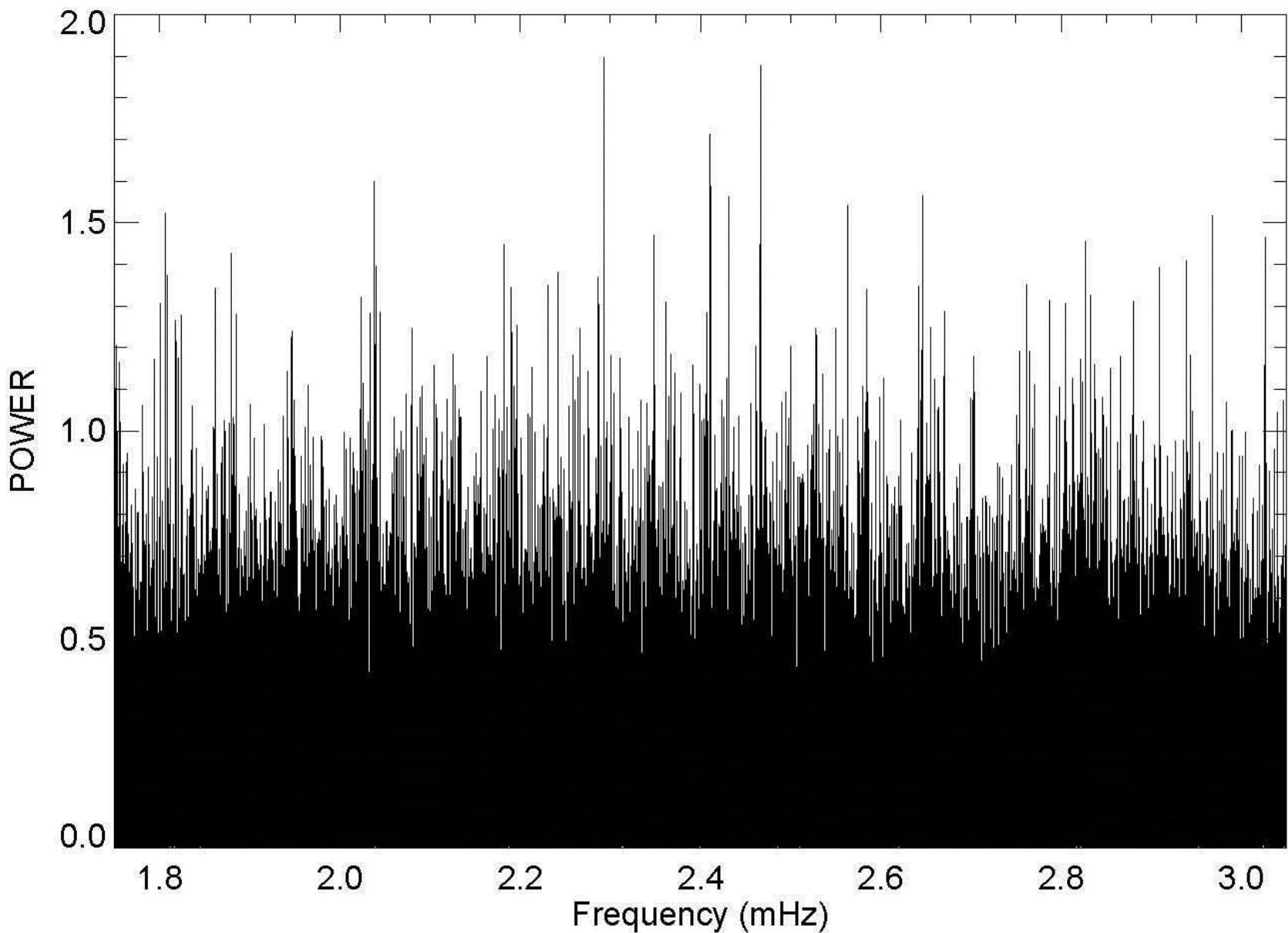


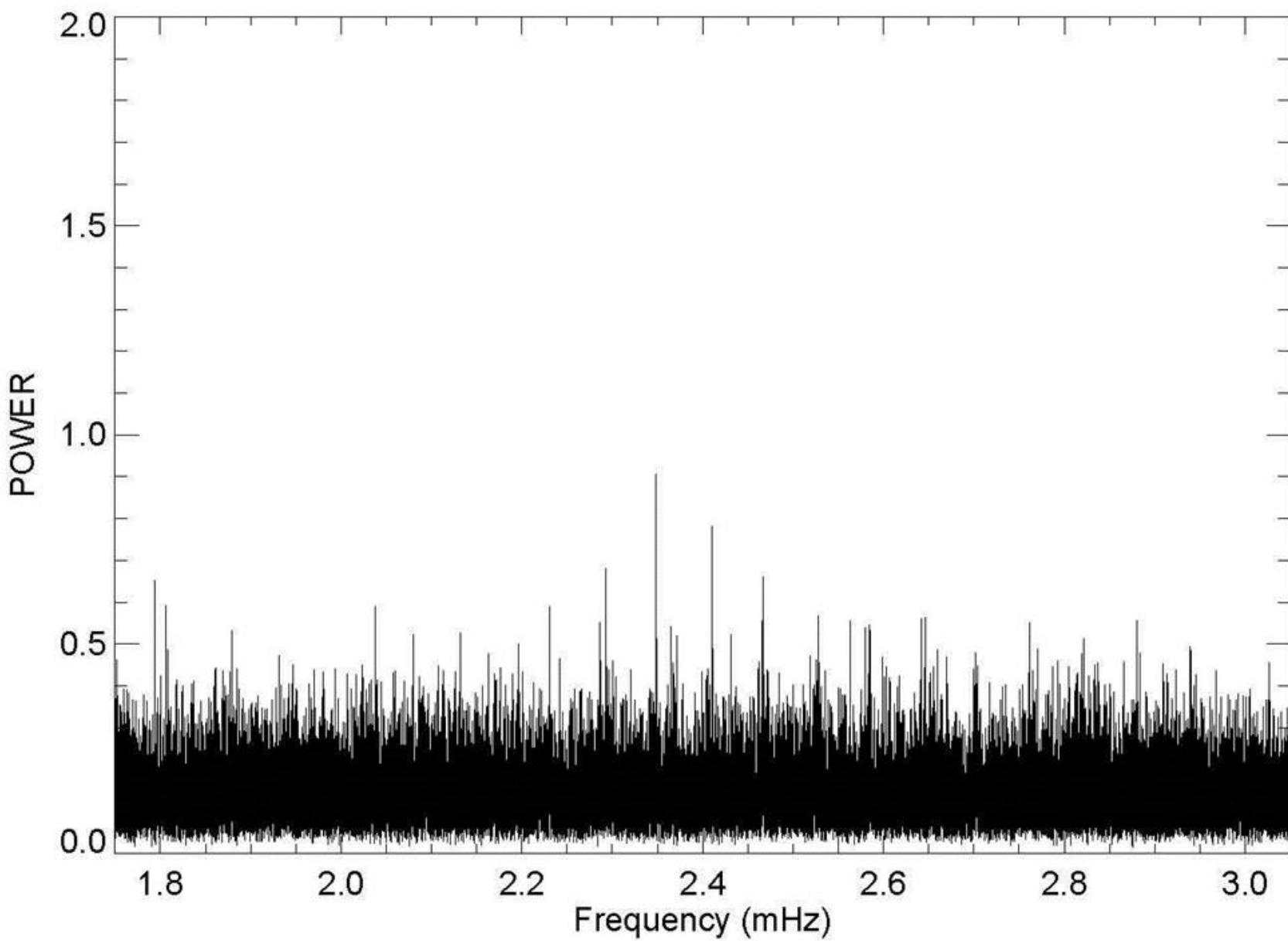


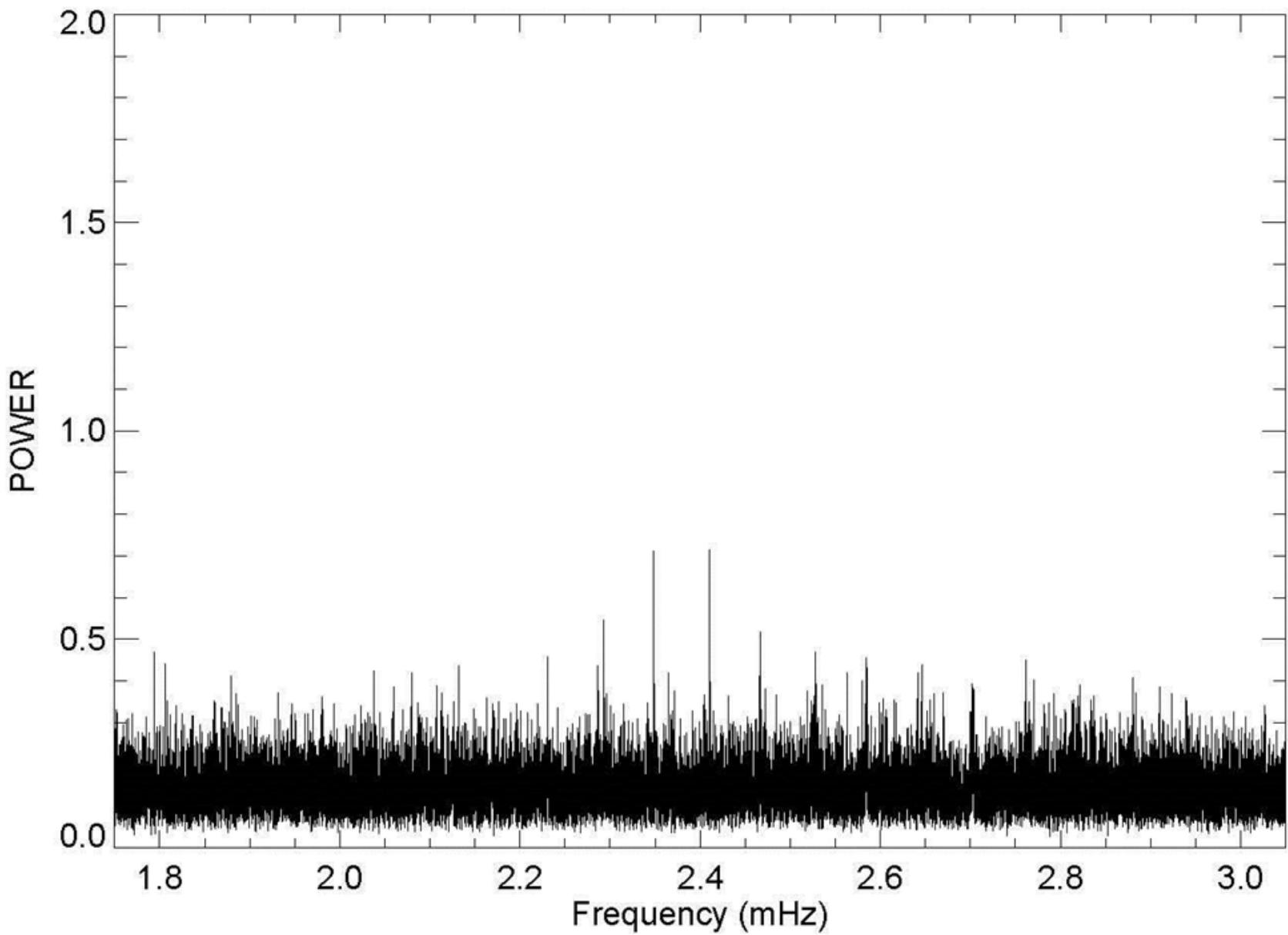
All data to date: **850 d**

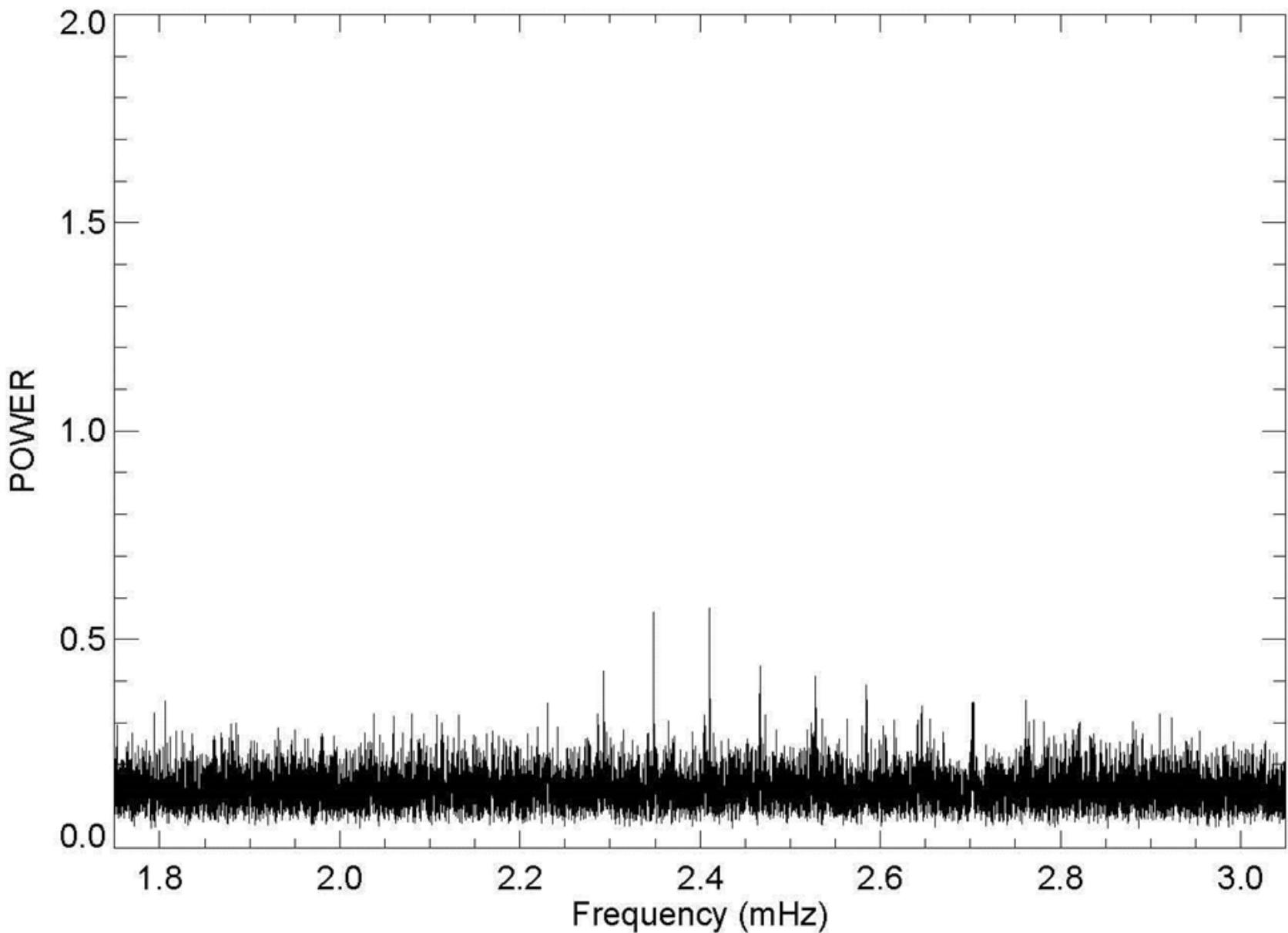


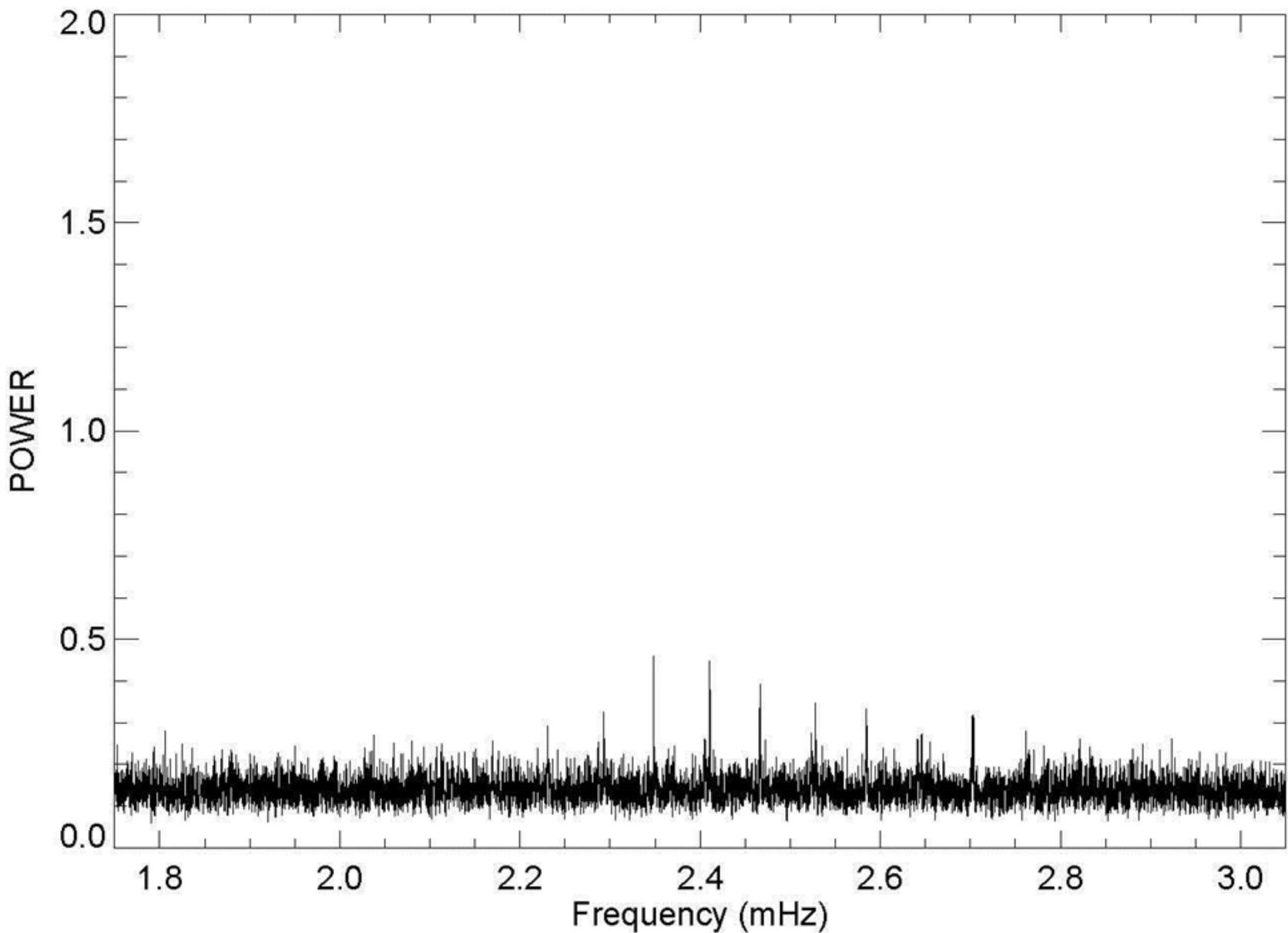
Batalha et al. 2011: **275d**

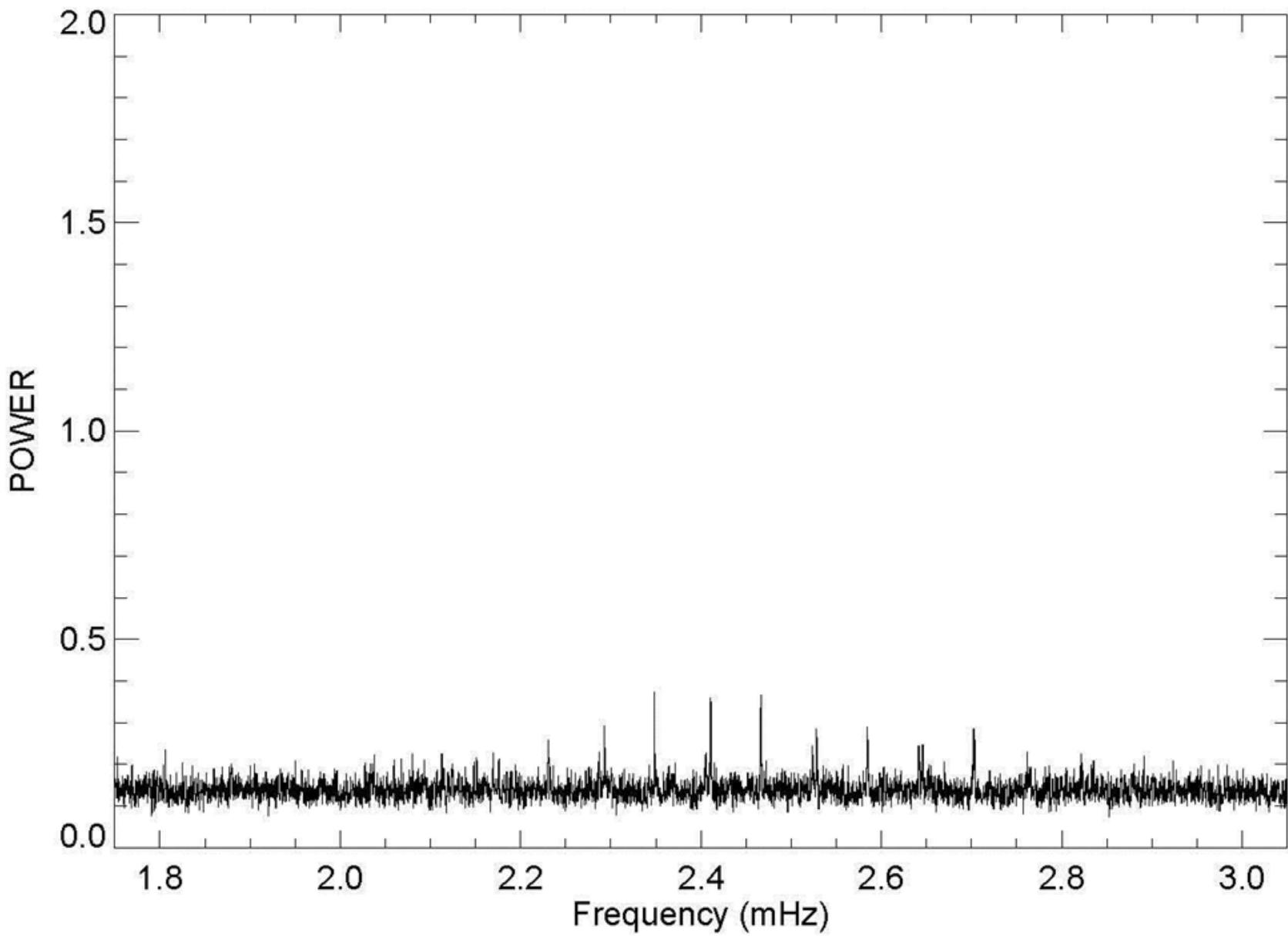


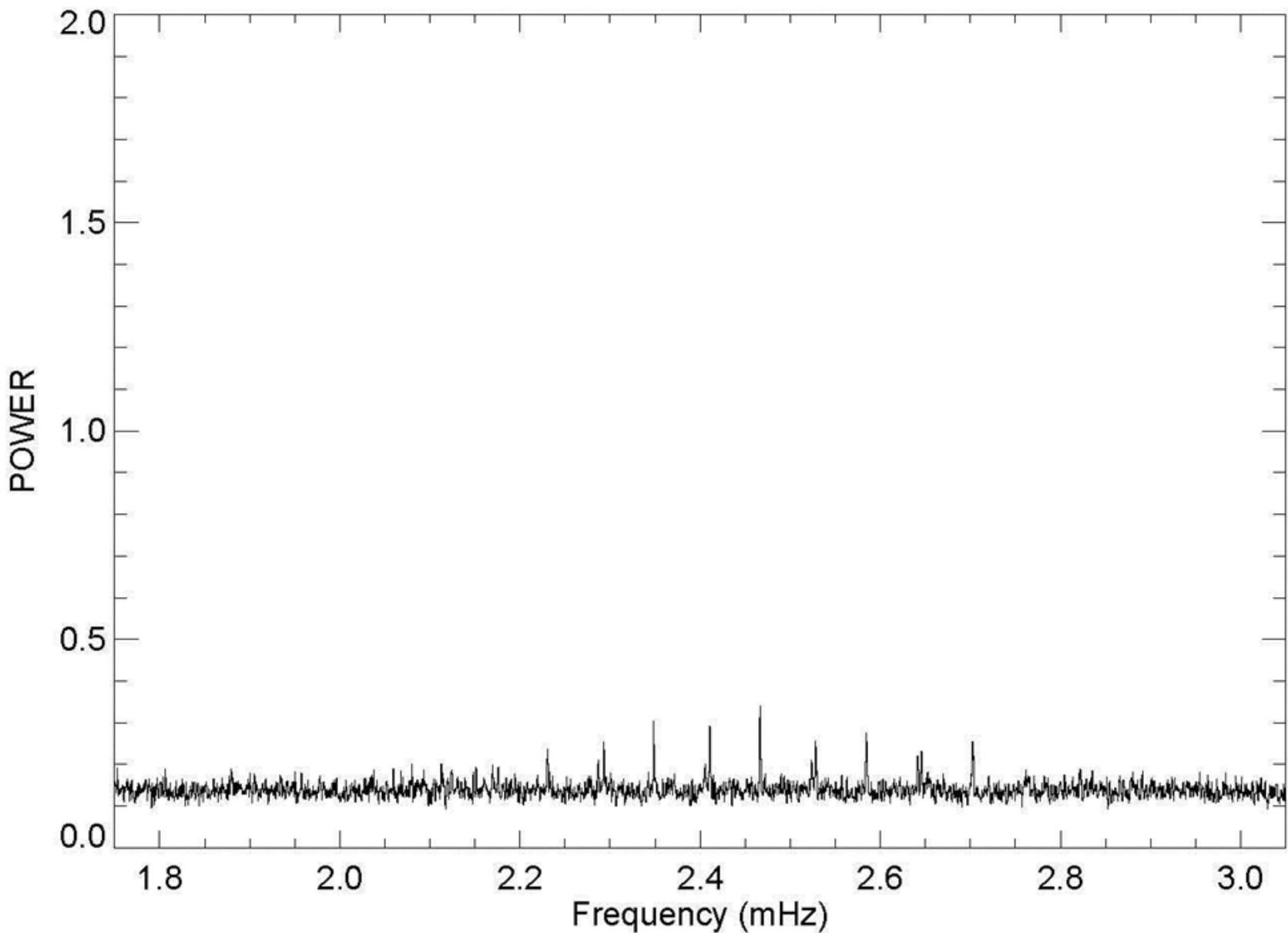


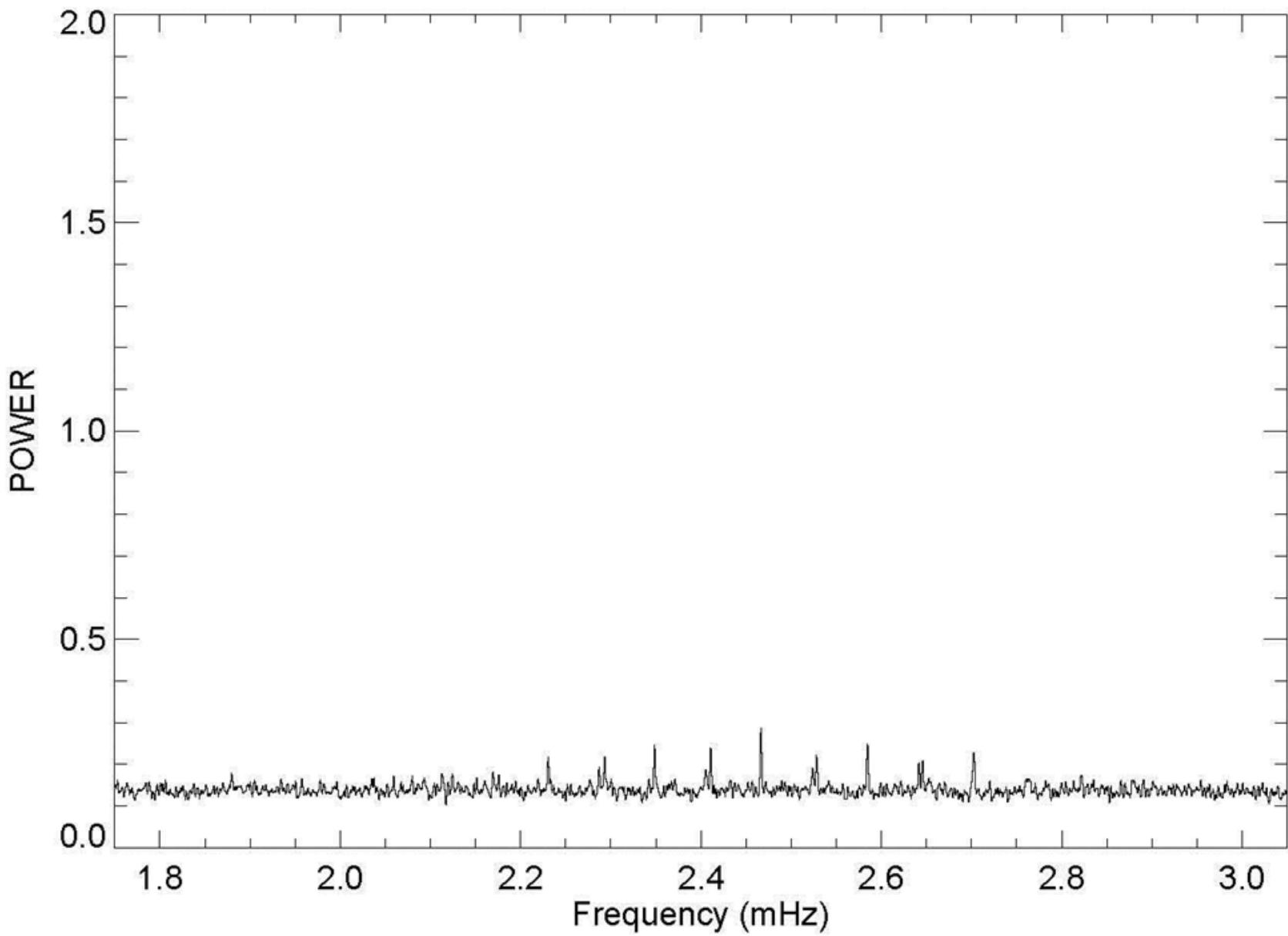


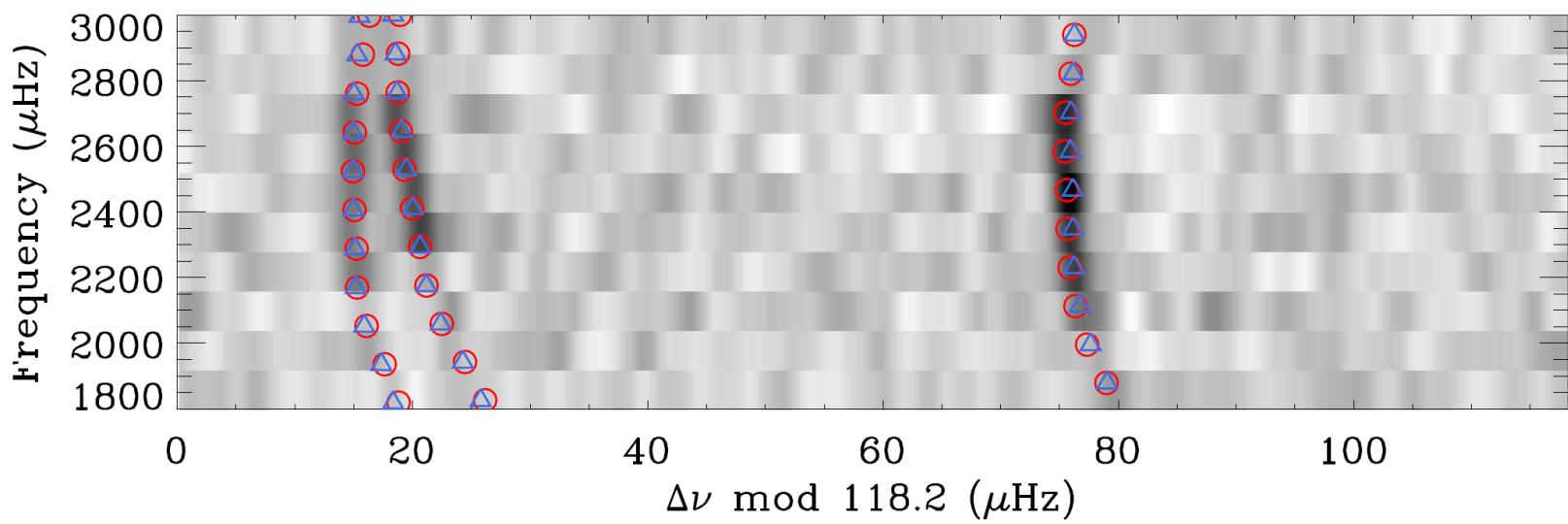
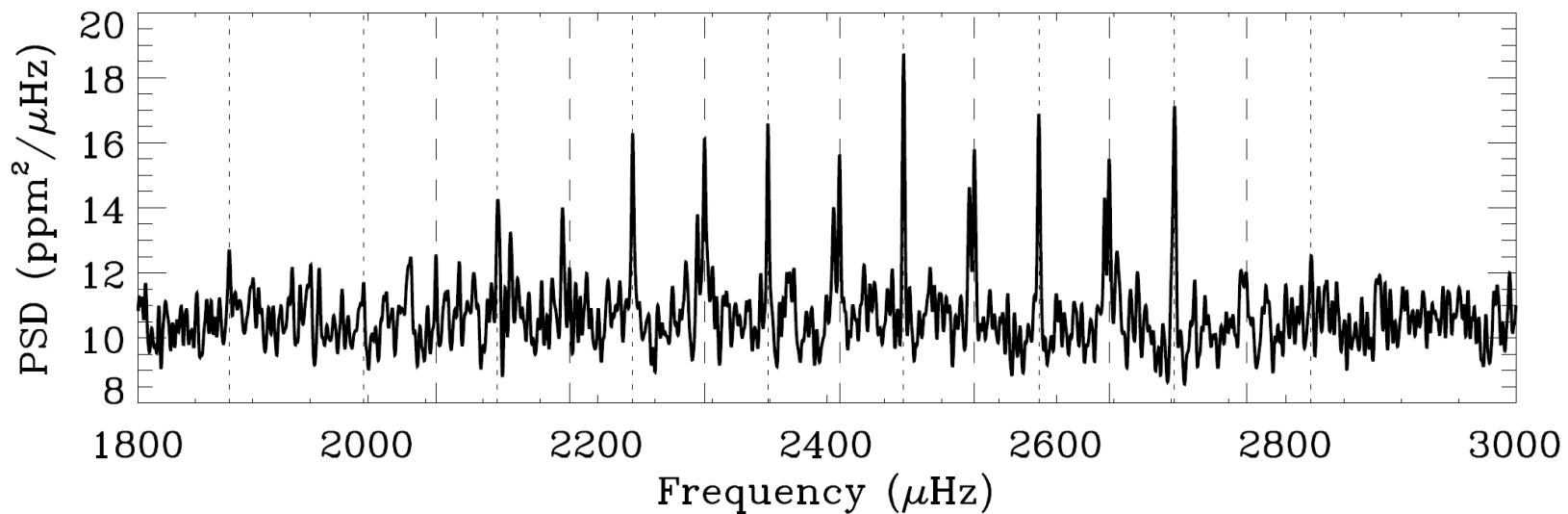










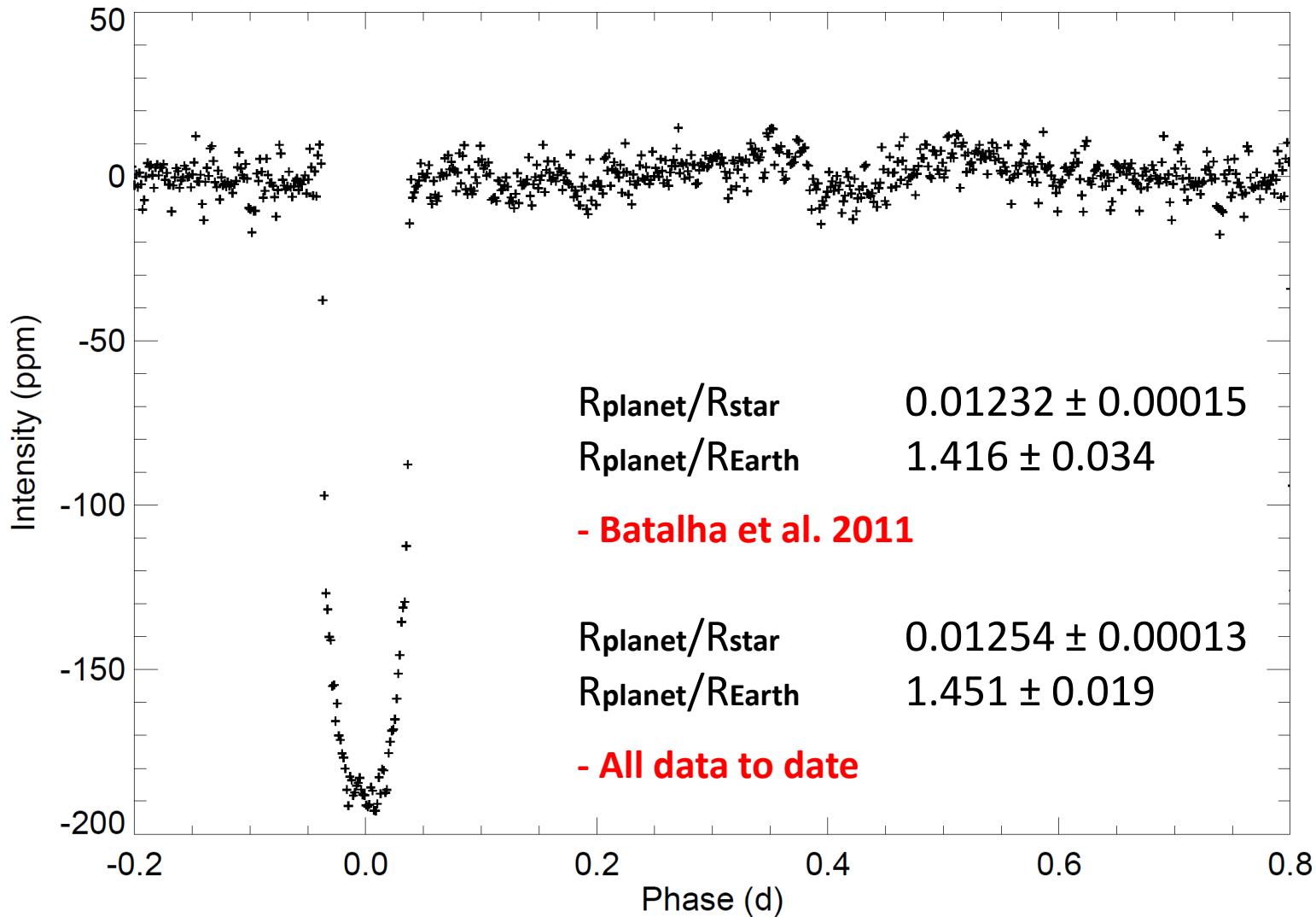


Mass (Msun)	$0.995 \pm 0.060$
Radius (Rsun)	$1.056 \pm 0.021$
Age (Gyr)	$11.9 \pm 4.5$

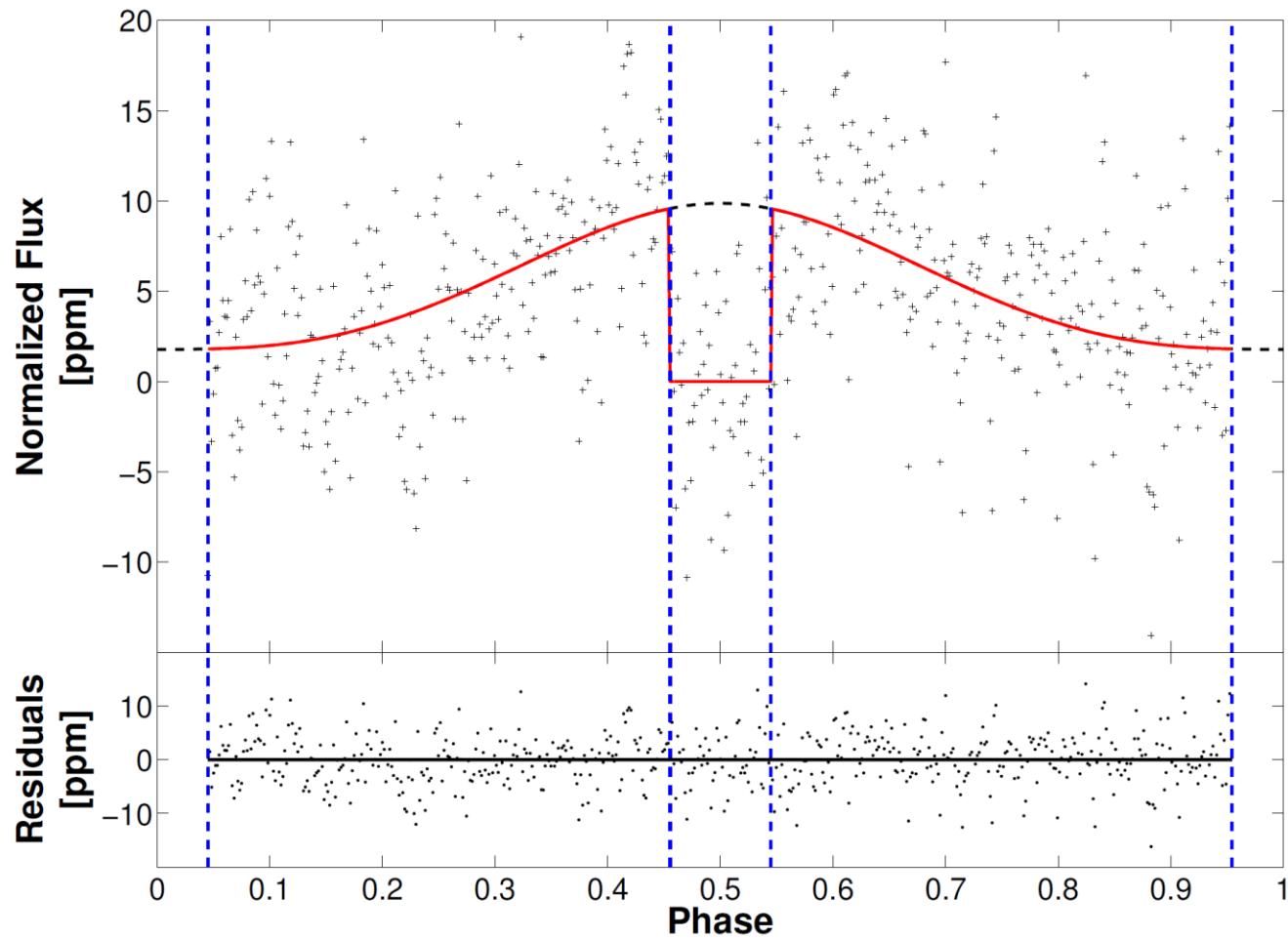
- **Batalha et al. 2011**

Mass (Msun)	$0.913 \pm 0.022$
Radius (Rsun)	$1.065 \pm 0.009$
Age (Gyr)	$10.4 \pm 1.4$

- **All data to date**



$$\delta_{\text{occultation}} \text{ (ppm)} \quad 9.9 \pm 1.0$$



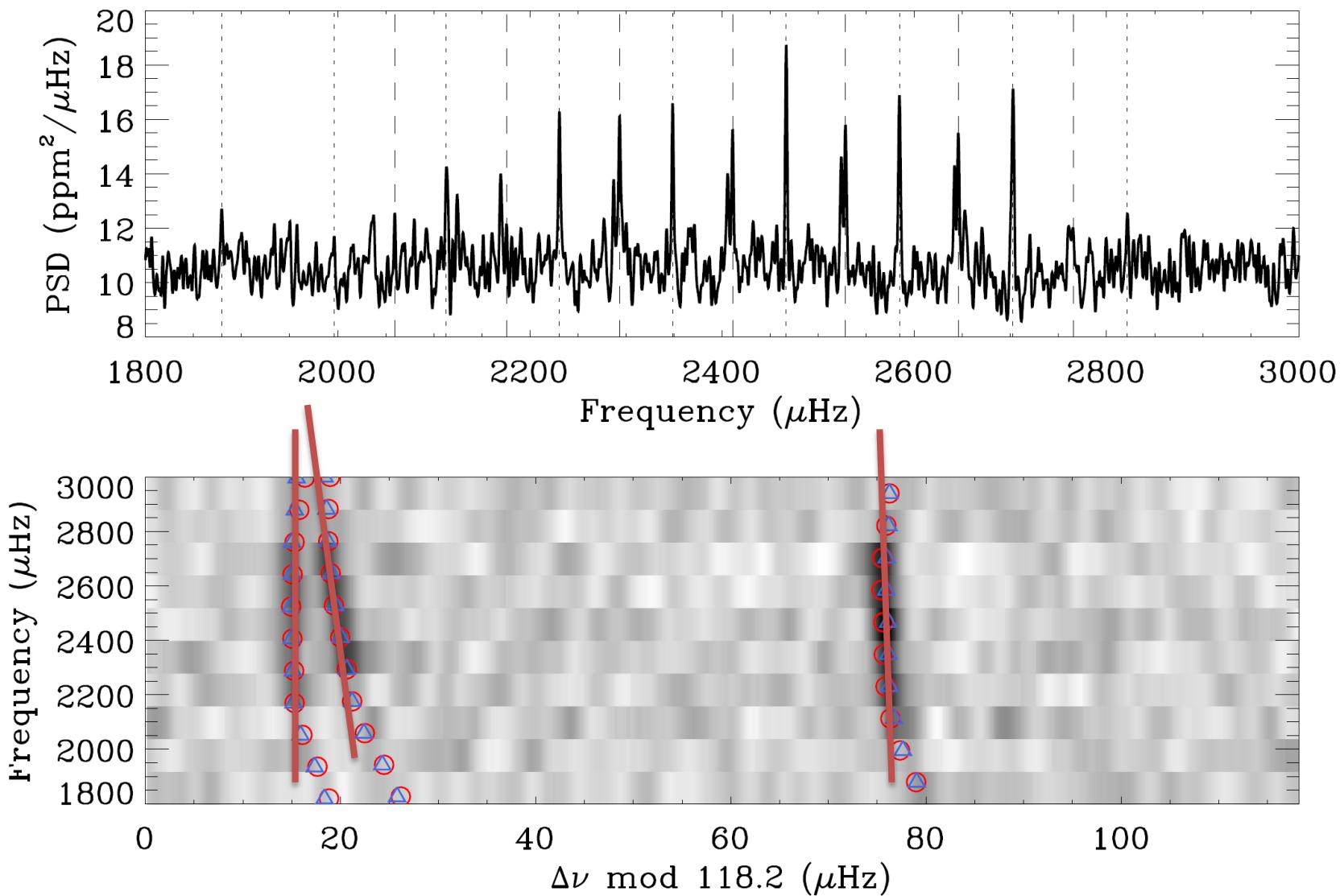
## Kepler-10:

Mass (Msun)	$0.913 \pm 0.022$	(2.4%)
Radius (Rsun)	$1.065 \pm 0.009$	(0.85%)
Age (Gyr)	$10.4 \pm 1.4$	(13%)

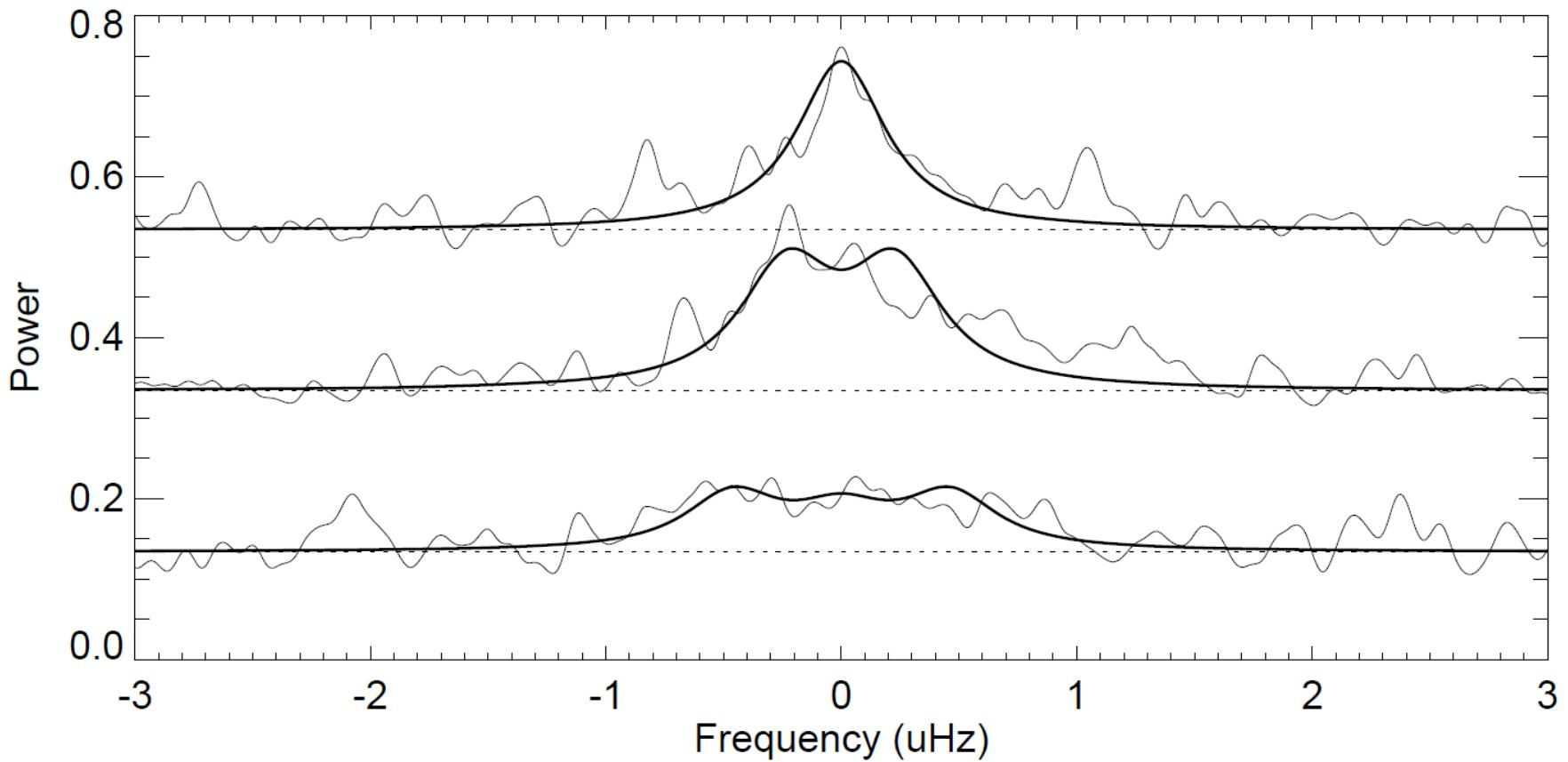
## Kepler-10b:

$R_{\text{planet}}/R_{\text{star}}$	$0.01254 \pm 0.00013$	(1.0%)
$R_{\text{planet}}/R_{\text{Earth}}$	$1.451 \pm 0.019$	(1.3%)

The key is to extend the length of the time series

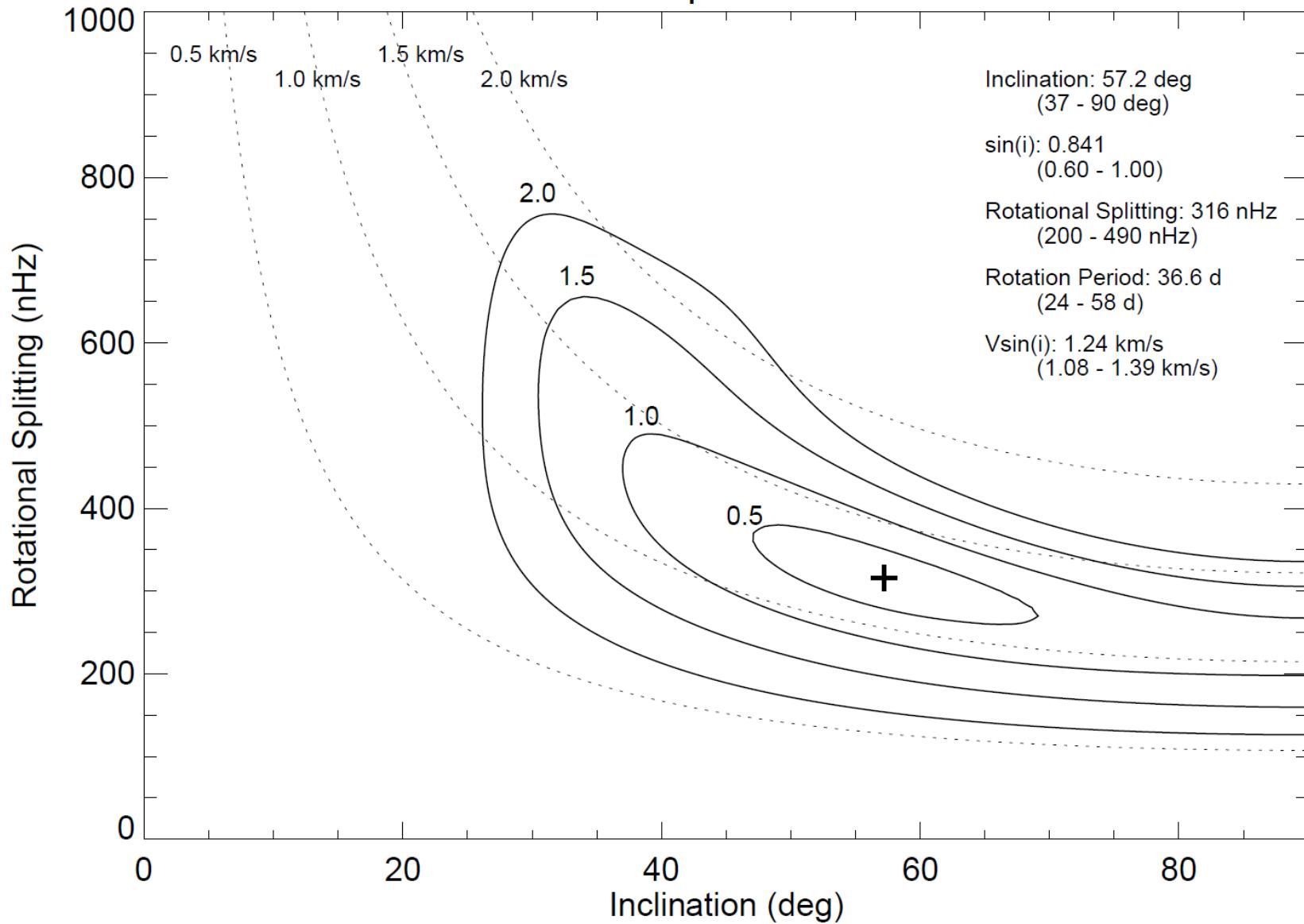


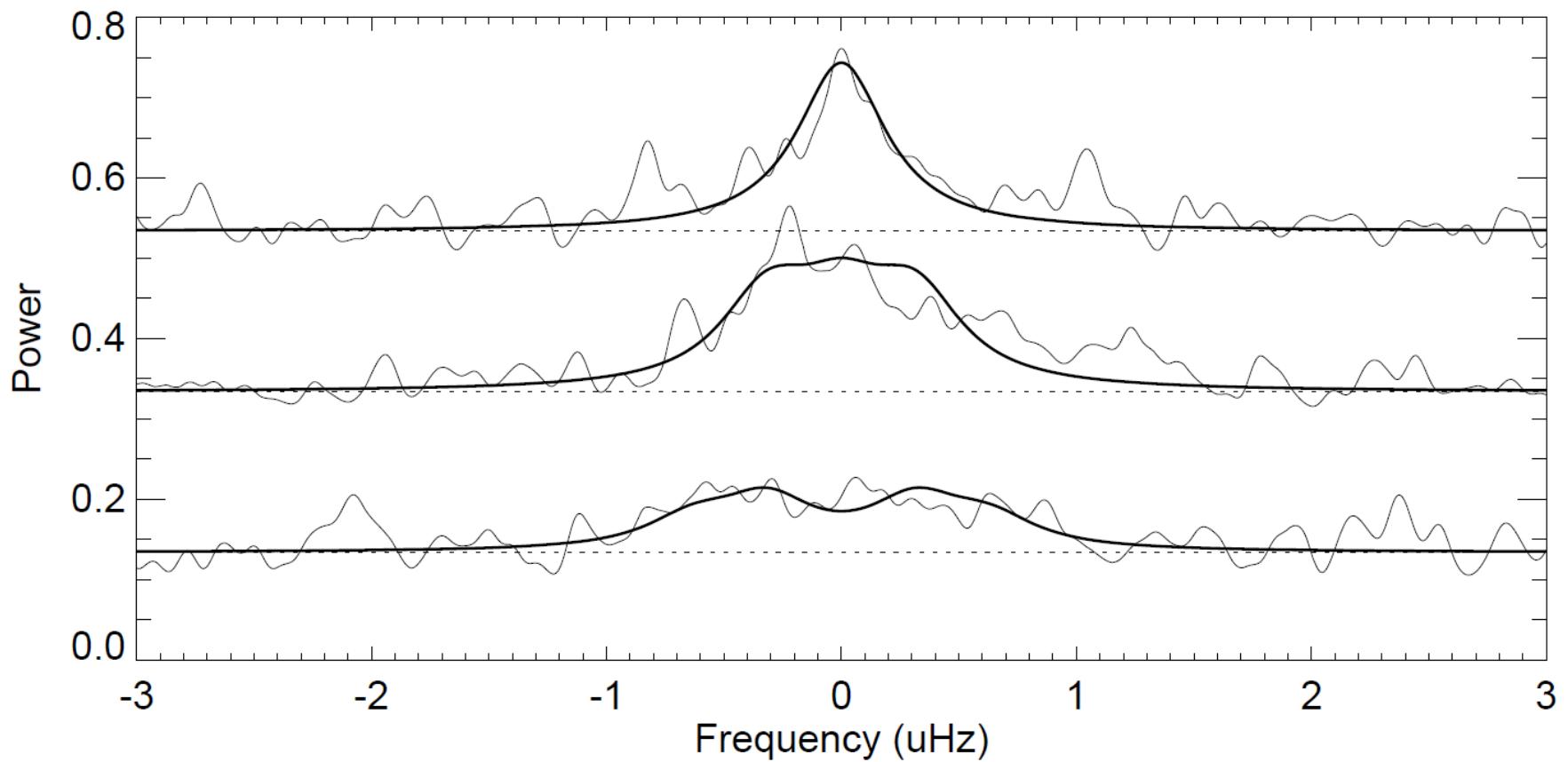
Rotation?



$i = 90 \text{ deg}$

# Kepler-10





# Kepler-10 frequencies

Power spectrum can be found via:

[http://astro.phys.au.dk/~hans/Kepler10/AMPLITUDE\\_SPECTRUM.txt](http://astro.phys.au.dk/~hans/Kepler10/AMPLITUDE_SPECTRUM.txt)

#	FRQ(μHz)	AMP(ppm)	PHASE	0.000000
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	FRQ	ERR	I	n
	1879.51	0.21	1	14
	1934.18	0.26	2	14
	1996.43	0.29	1	15
	2059.37	0.22	0	16
	2112.69	0.21	1	16
	2169.72	0.20	2	16
	2175.91	0.22	0	17
	2230.66	0.15	1	17
	2287.09	0.18	2	17
	2293.20	0.15	0	18
	2348.48	0.13	1	18
	2405.81	0.20	2	18
	2410.80	0.13	0	19
	2466.60	0.11	1	19
	2523.72	0.19	2	19
	2528.10	0.15	0	20
	2584.59	0.13	1	20
	2641.69	0.17	2	20
	2645.83	0.16	0	21
	2702.47	0.14	1	21
	2761.50	0.25	2	21
	2763.67	0.26	0	22
	2821.87	0.23	1	22