



Photospheric counter Evershed flows in the penumbra of sunspots

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M. Rempel³

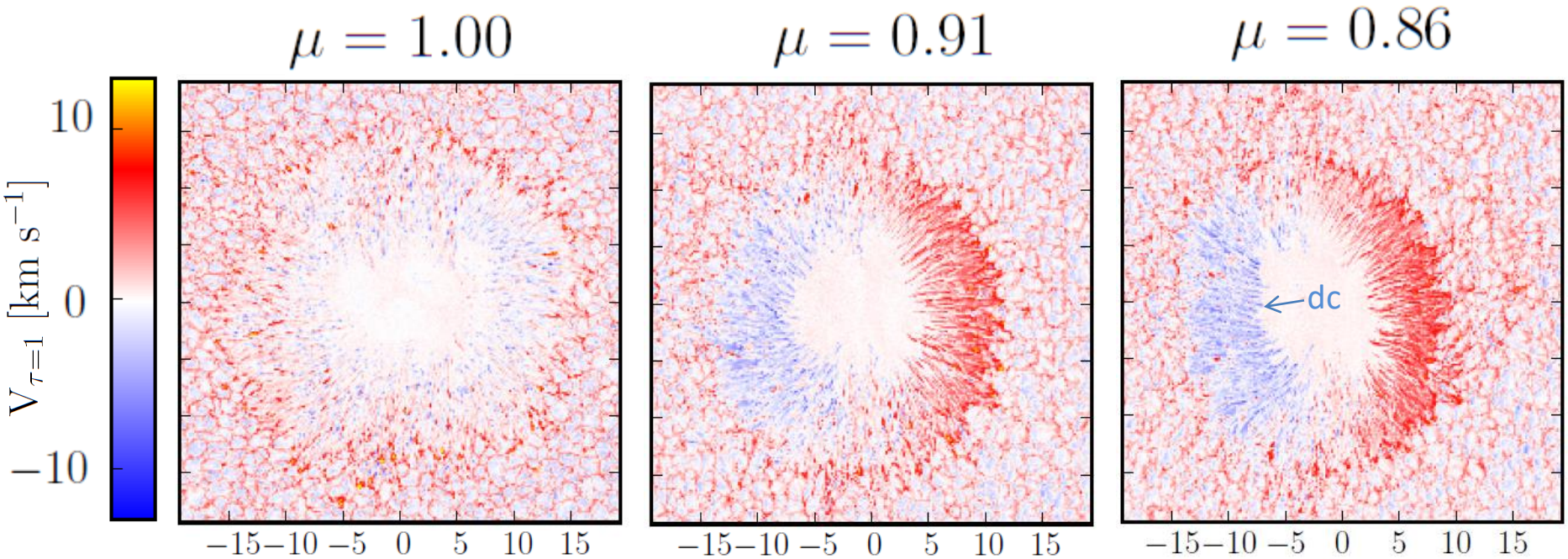
SOLARNET IV MEETING
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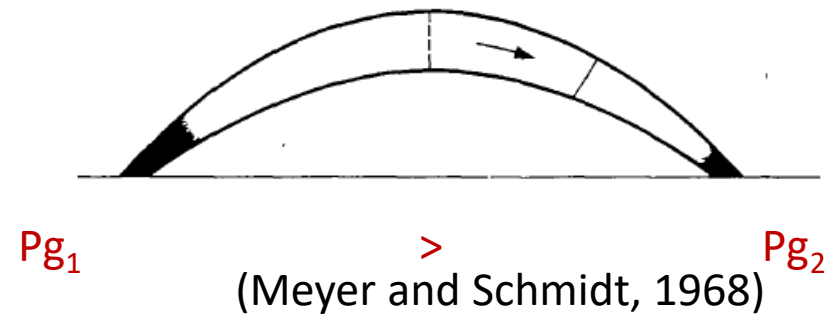
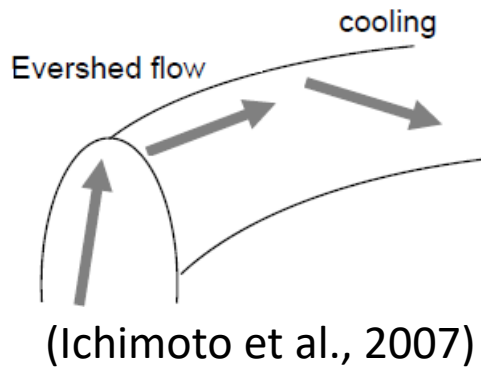
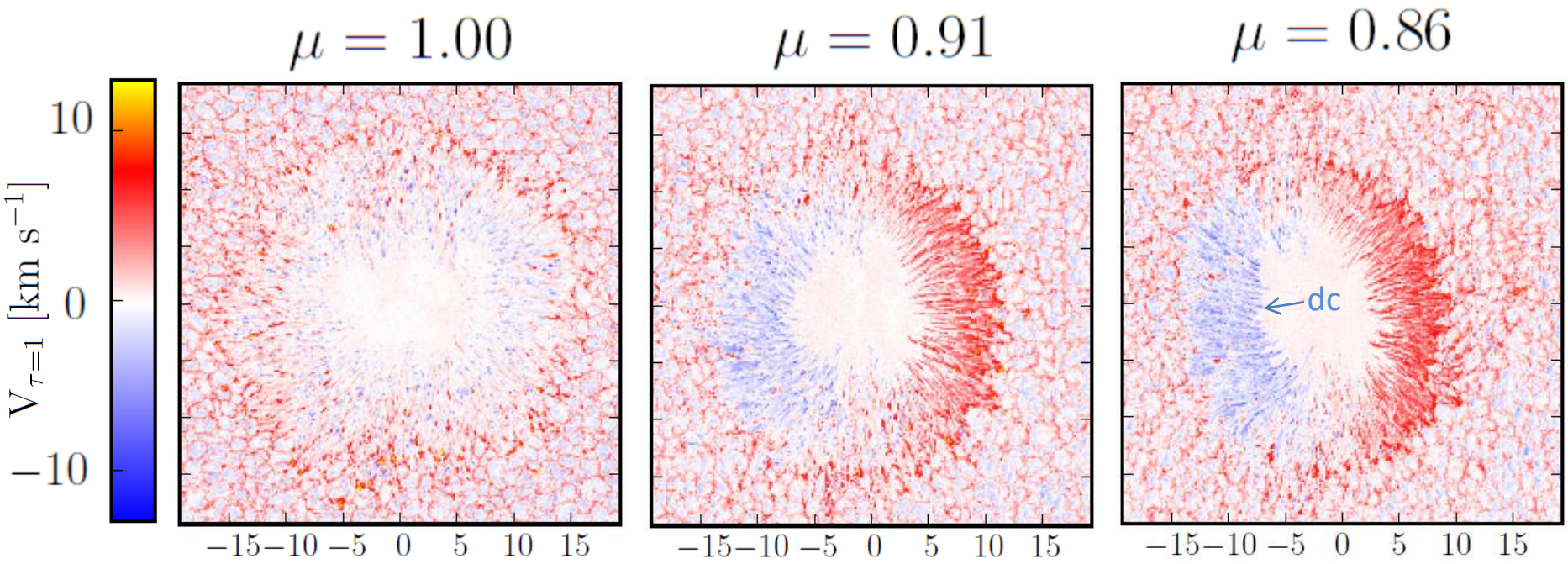
³High Altitude Observatory, NCAR, Boulder, USA

The Evershed effect

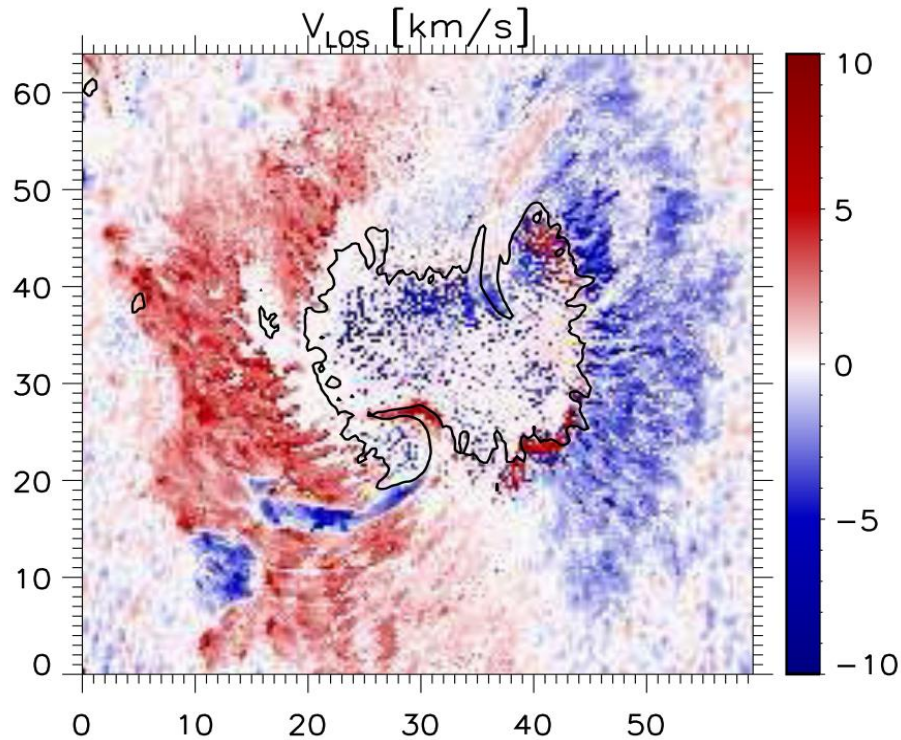


AR10933, Hinode SOT/SP
SPINOR inversions by M. van Noort

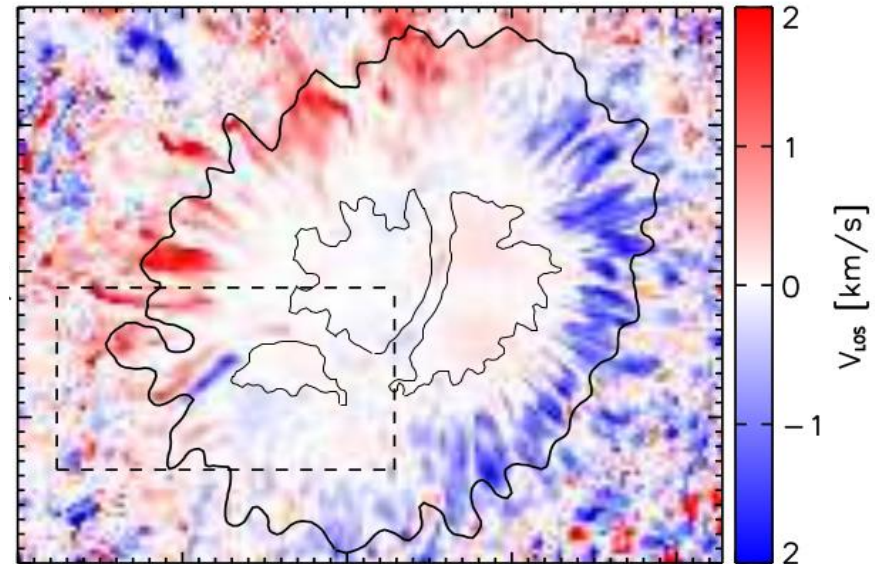
The Evershed effect



Photospheric counter-Evershed flows (singular filaments)

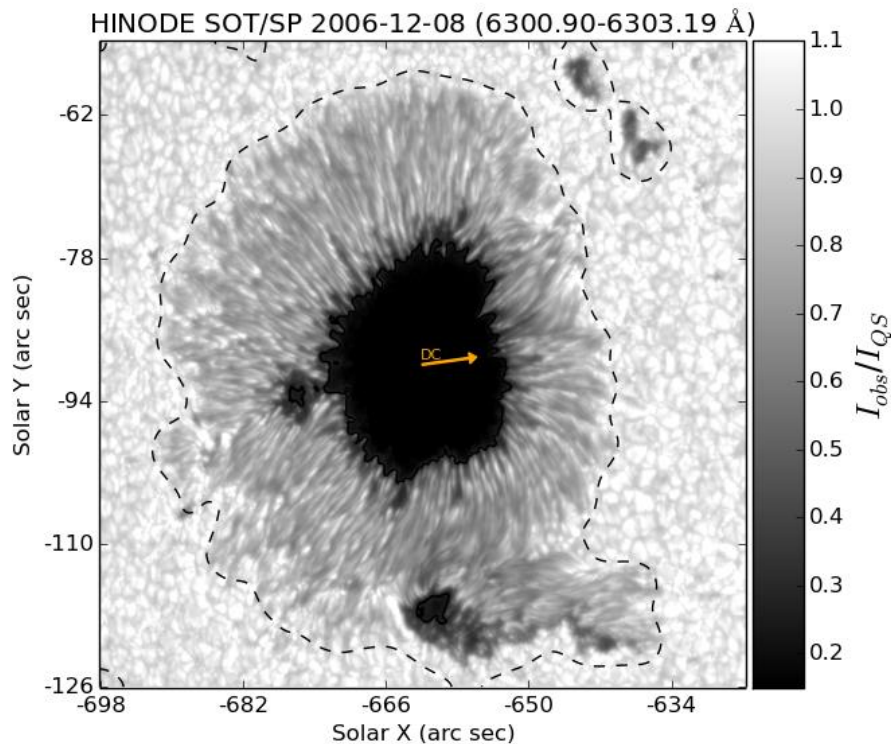


NOAA AR 11302 Hinode/SP
(Kleint and Sainz Dalda, 2013)

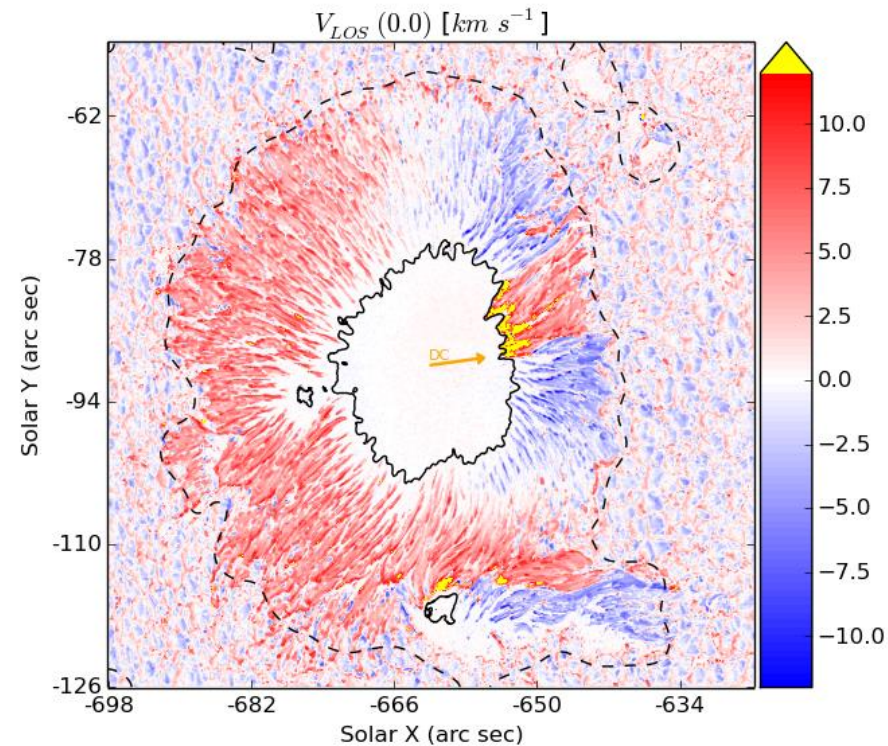


NOAA AR 11271 Hinode/SP
(Louis et al., 2014)

Photospheric counter-Evershed flows (large penumbral region)



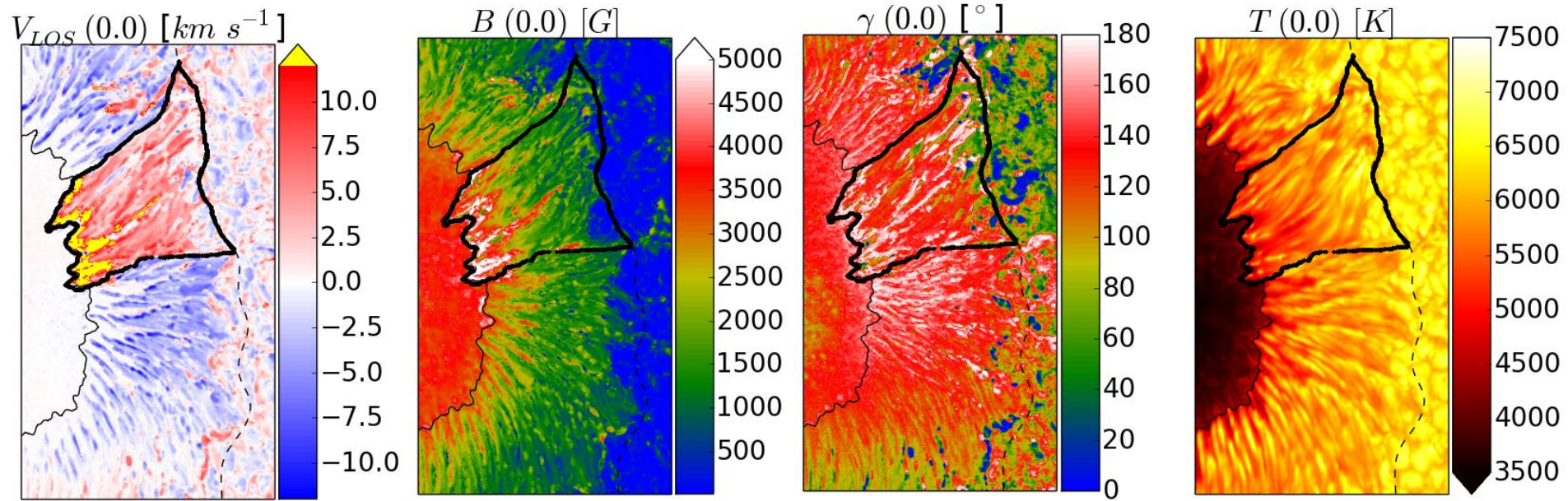
NOAA AR 10930 Hinode/SP
1 pixel=0.16"



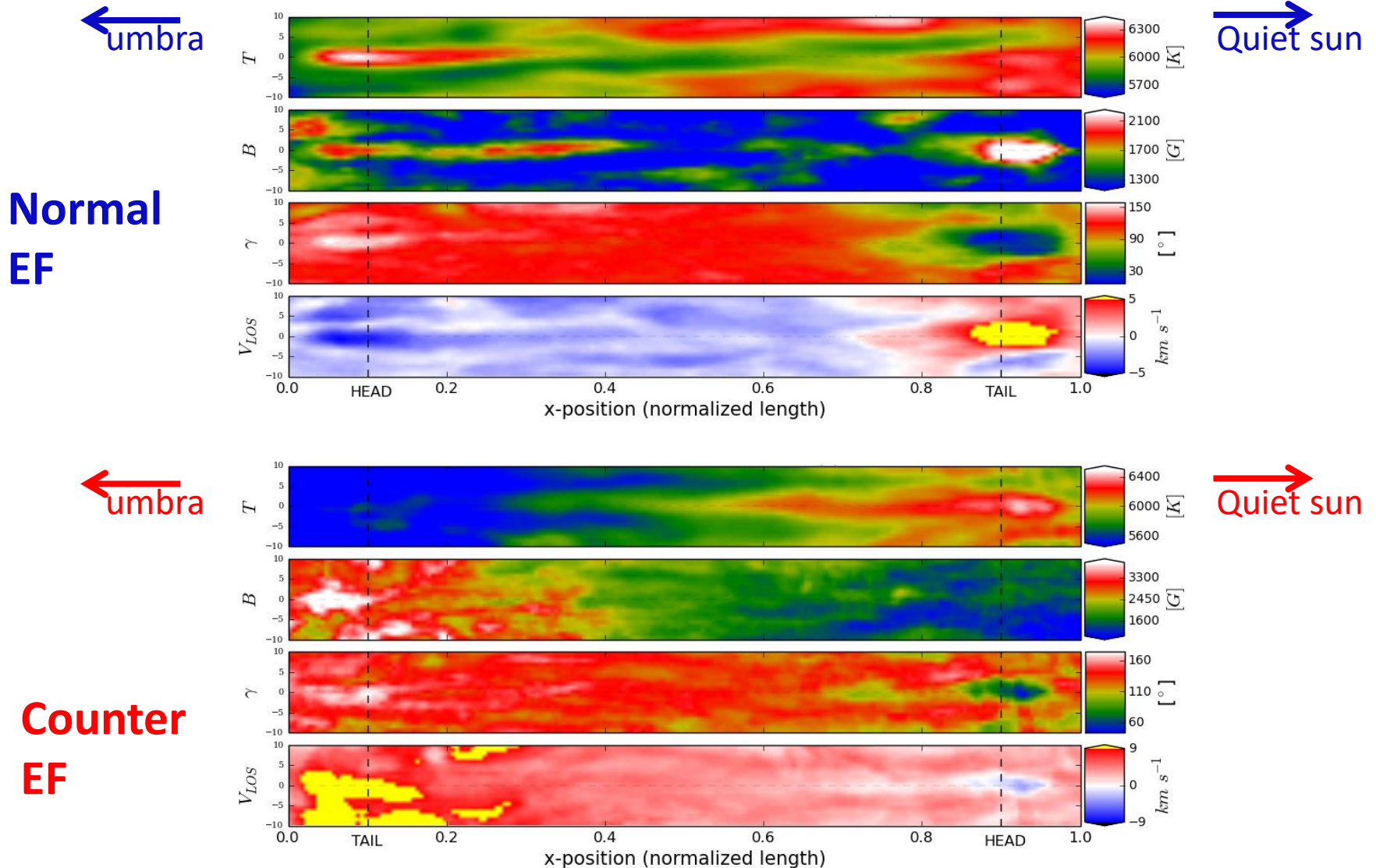
SPINOR 2D inversions
1 pixel=0.08 "

heliocentric angle: $\theta \approx 47^\circ$

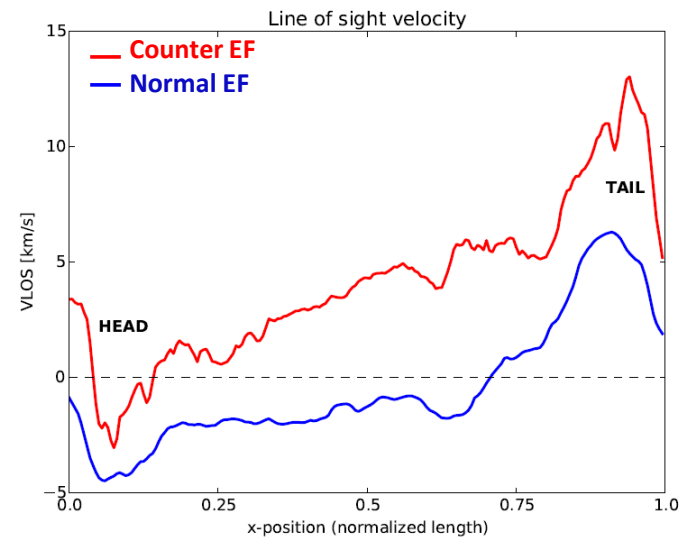
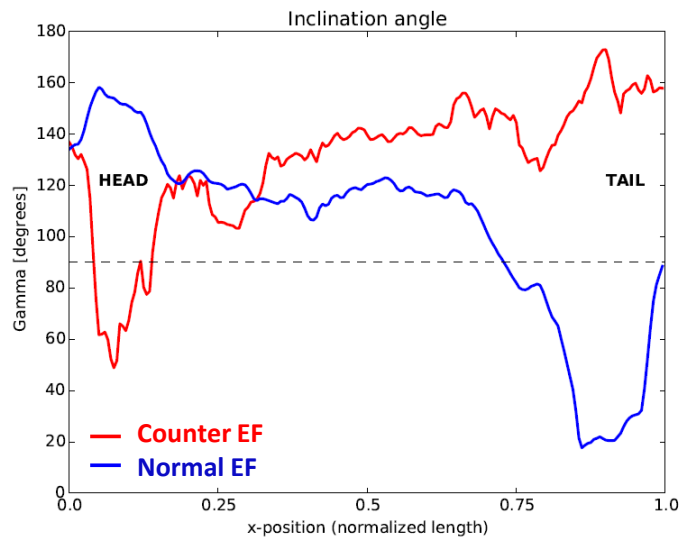
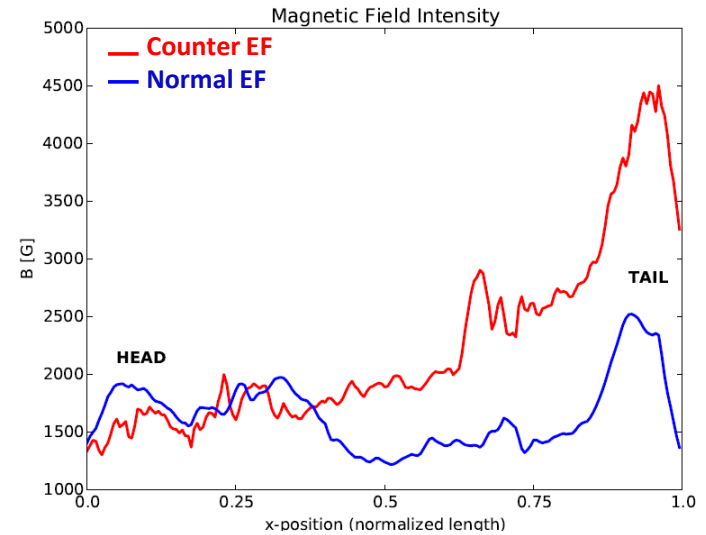
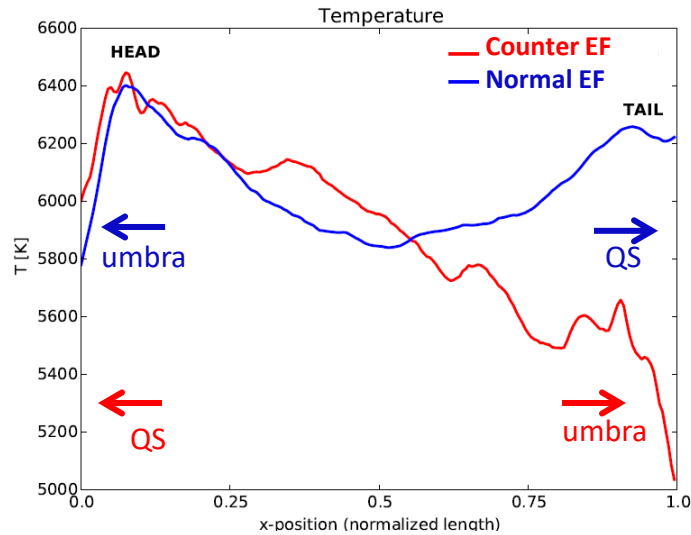
SPINOR 2D: $\log(\tau)=-2.0, -0.8$ and 0



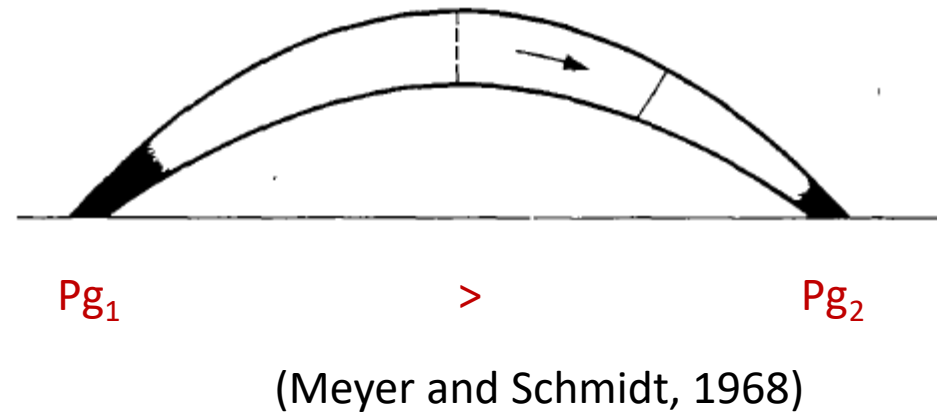
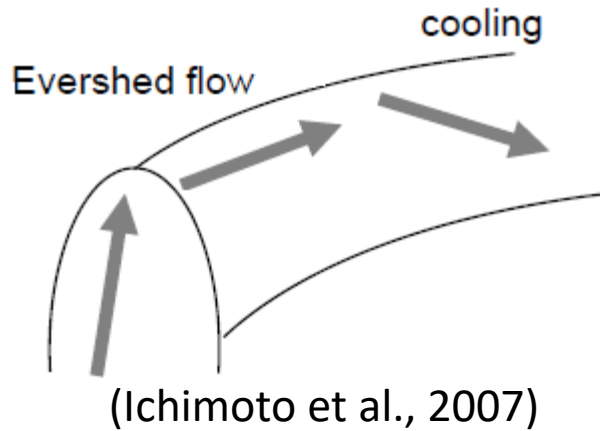
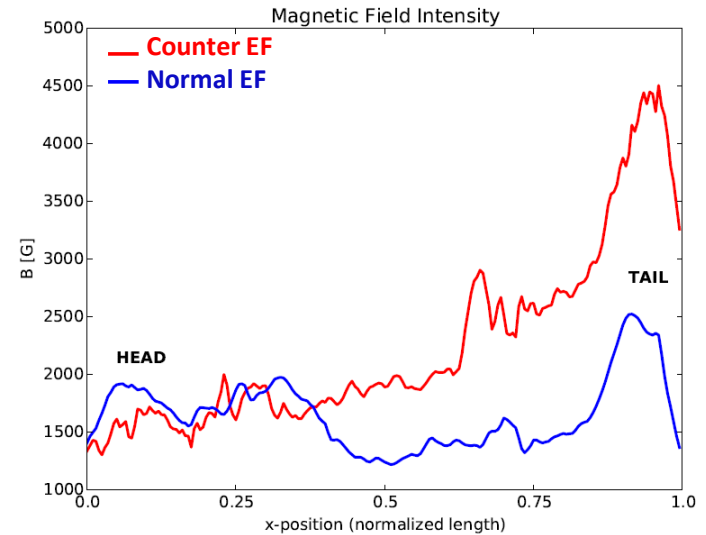
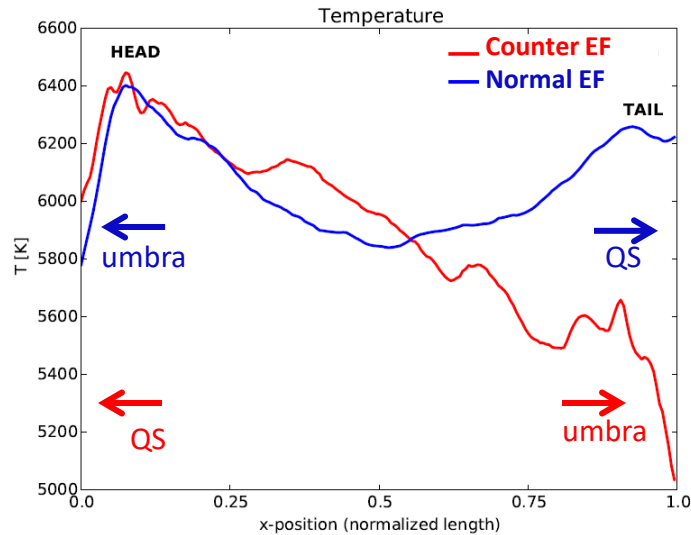
Characteristic filaments



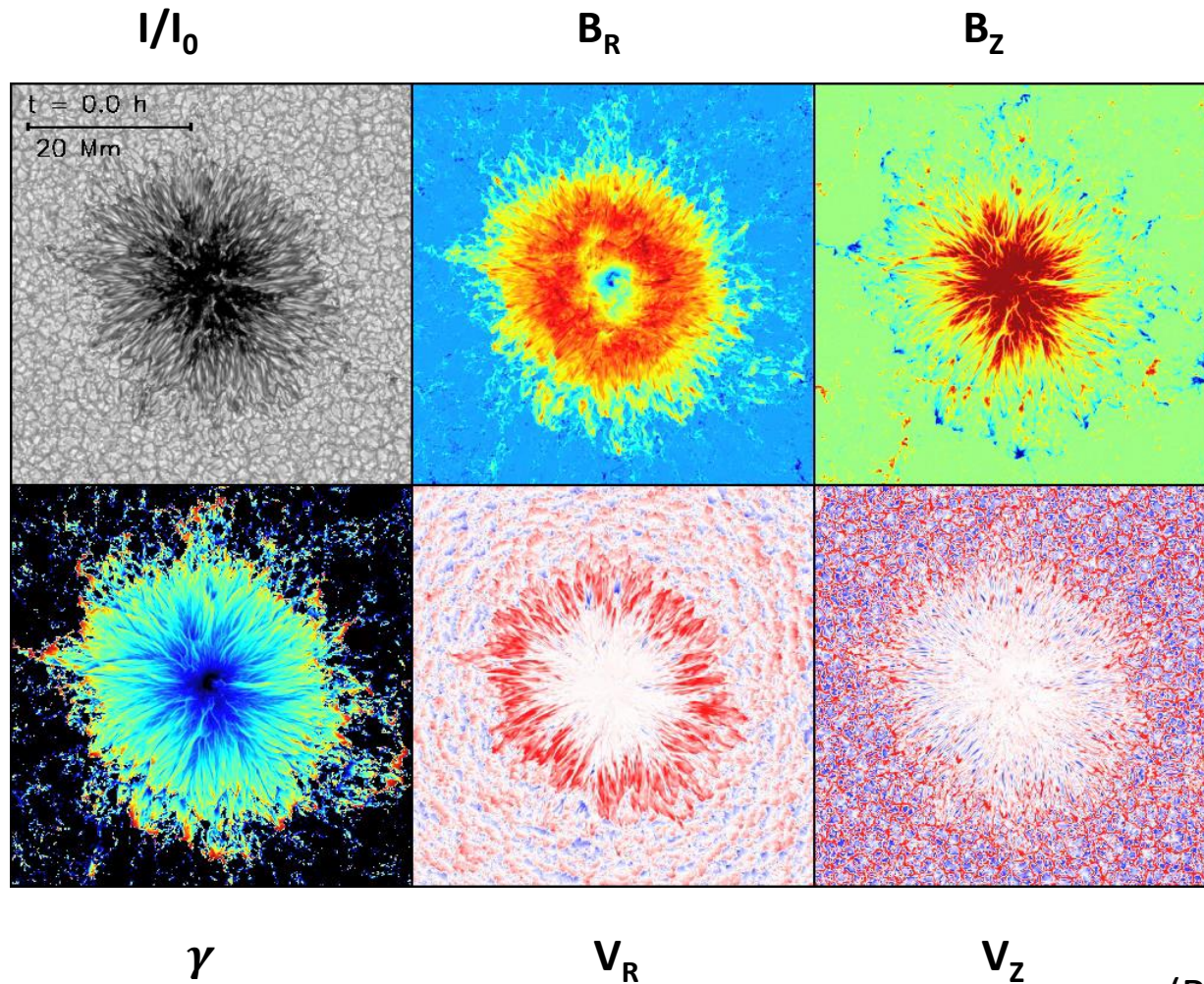
Central axes of filaments



Central axes of filaments



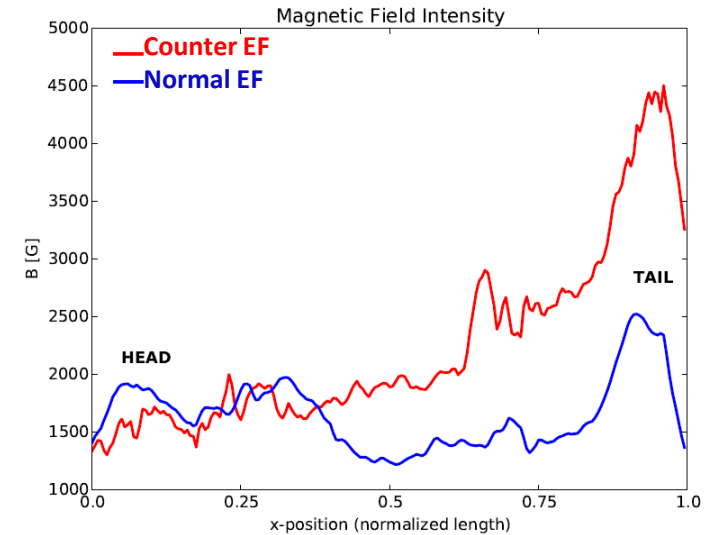
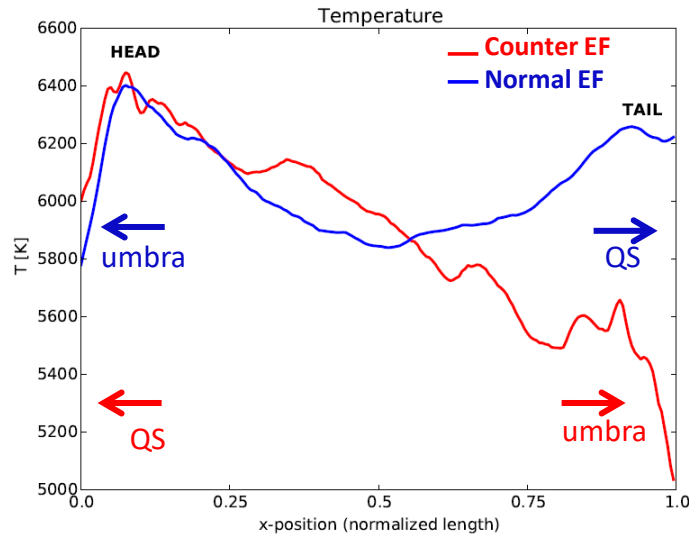
MURaM simulation: Formation of a penumbra



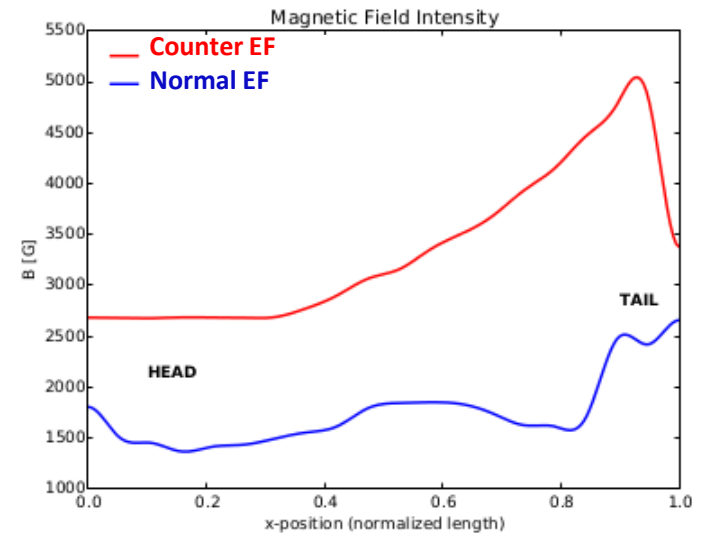
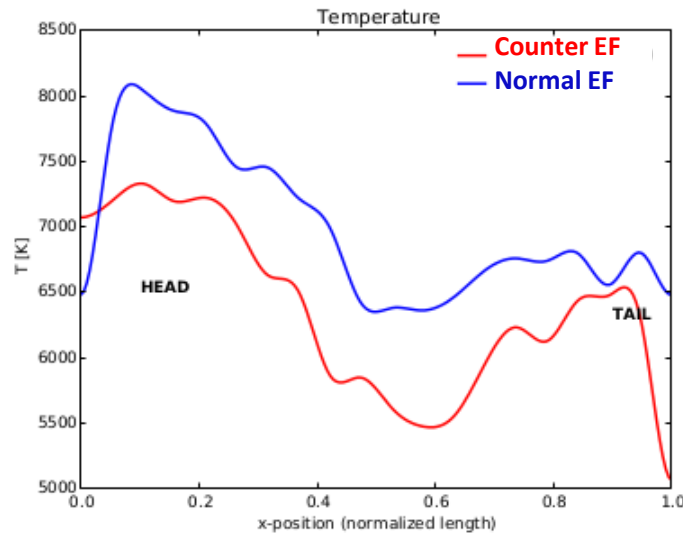
(Rempel 2015)

Comparison: Observations vs simulations

Observations
 $\tau = 1$

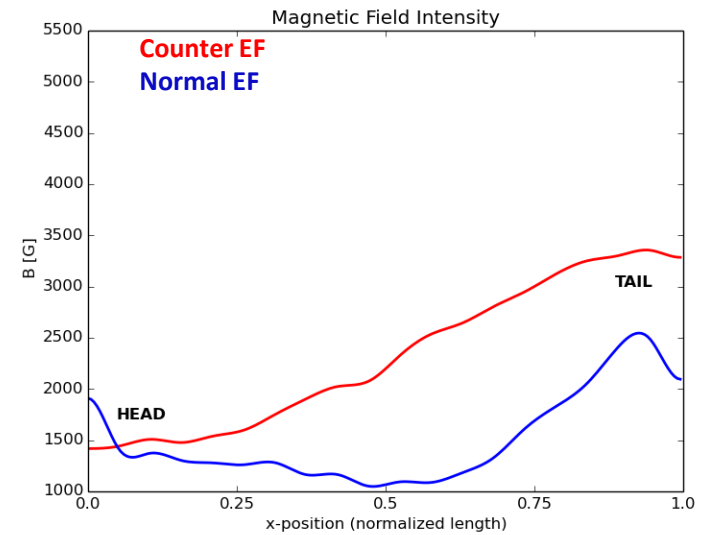
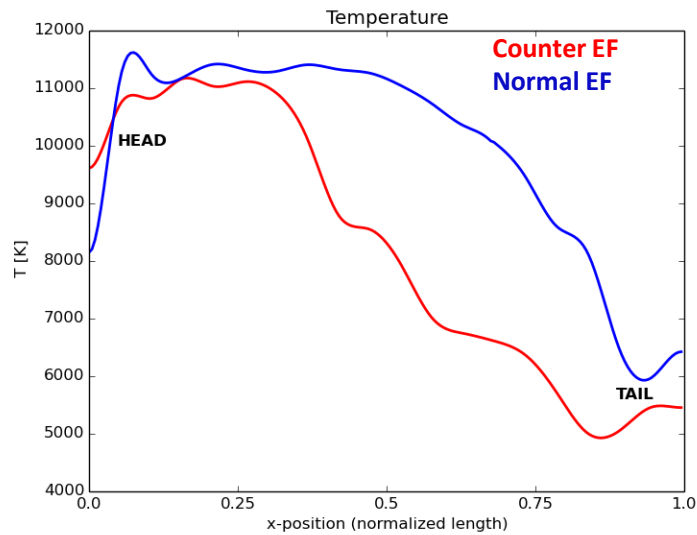


Simulations
 $\tau = 1$

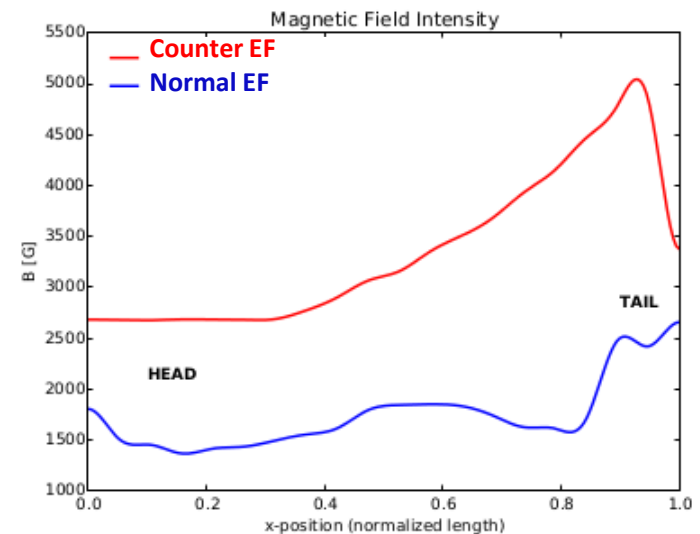
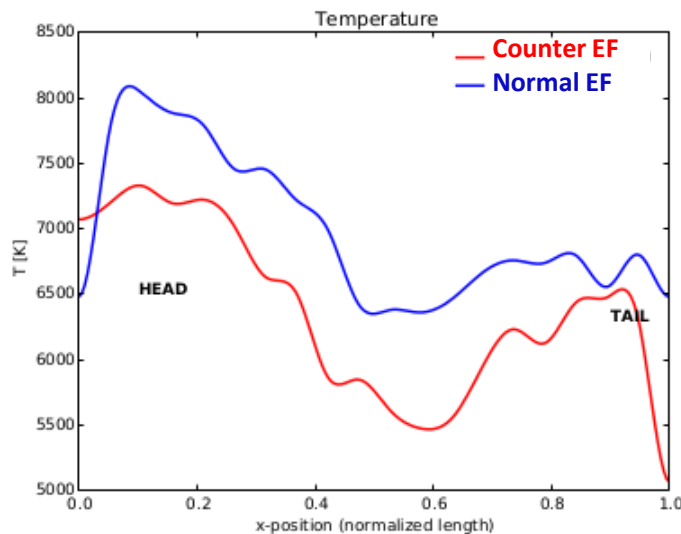


Comparison: optical depth vs geometrical height

Simulations
 $h = -226 \text{ km}$

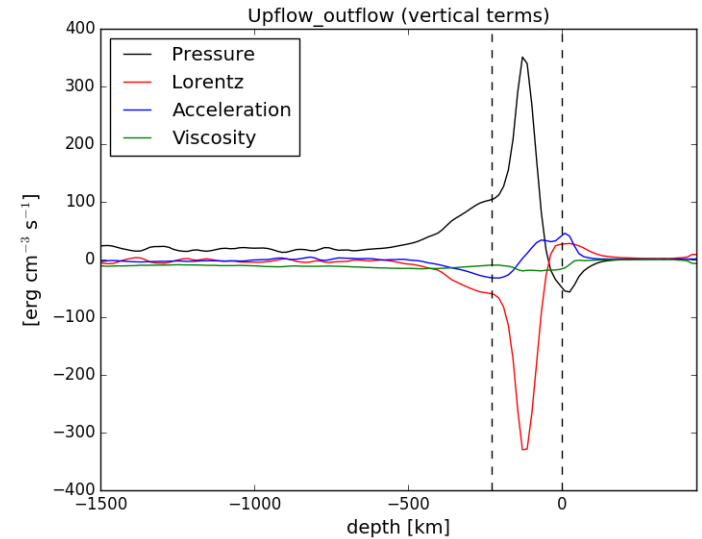
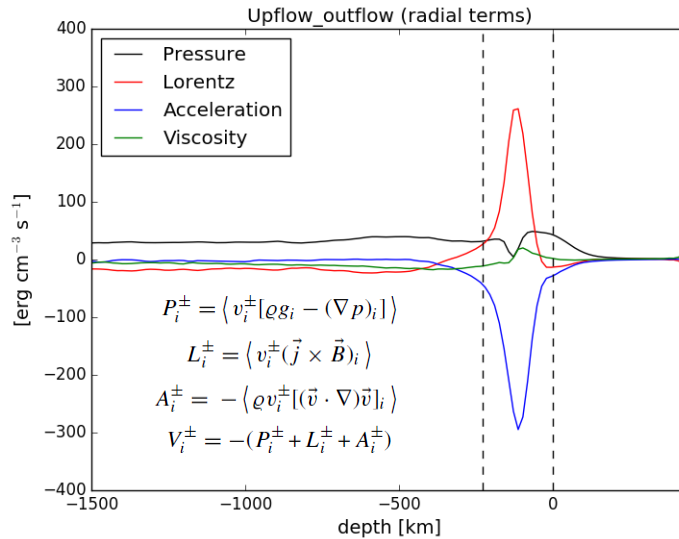


Simulations
 $\tau = 1$

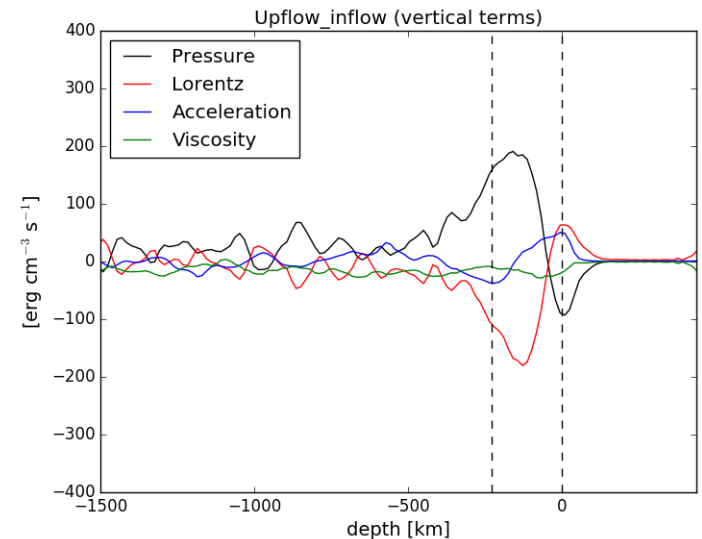
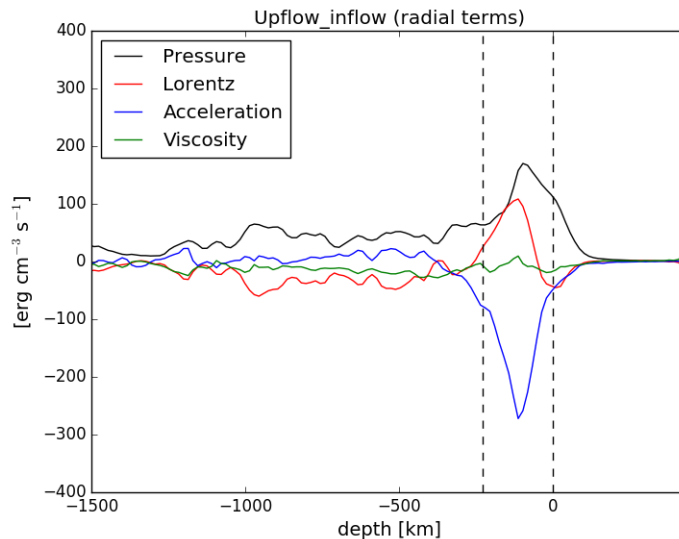


Driving mechanisms: 10 h average

Normal EF



Counter EF



Summary and outlook

- Unusual observation of photospheric CEFs in a large penumbral region
 - Photospheric flows are confined in penumbral filaments
 - NEF and CEF show anti-correlated flow structures
 - Both, ∇T and ∇B are consistent with the flow direction in each case, NEF and CEF.
- MHD simulations reproduce CEFs
 - Qualitative similarities as in observations near $\tau = 1$ (but no compact footpoints)
 - NEF and CEF are driven in a thin boundary layer near $\tau = 1$
 - Driving mechanisms:
 - NEF: convection + magnetic deflection (quasi-stationary)
 - CEF: siphon flow + magnetic deflection (transient)
- Future work
 - To study the temporal evolution of CEFs: observations and simulations.