

The external component: surrounding matter as an agent in star formation

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How does an external matter distribution effect a molecular cloud?

Recent analytical studies have demonstrated that the media surrounding a molecular cloud can importantly alter the gravitational potential to which it is subject, and contribute to its collapse or disruption (Ballesteros-Paredes, 2006; Gómez et al. 07). We study three cases where this effect can be important.

Internal gravitational energy $E_{grav} = \frac{1}{2} \int_V \rho_{cloud} \phi_{cloud} dV$

External gravitational energy $W_{ext,i} = - \int_V x_i \frac{\partial \phi_{ext}}{\partial x_i} \rho_{cloud} dV$

1. Dark matter halos as a whole
2. Dark and barionic matter perturbations
3. Tidal influence from the Galaxy

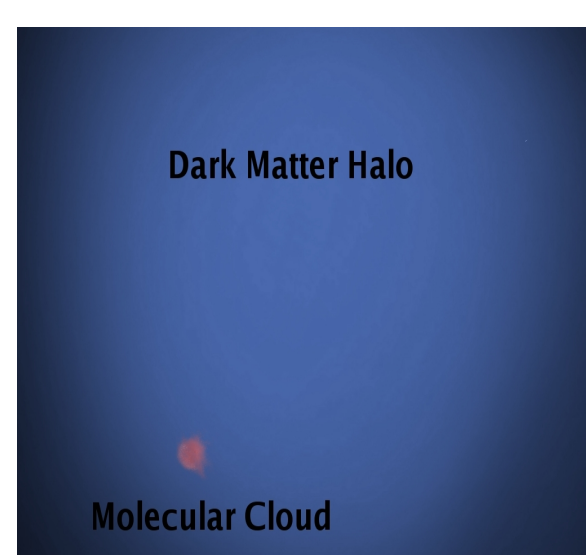
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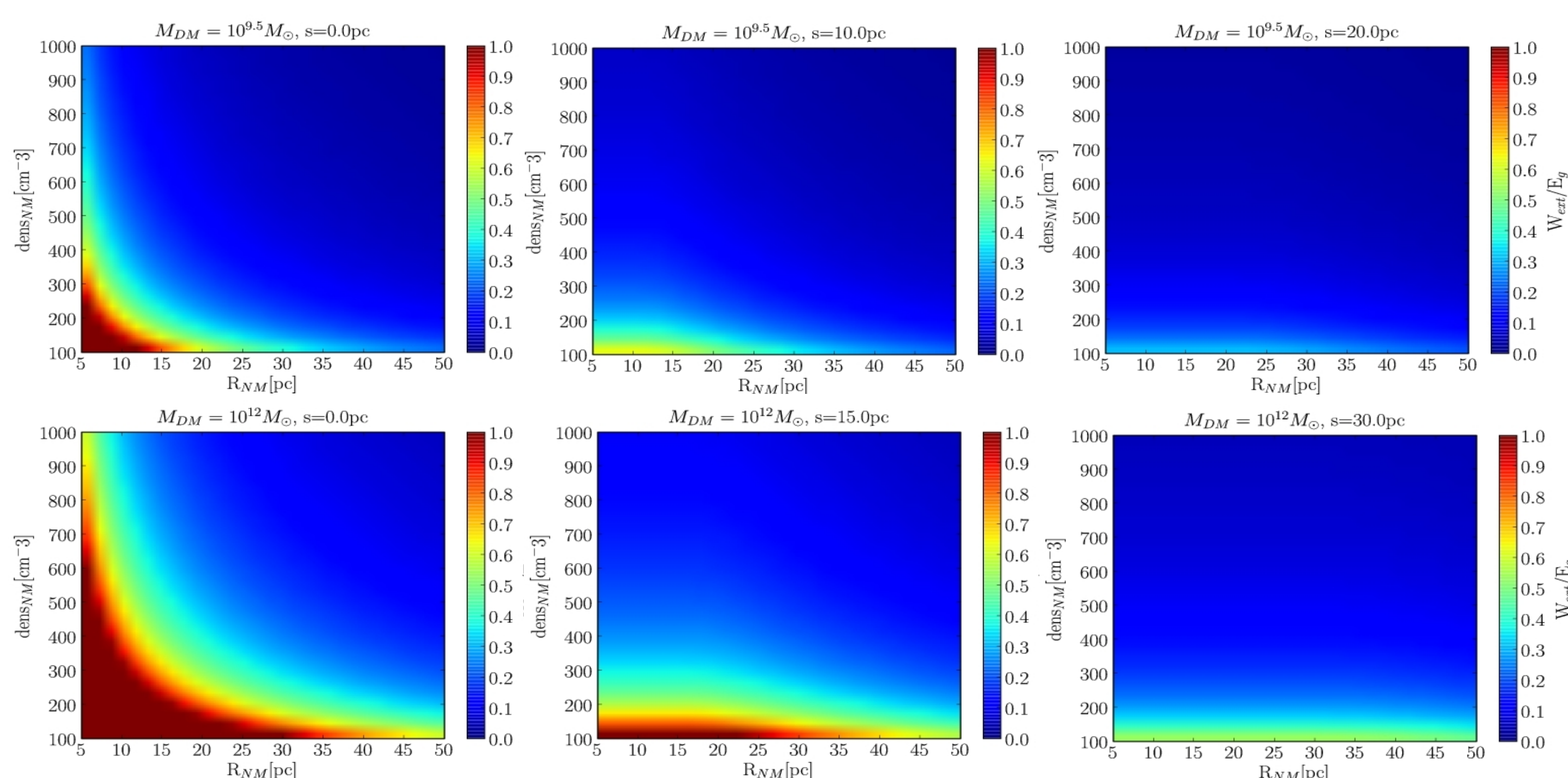
The dark matter contribution

PERTURBATION

Constant density molecular cloud embedded in a NFW dark matter distribution.



The halo extended potential is felt by the cloud and contributes as an external agent. Two very different length scales are involved: the cloud and the halo.



The halo profile is important only for clouds in galaxy centers and contributes to deplete their gas.

The presence of a molecular cloud perturbs the otherwise homogeneous dark matter background, which in turn alters the potential felt by the cloud and contributes to its collapse.

Cloud with Plummer density profile

$$\rho_P = \frac{3M}{4\pi a^3} \left(1 + \frac{r^2}{a^2}\right)^{-5/2}$$

Homogeneous and isotropic dark matter background:

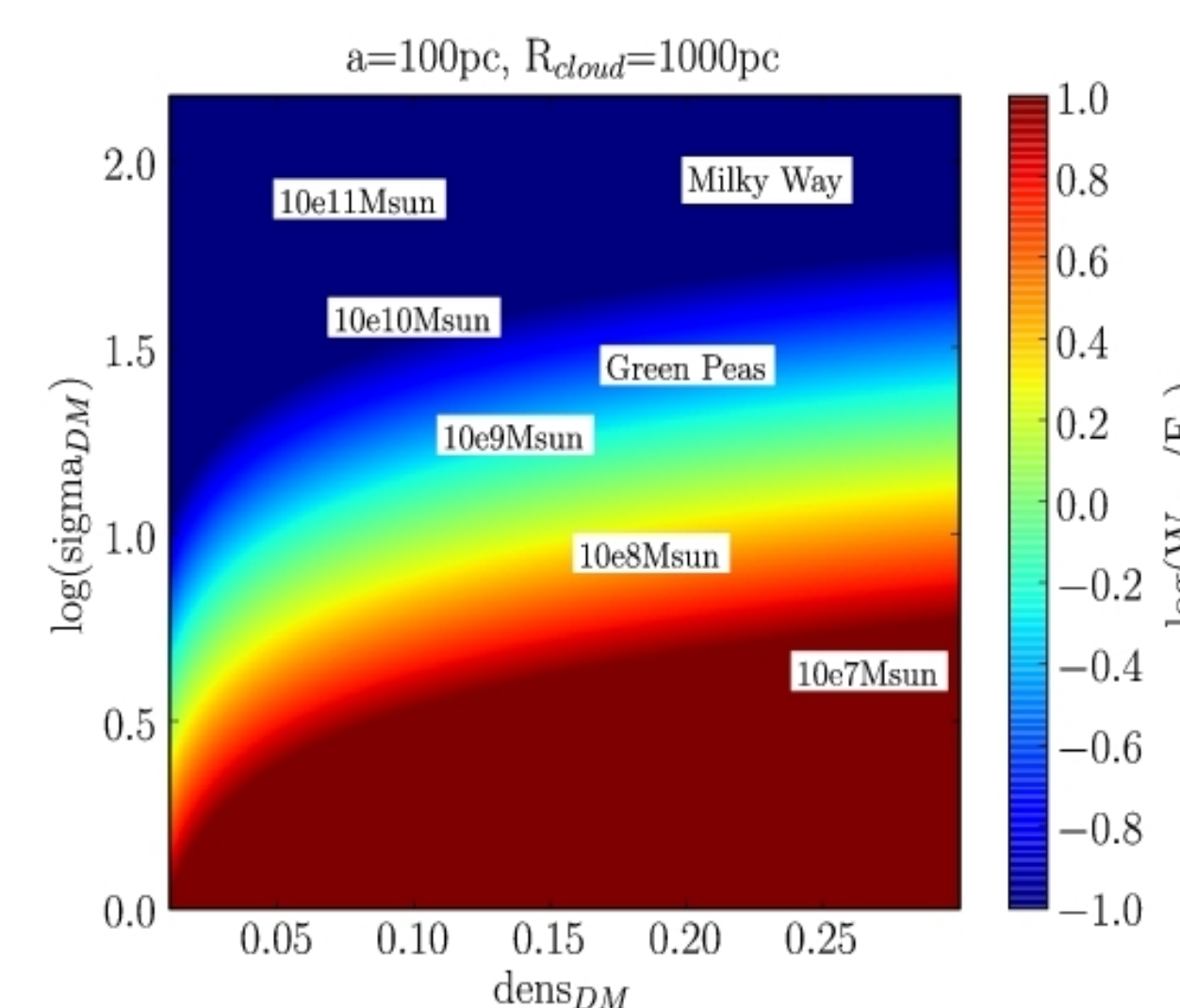
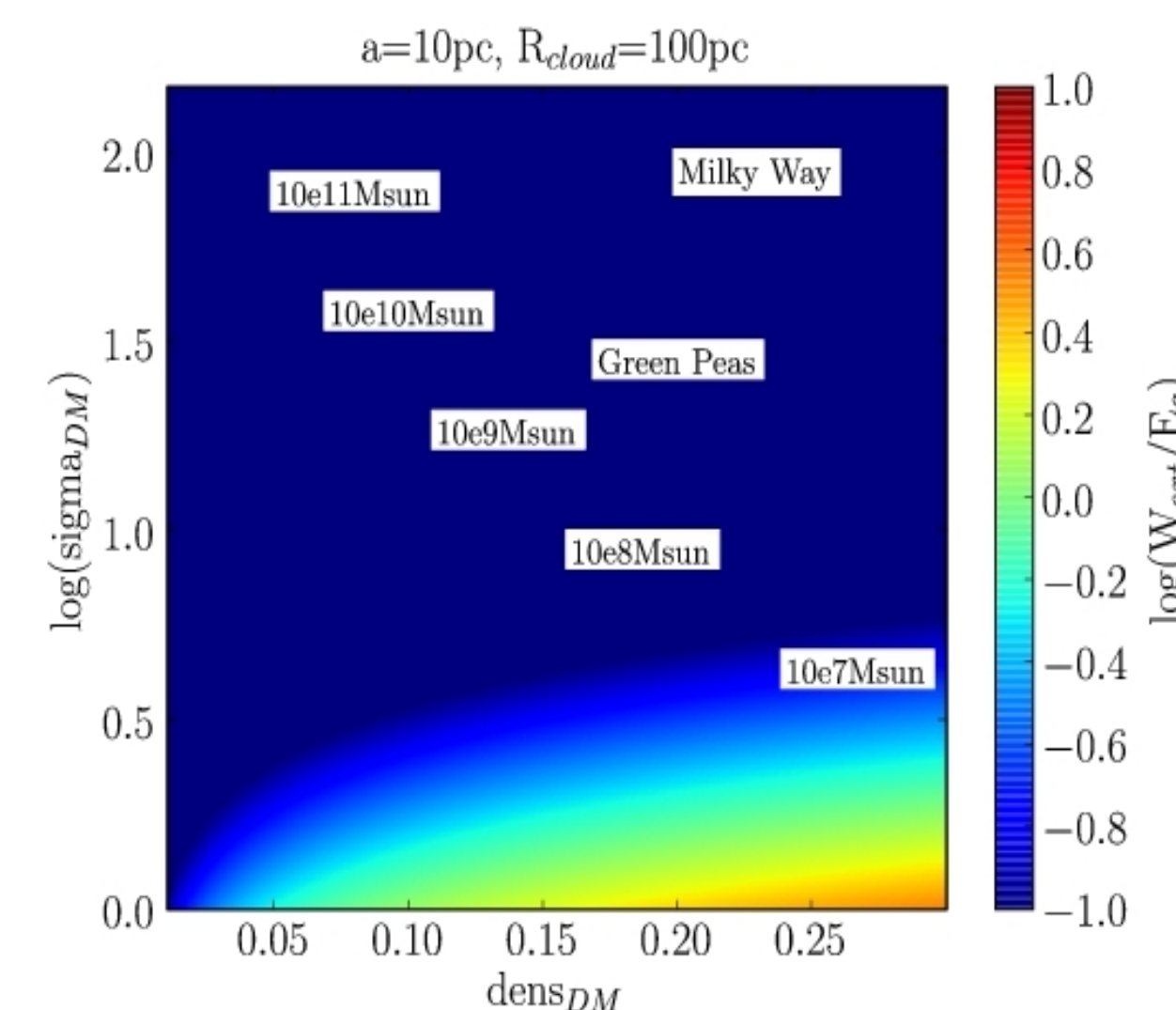
constant density ρ_0
 constant velocity dispersion σ

External to internal energy ratio

$$\frac{W_{ext}}{E_g} = -1.8025 \left(\frac{a}{10\text{pc}}\right)^2 \left(\frac{\rho_0}{M_\odot \text{pc}^{-3}}\right) \left(\frac{\sigma}{\text{kmsec}^{-1}}\right)^{-2} \frac{f_{ext}(R/a)}{f_g(R/a)}$$

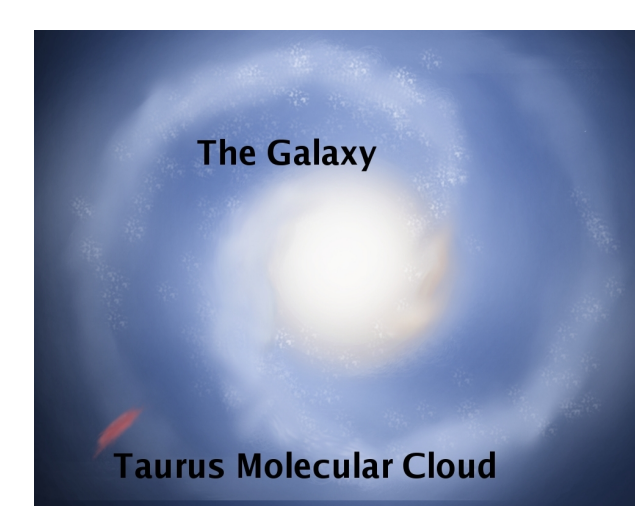
SIMULATIONS IN PREPARATION

The interrelation with dark matter is relevant for big clouds in small galaxies. Could explain starbursts in Green Peas.

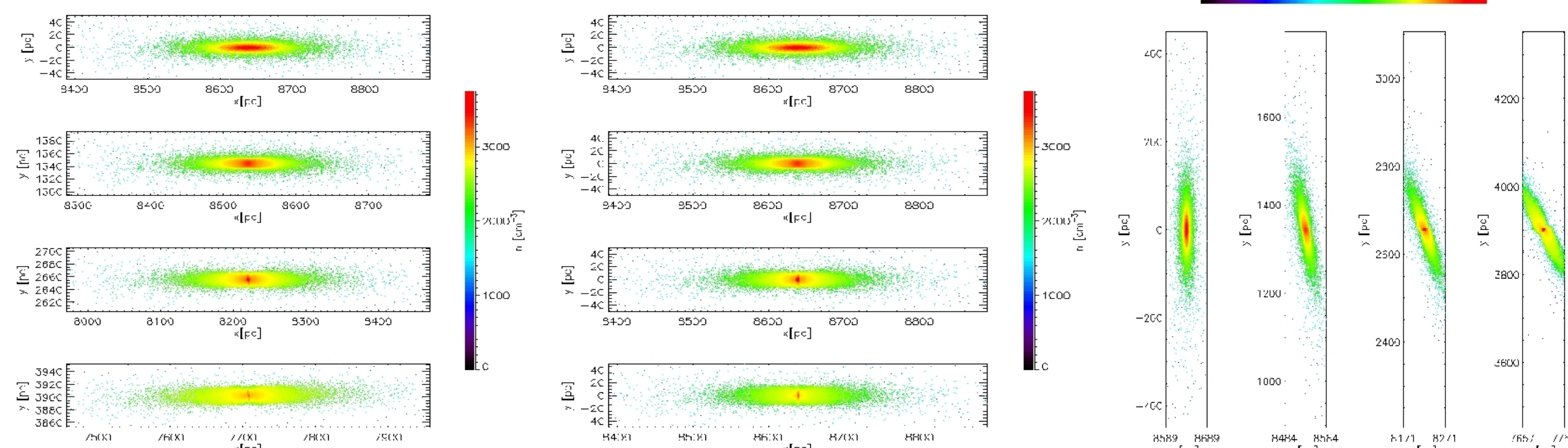


Tidal forces from the Galaxy: Taurus molecular cloud

An oblate ellipsoidal Plummer spheroid resembling Taurus molecular cloud is embedded in an Allen-Santillán (1991) Galactic potential for the Galaxy, which will cause an external potential energy. Two different alignments of the cloud with respect to the Galaxy are considered and a cloud without any external potential for comparison.



t = 0 Myr



t = 6 Myr

t = 12 Myr

t = 17 Myr

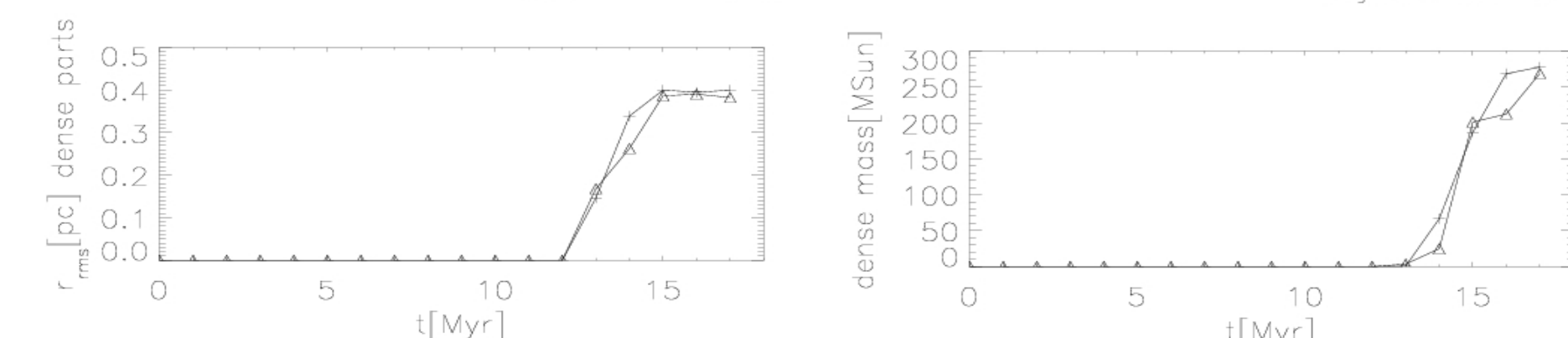
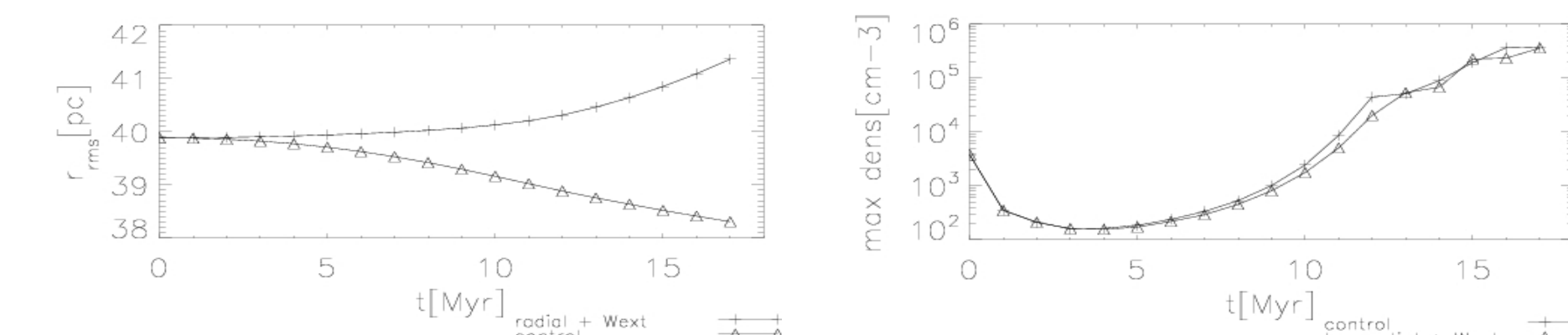
Radial + West.

Control (no external potential)

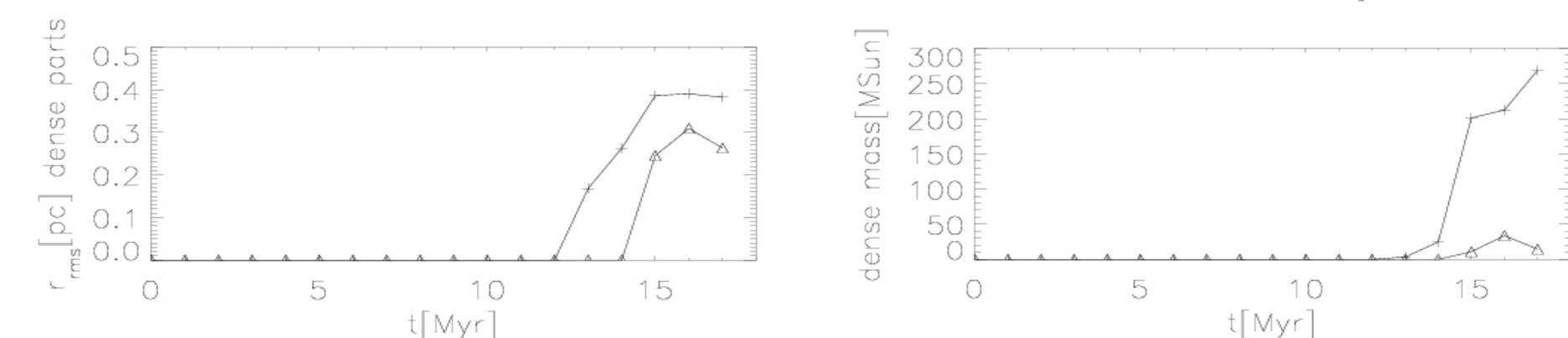
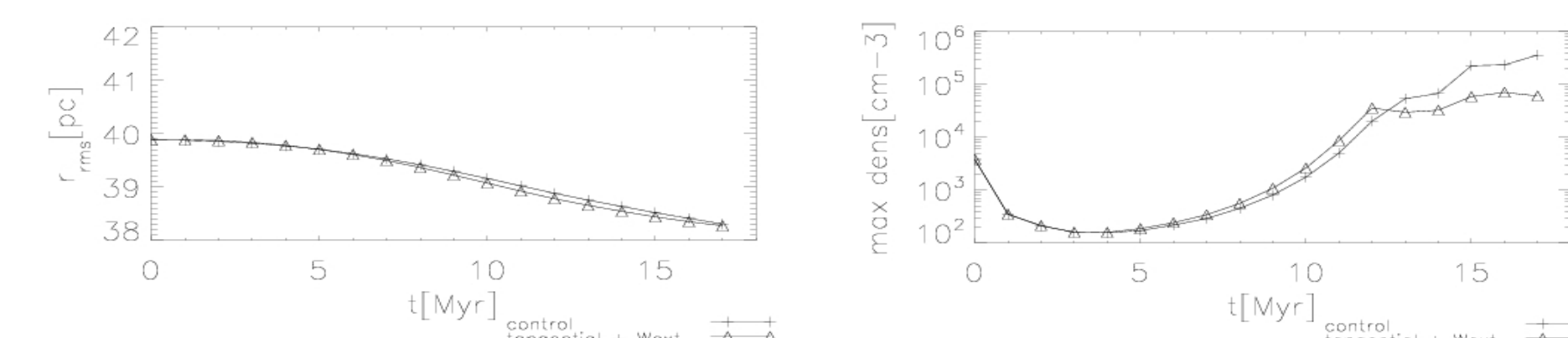
Tangential + West.

The tidal forces from the Galactic mass distribution clearly cause global disruption of the cloud, specially for radial alignment.

A radially aligned cloud is globally disrupted



A tangential cloud still collapses



HI-RES SIMULATIONS STILL RUNNING!