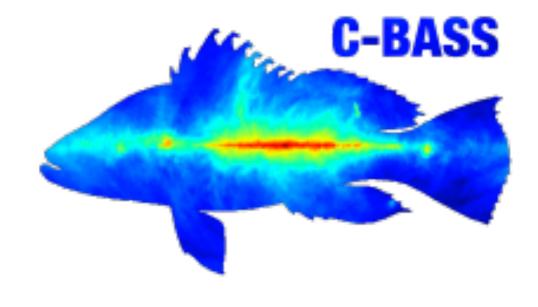
Template Fitting Analysis with C-BASS

Stuart Harper on behalf of the C-BASS collaboration

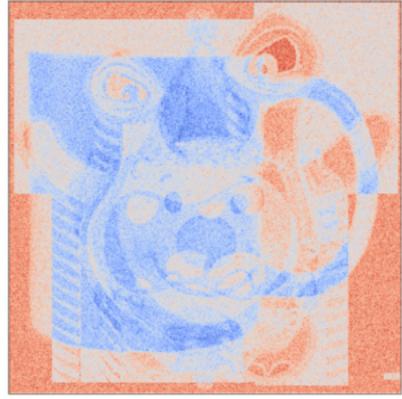
Radio Foregrounds Conference, IAC, 2018



The University of Manchester



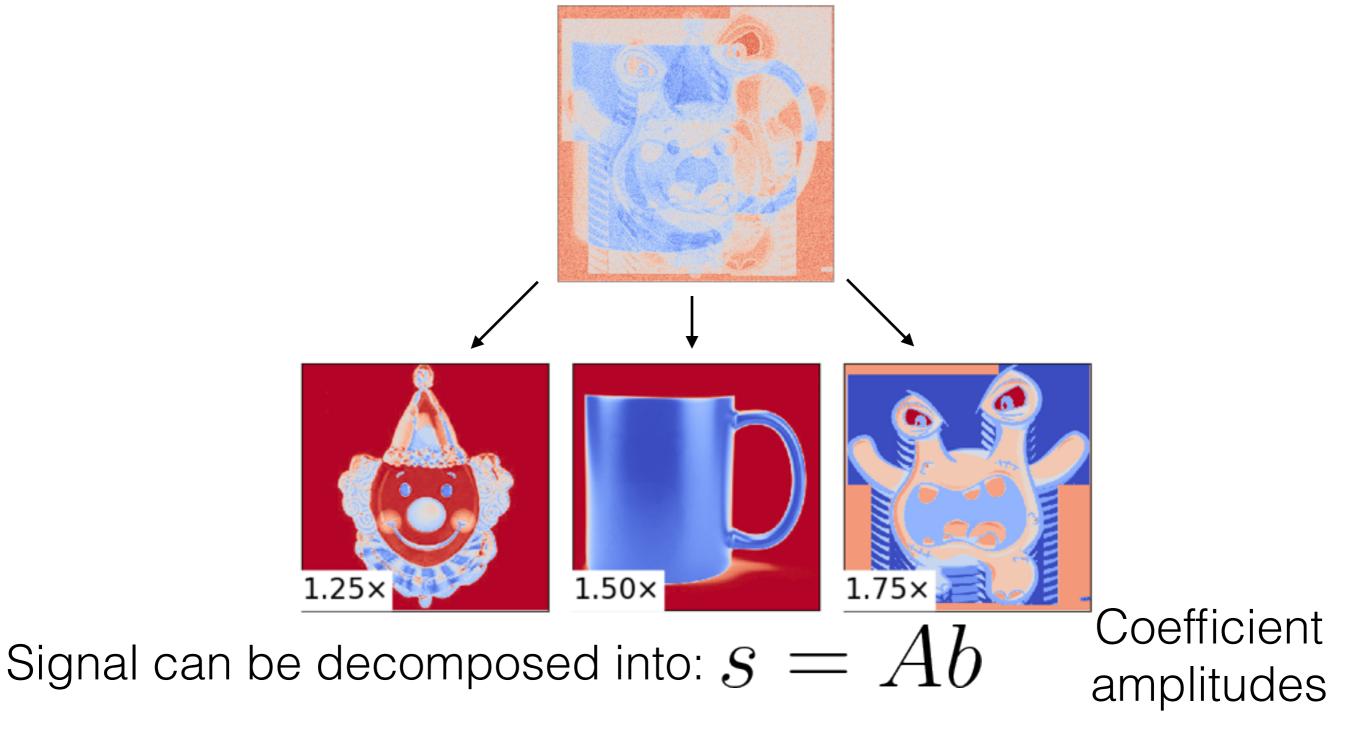
We have an observation:



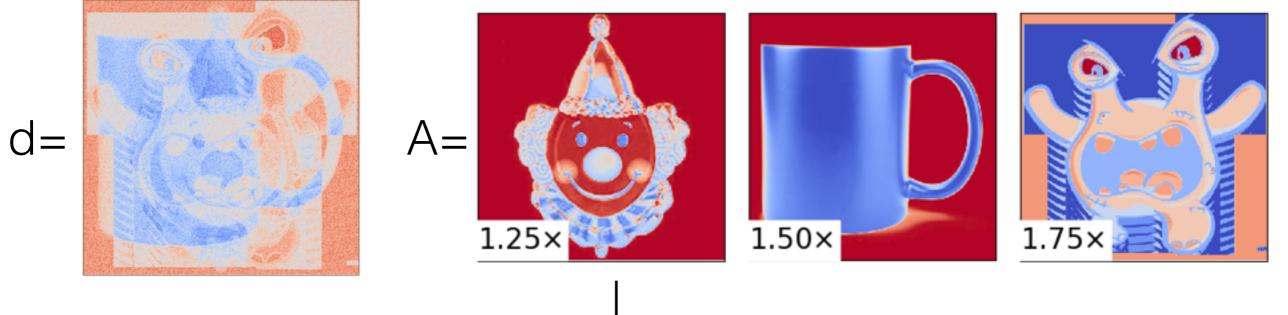
That can be decomposed into two components:

$$d = s + n$$
Noise
Signal

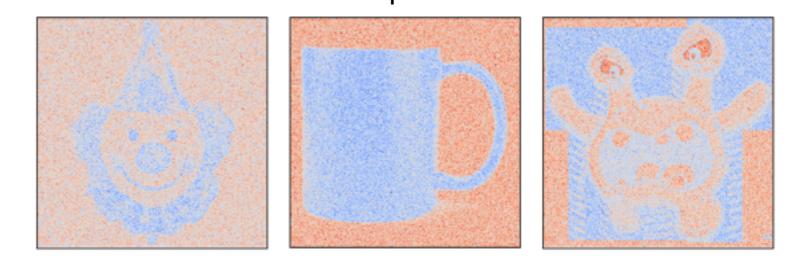
Signal composed of templates/matched filtered



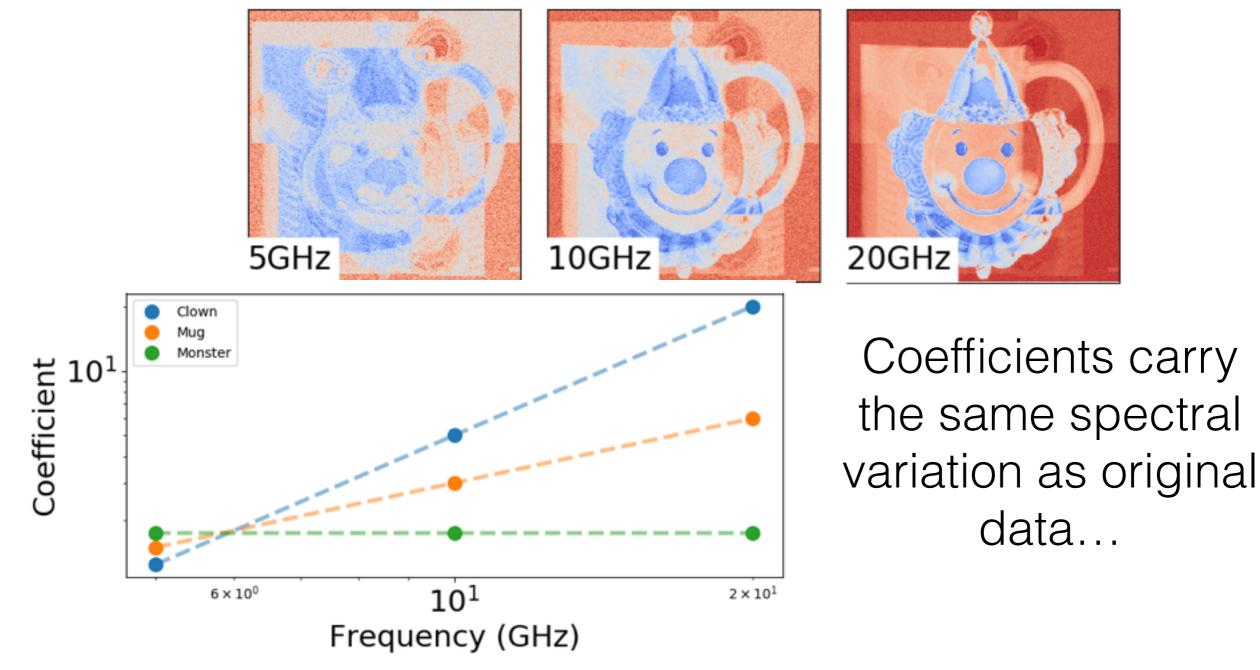
Matched filters/templates



$$b = (AN^{-1}A^{\mathsf{T}})^{-1}(AN^{-1}d^{\mathsf{T}})$$



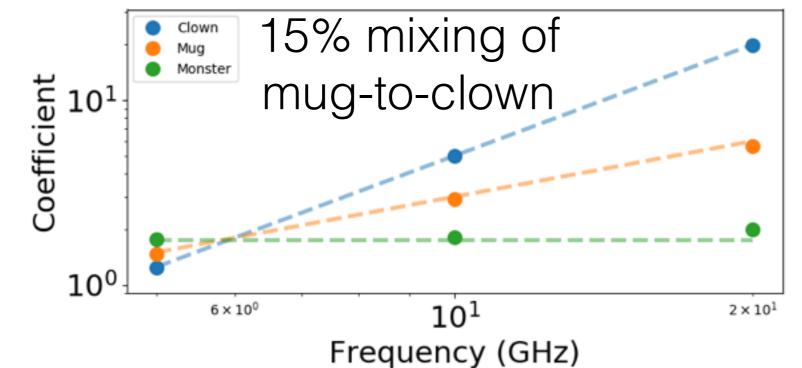
Spectral variation in components can be captured:



Stuart Harper, Radio Foregrounds 2018

Spectral variation in components can be captured:

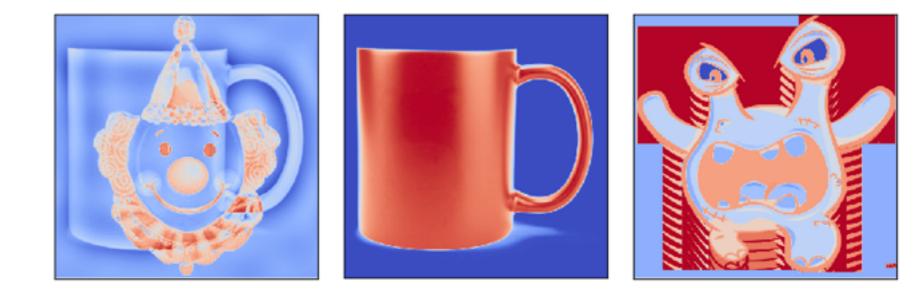


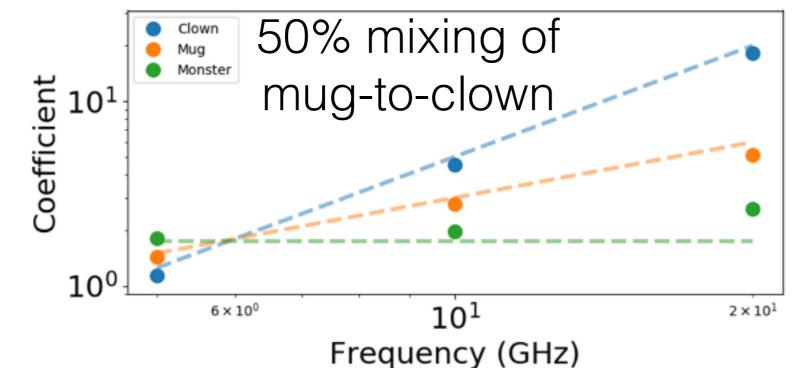


Bias' from: template correlations template defects gradients spectral variations

Stuart Harper, Radio Foregrounds 2018

Spectral variation in components can be captured:

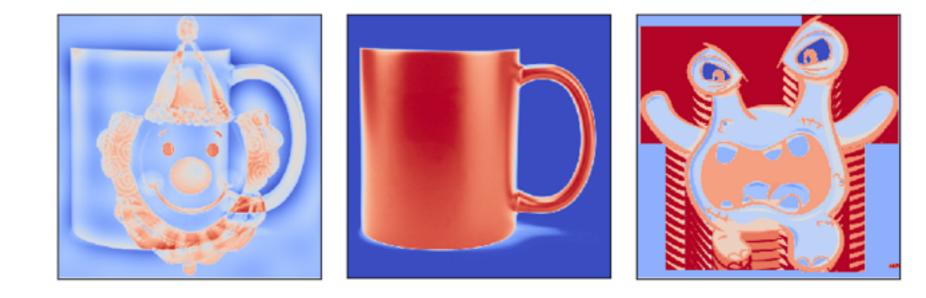


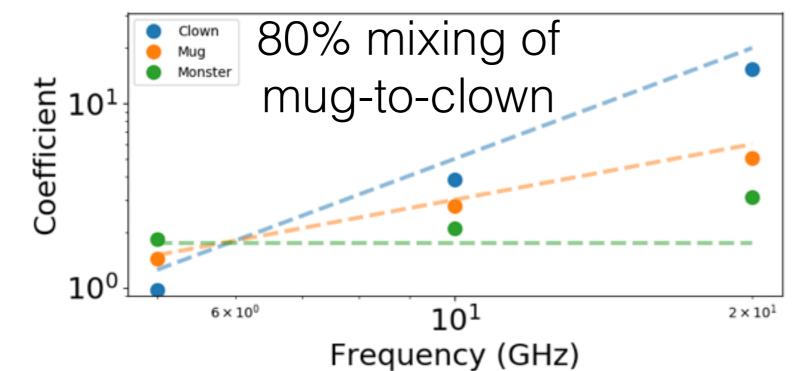


Bias' from: template correlations template defects gradients spectral variations

Stuart Harper, Radio Foregrounds 2018

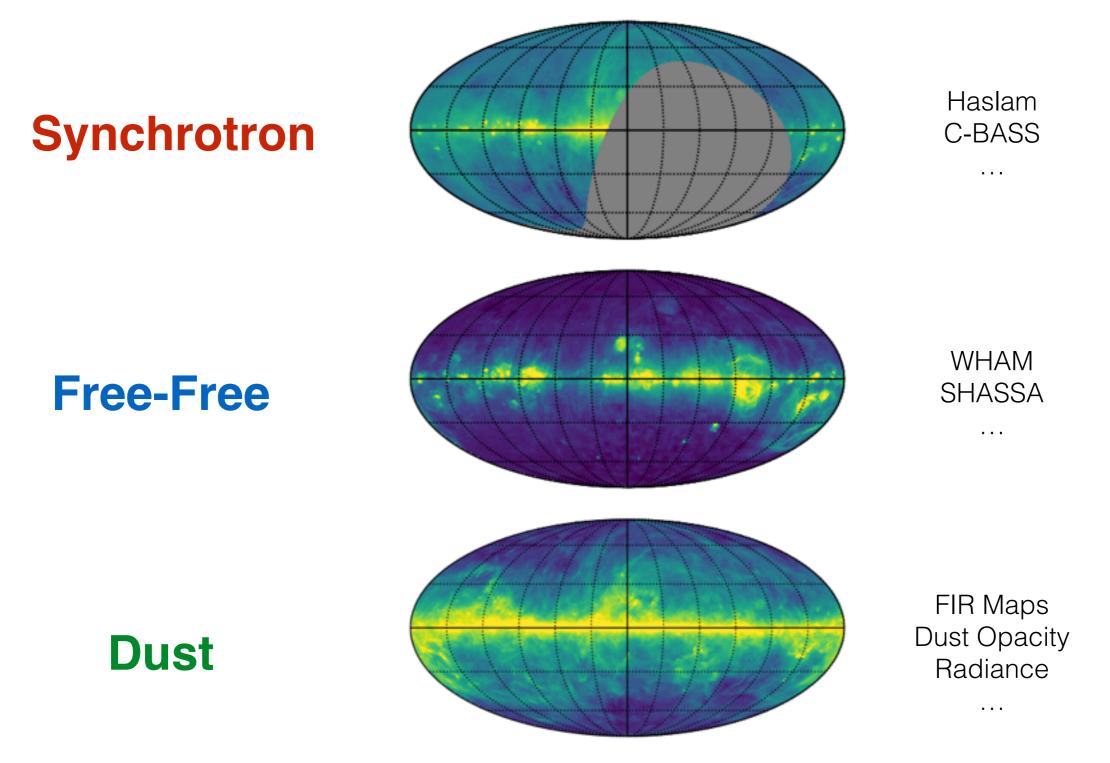
Spectral variation in components can be captured:





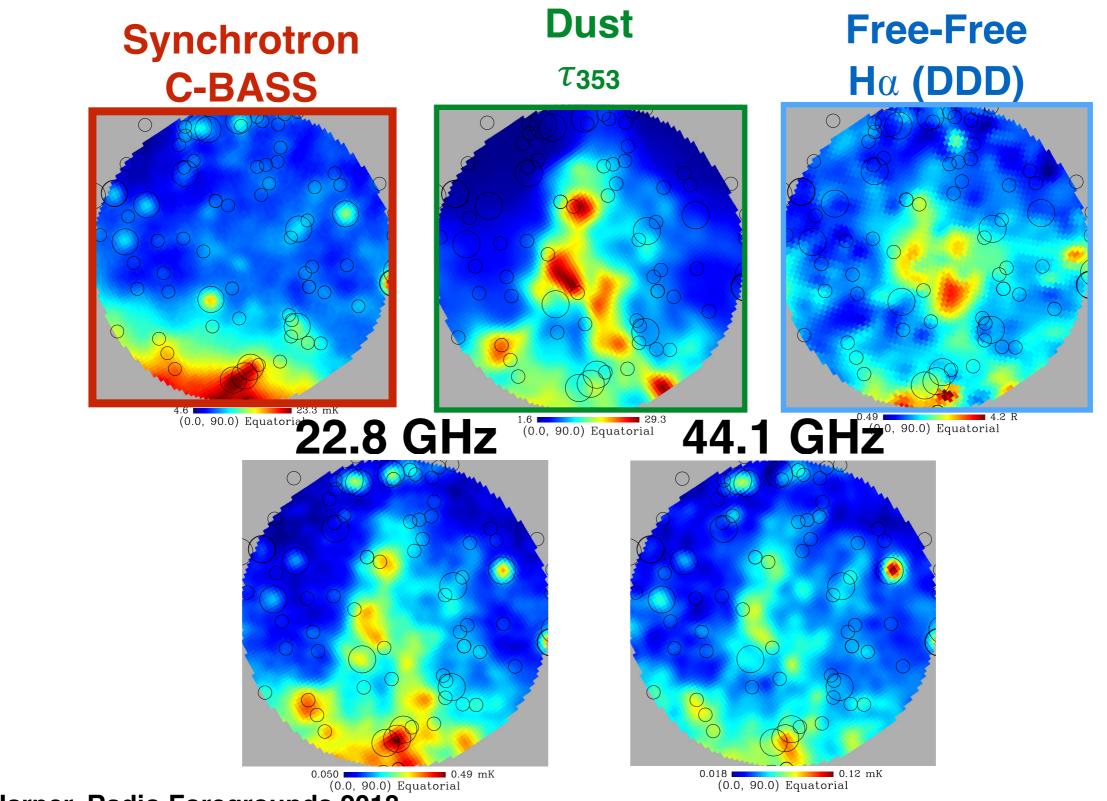
Bias' from: template correlations template defects gradients spectral variations

Stuart Harper, Radio Foregrounds 2018



NCP

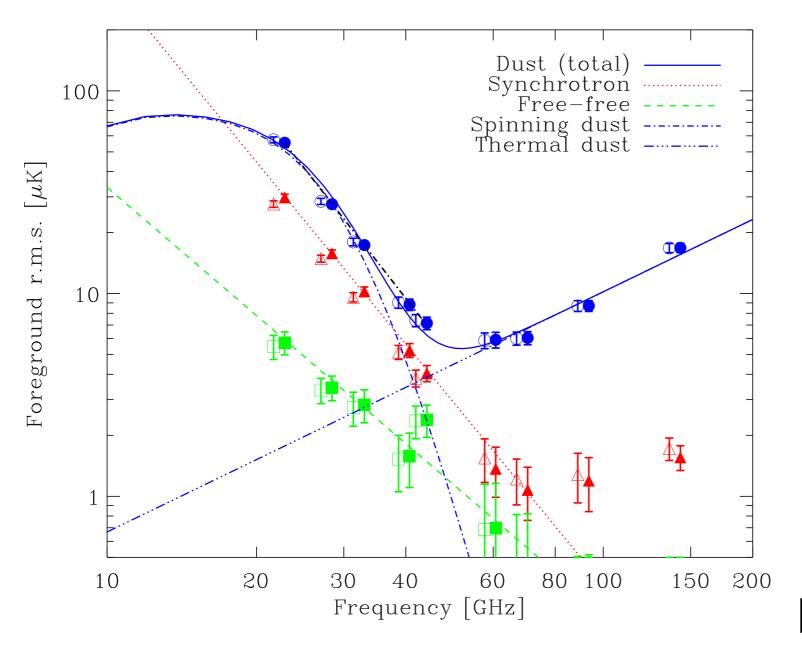
C-BASS: Constraining diffuse Galactic radio emission in the North Celestial Pole region, *Dickinson, C. + C-BASS Collaboration*, 2018



NCP Results

C-BASS: Constraining diffuse Galactic radio emission in the North Celestial Pole region, *Dickinson, C. + C-BASS Collaboration*, 2018

RMS of individual NCP components



Comparison with **Commander** expected values (22.8GHz)

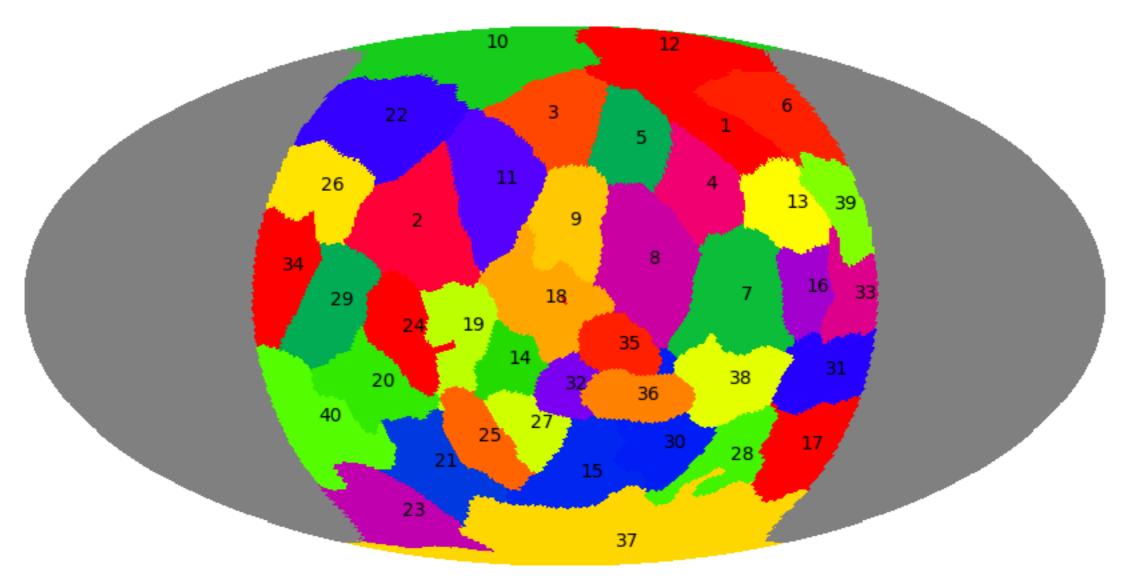
Component	CC	Planck 2015
Synchrotron	29.9 ± 1.1	11.8
Free-free	5.7 ± 0.7	46.9
AME / dust	55.4 ± 1.8	45.4
Thermal dust		1.2
Total foreground	88.8 ± 3.6	84.0

Larger AME component... 60% of total emission

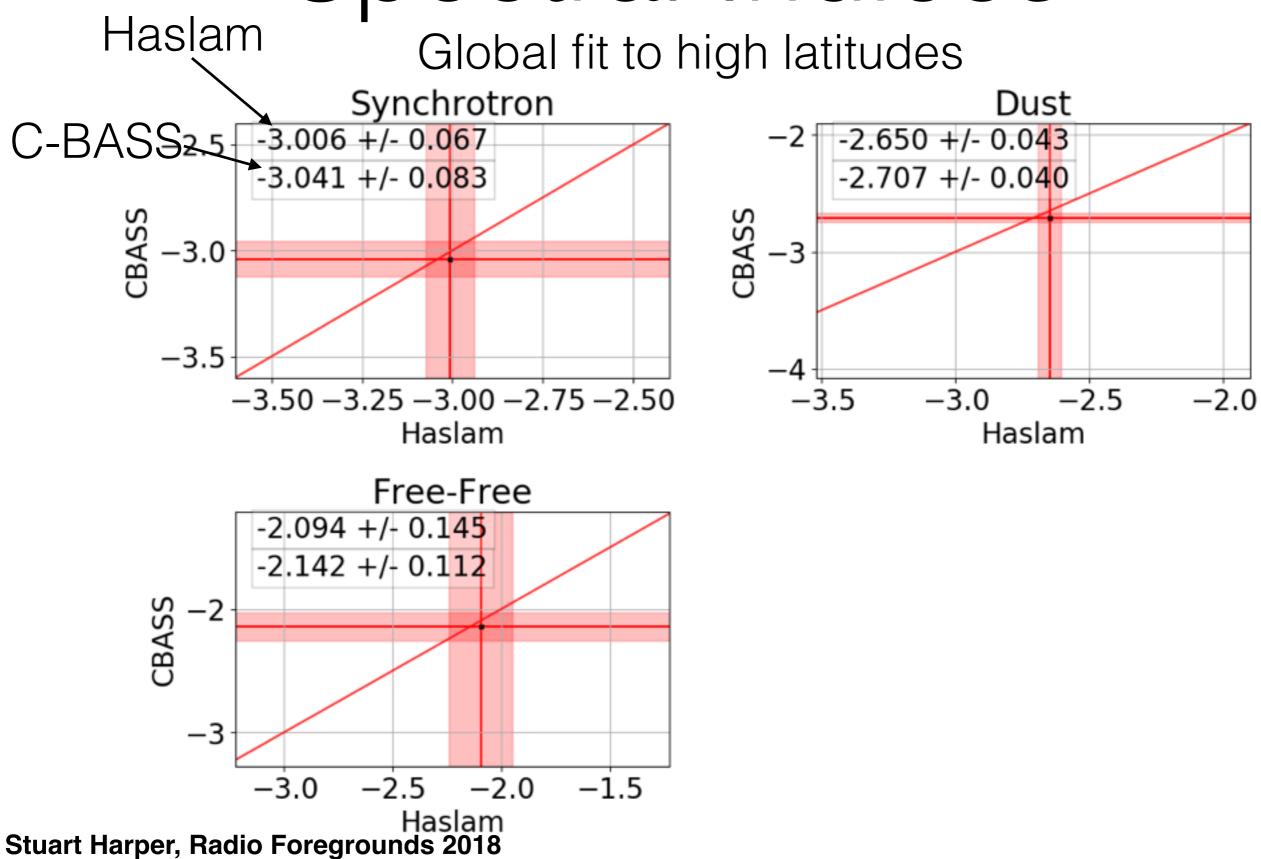
b_dust = 9.59 K/tau_353

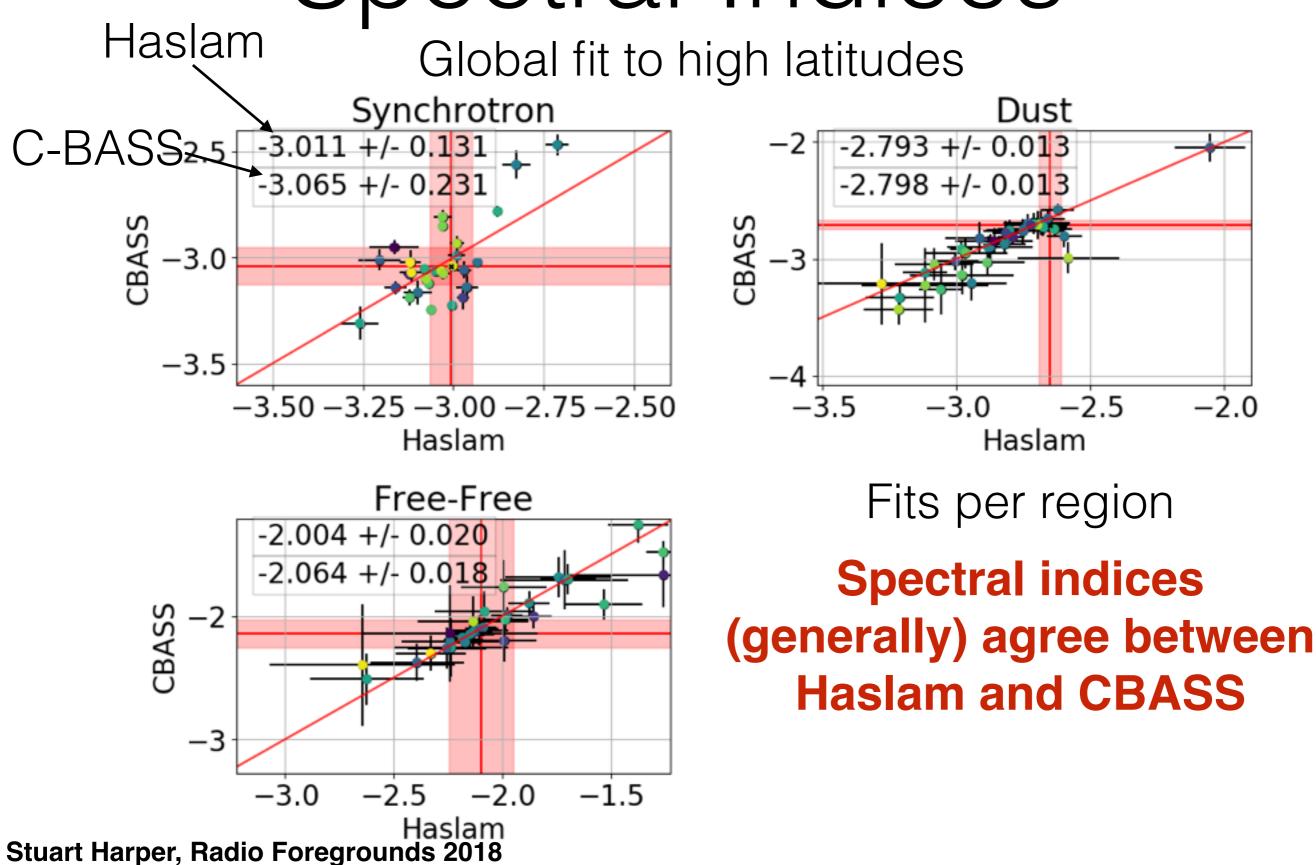
Looking at the whole sky

Break sky up into distinct regions (see Luke's talk)...

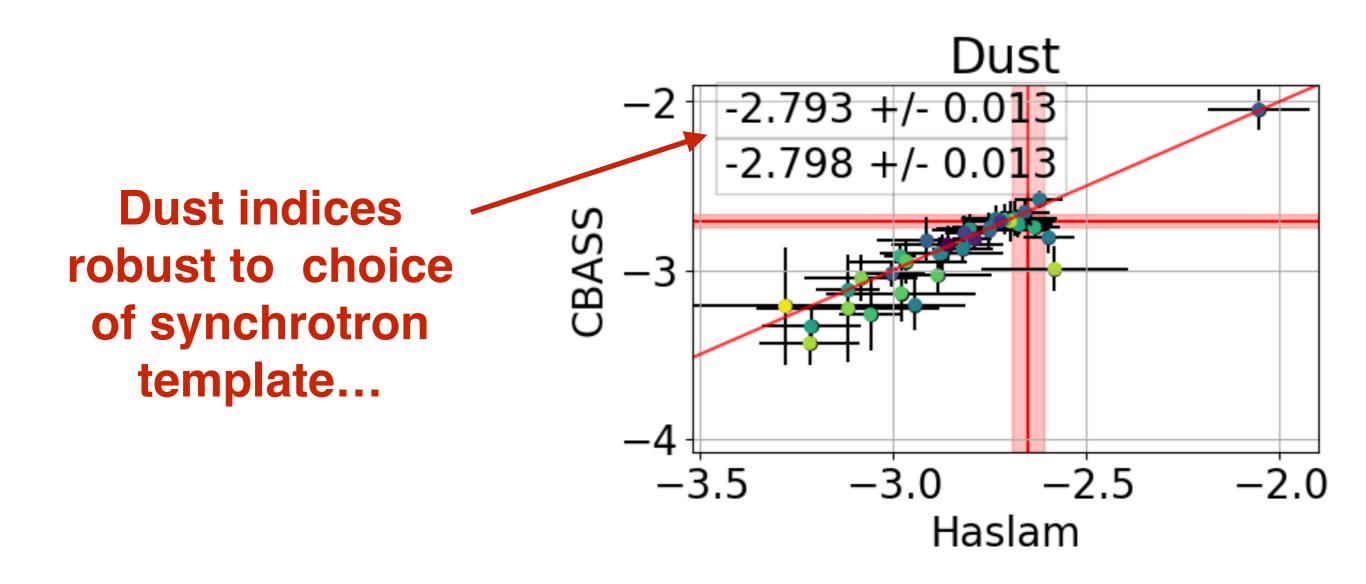


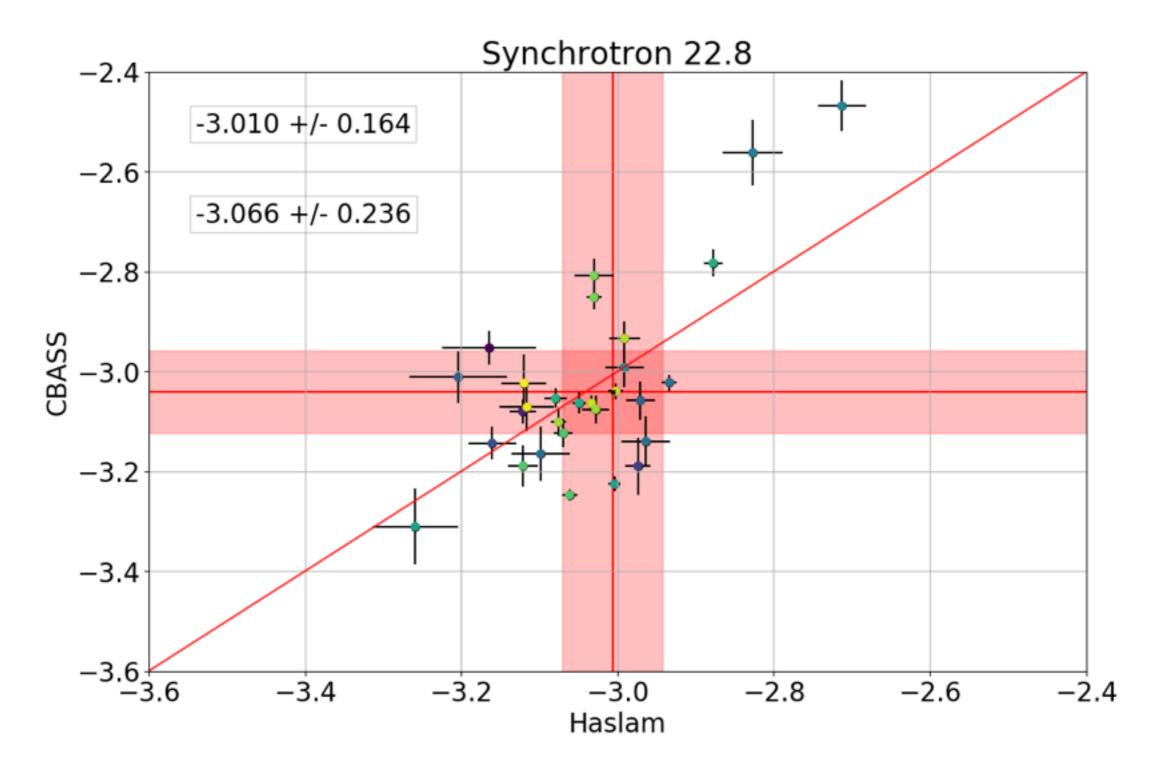
Mollweide project looking down on the NCP Stuart Harper, Radio Foregrounds 2018



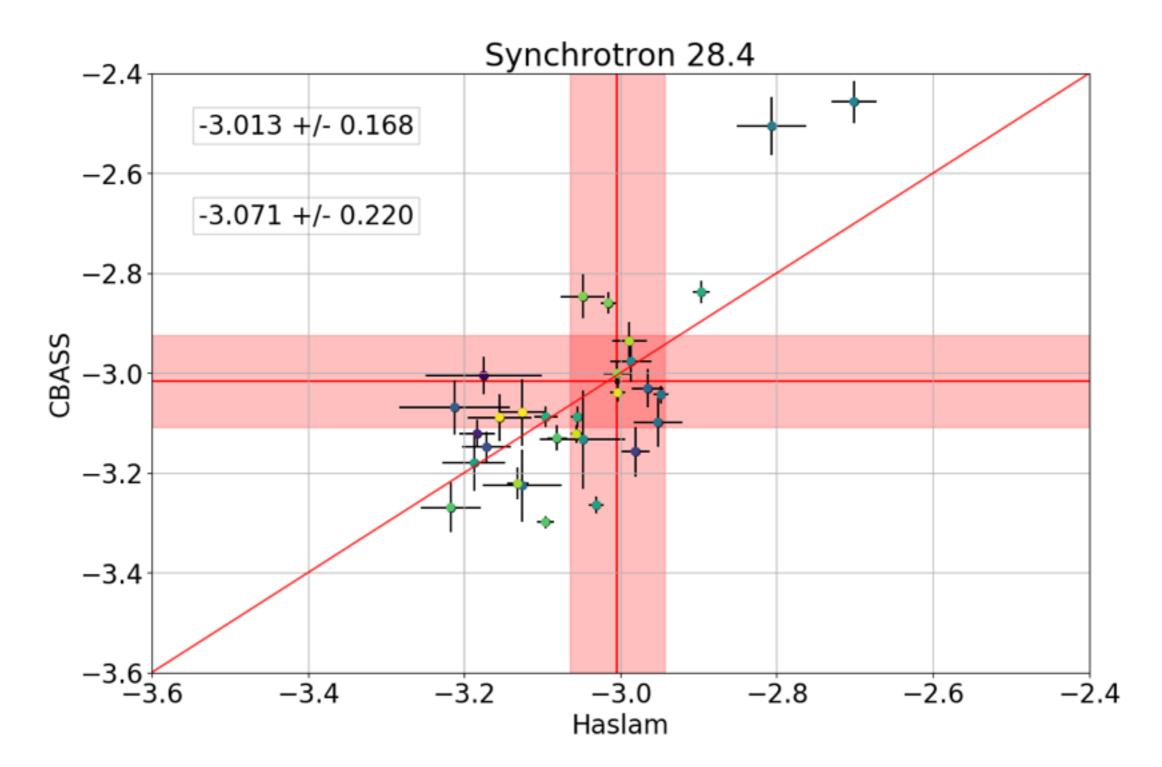


Global fit to high latitudes

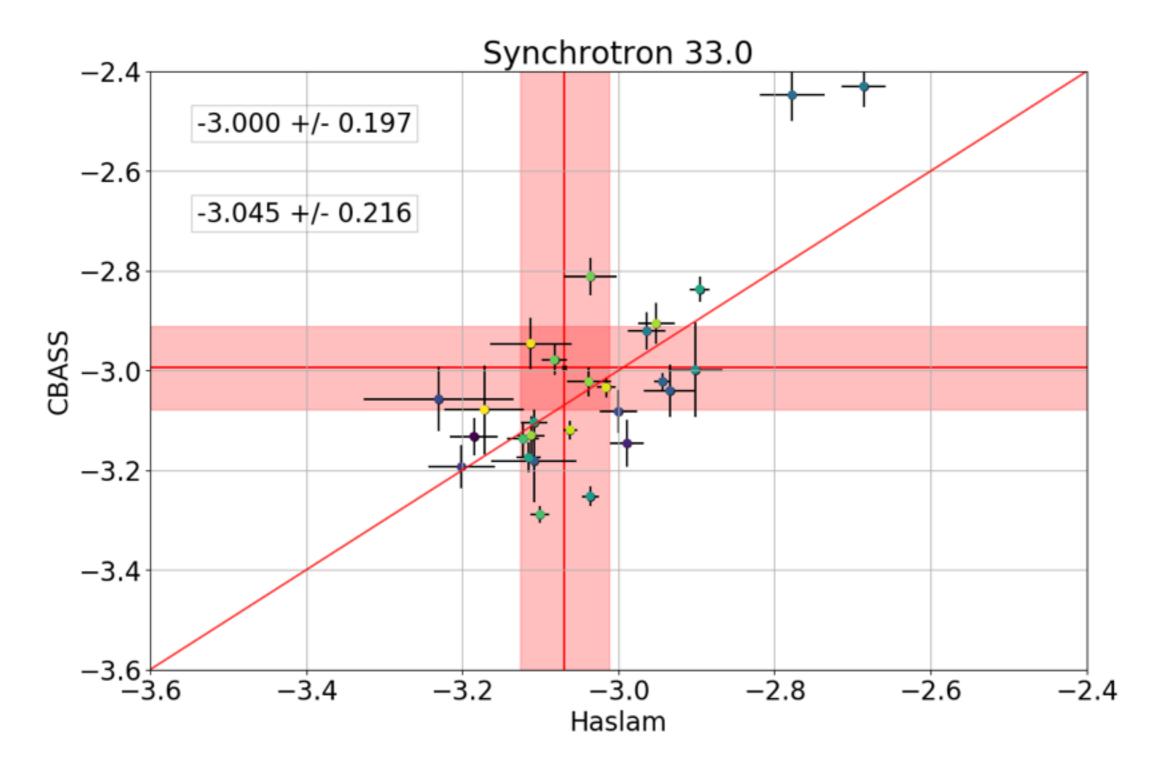




Stuart Harper, Radio Foregrounds 2018

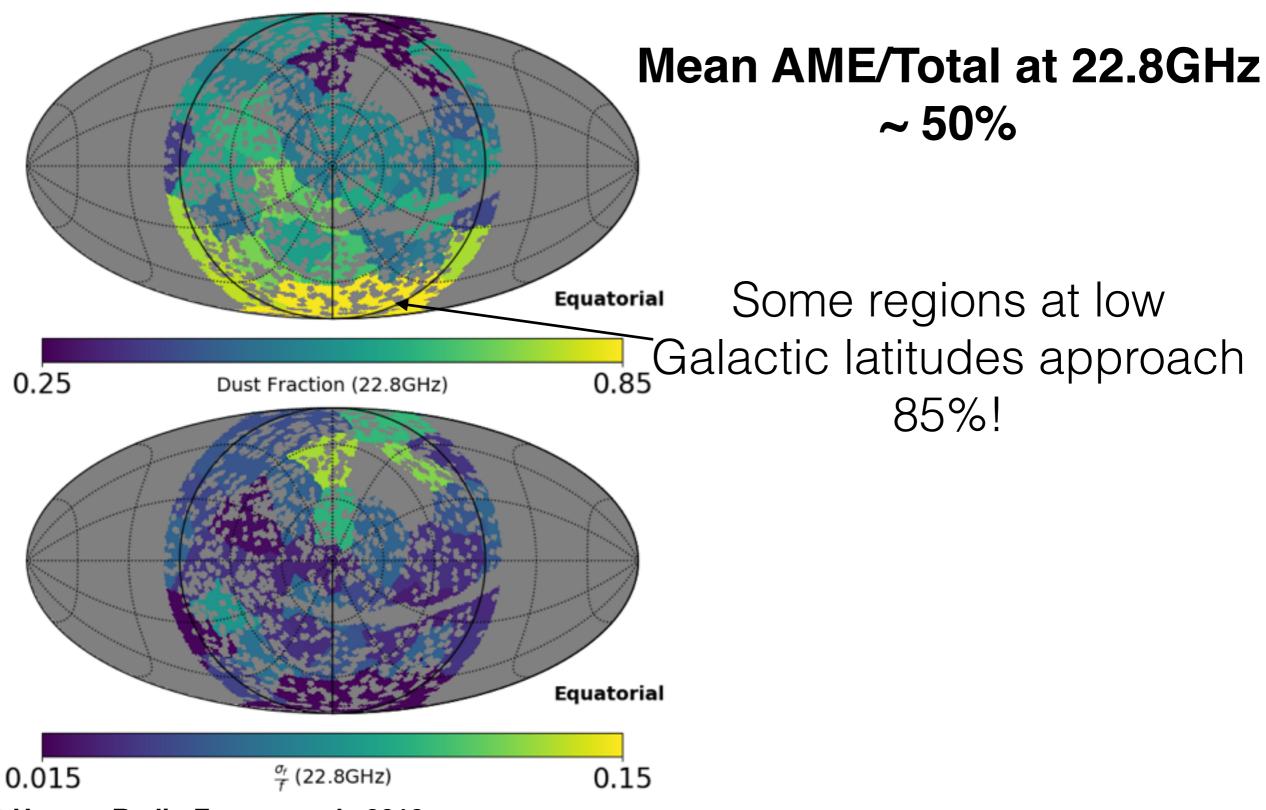


Stuart Harper, Radio Foregrounds 2018

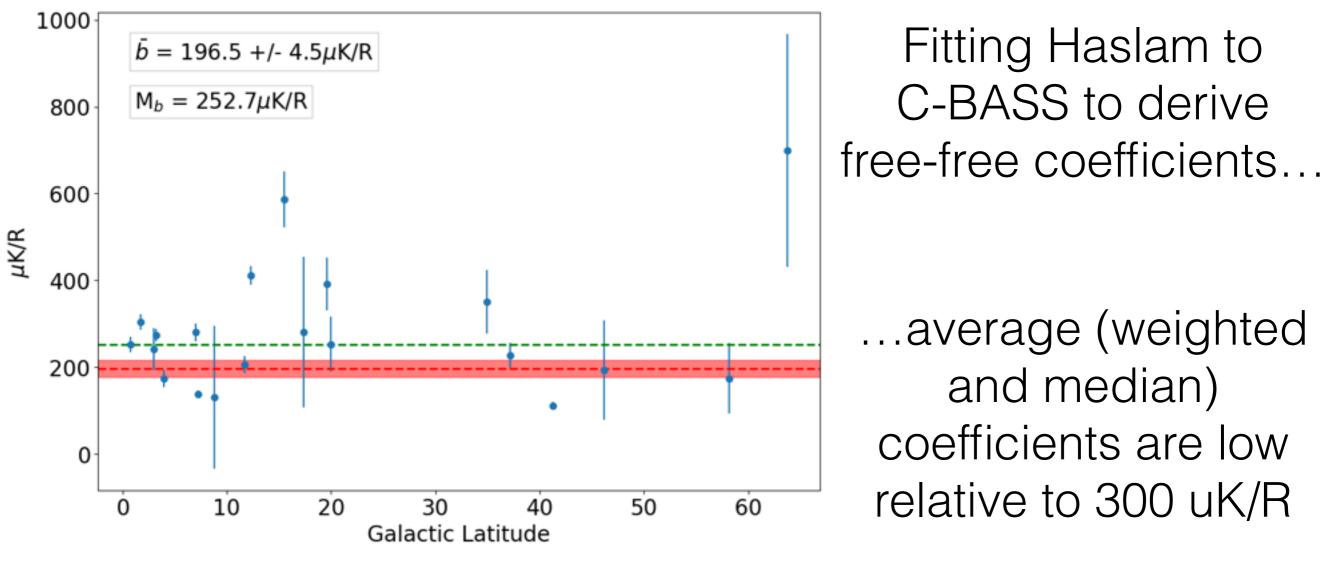


Stuart Harper, Radio Foregrounds 2018

AME Fractions



Ha Ratios



Possible evidence for

$\mbox{H}\alpha$ light dust scattering at the 15 to 30% level

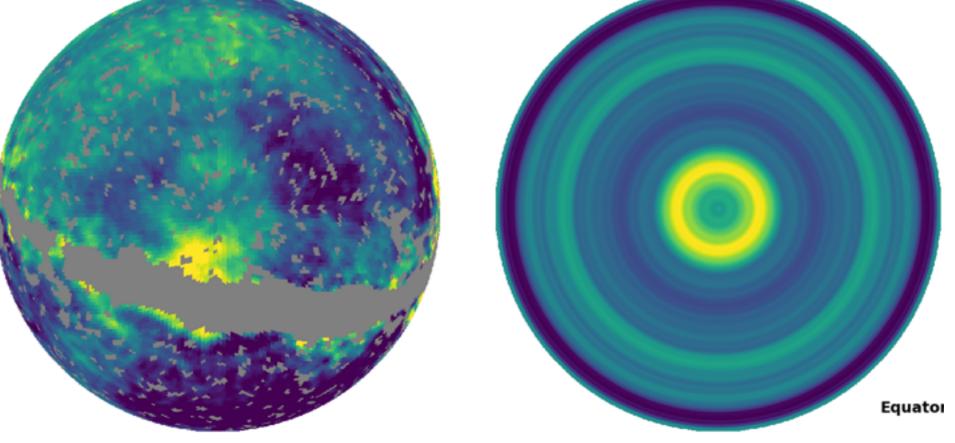
at low to medium latitudes...

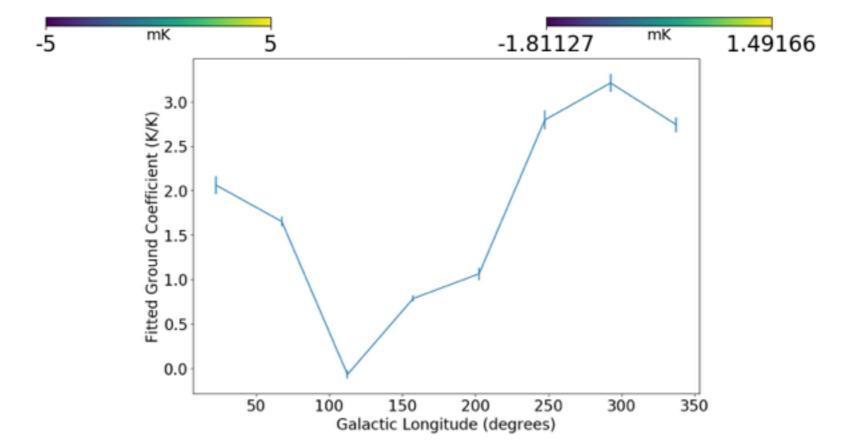
Conclusions

- Template fitting is a robust method for studying foreground emission on large scales (subject to caveats!)
- C-BASS, on average, as a synchrotron template is consistent with Haslam
- However, there are lots interesting individual regions that deviate up to the 10% level in dust and synchrotron.
- H α scattering up to the 30% level with lots of regional variability

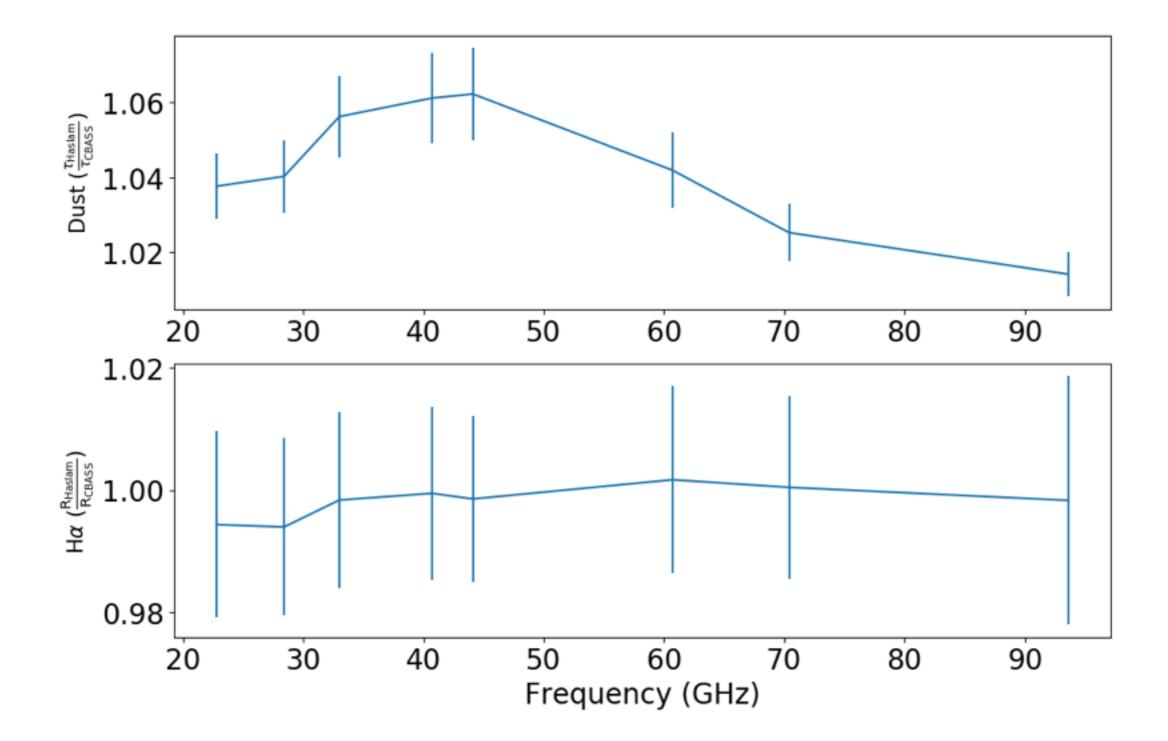
Extra Slides

Ground Contamination

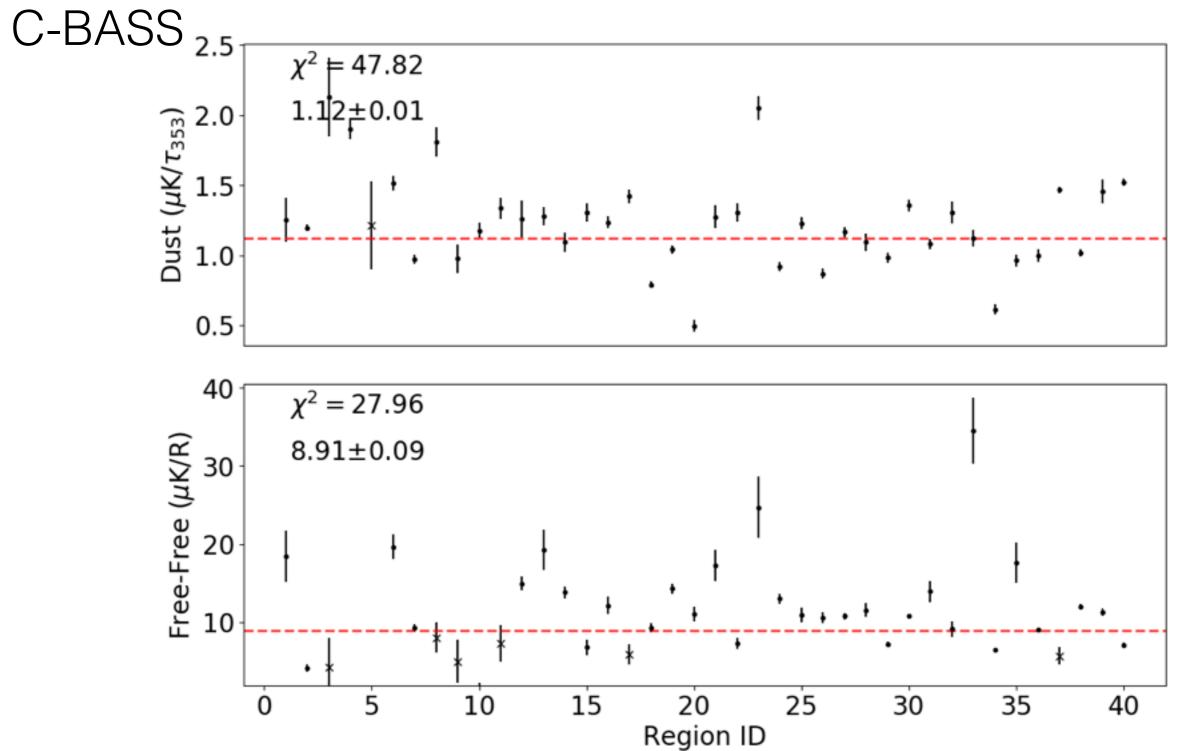




Ratios of Coefficients



Ratios of Coefficients



Ratios of Coefficients

Haslam

