

# The BRITE SONG

Aldebaran, Beta Cas et al. – the headlines



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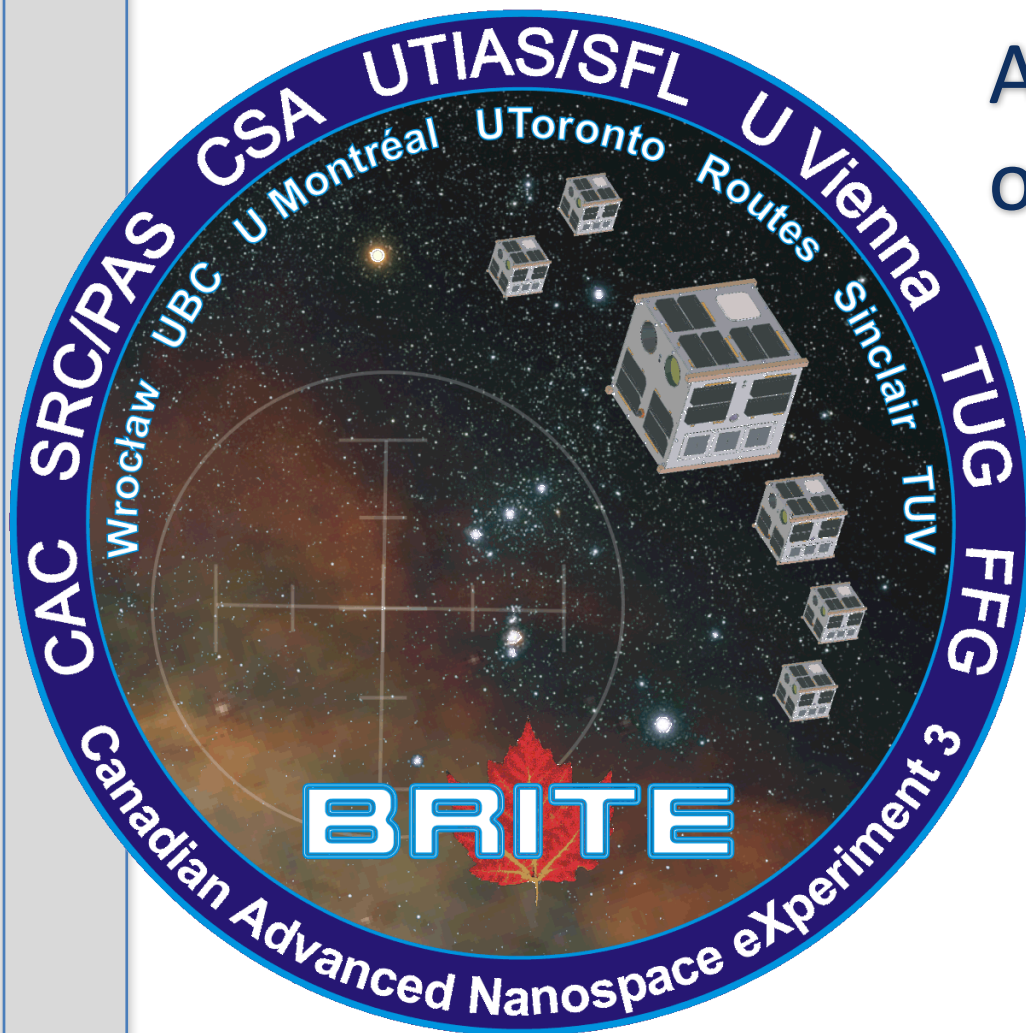
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**Konstanze Zwintz** (Innsbruck University)

**Thomas Kallinger, Werner W. Weiss** (Vienna Univ.)

# BRight Target Explorer

A Constellation of 5  
operational nanosatellites:  
20 x 20 x 20 cm



**AUSTRIA:** 02/2013  
UniBrite (R) & TUGSat (B)



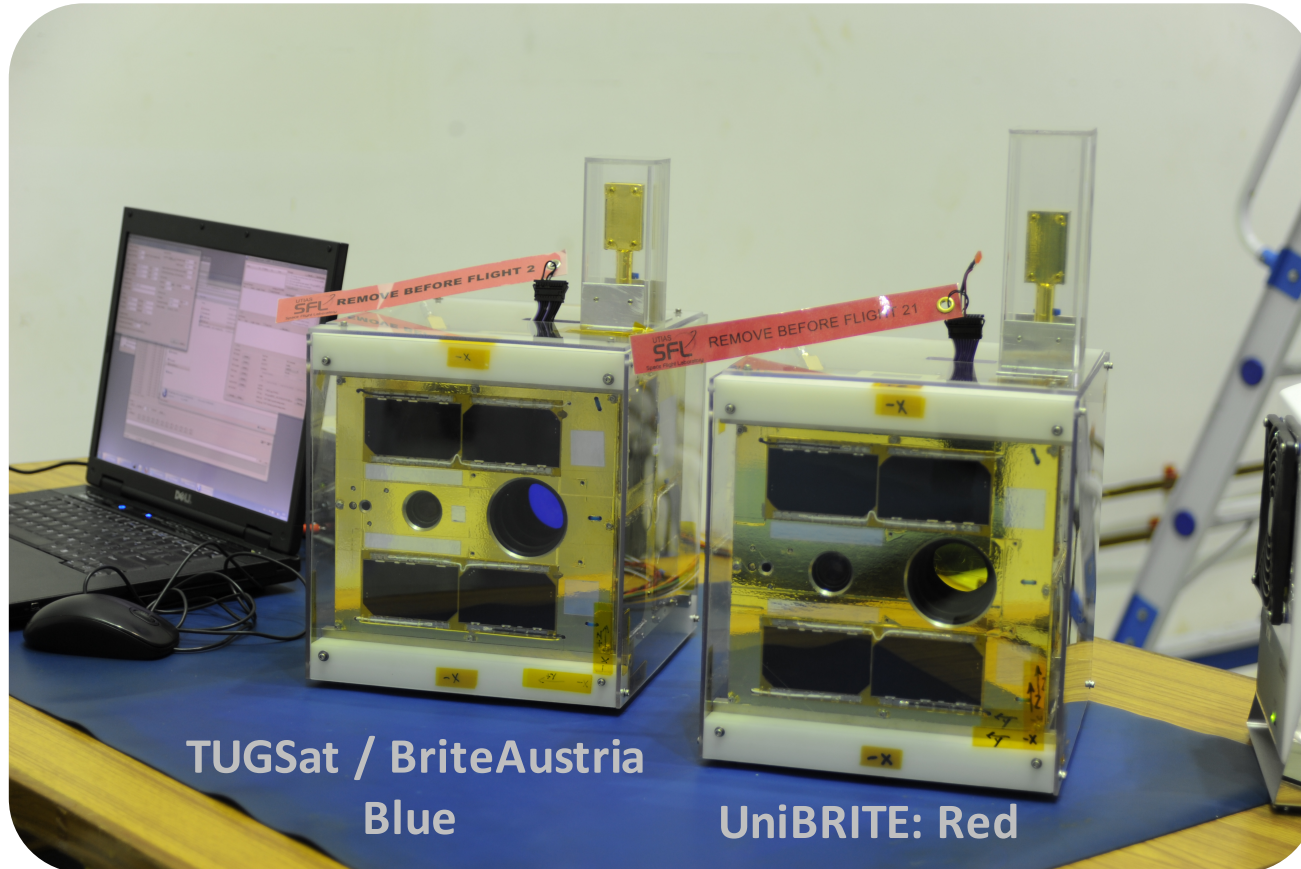
**CANADA:** 06/2014  
Toronto (R) [Montreal, B]



**POLAND:** 11/2013 & 08/2014  
Lem (B) & Hewliusz(R)

# TugSat & UniBRITE

The first Austrian Satellites



TUGSat / BriteAustria  
Blue

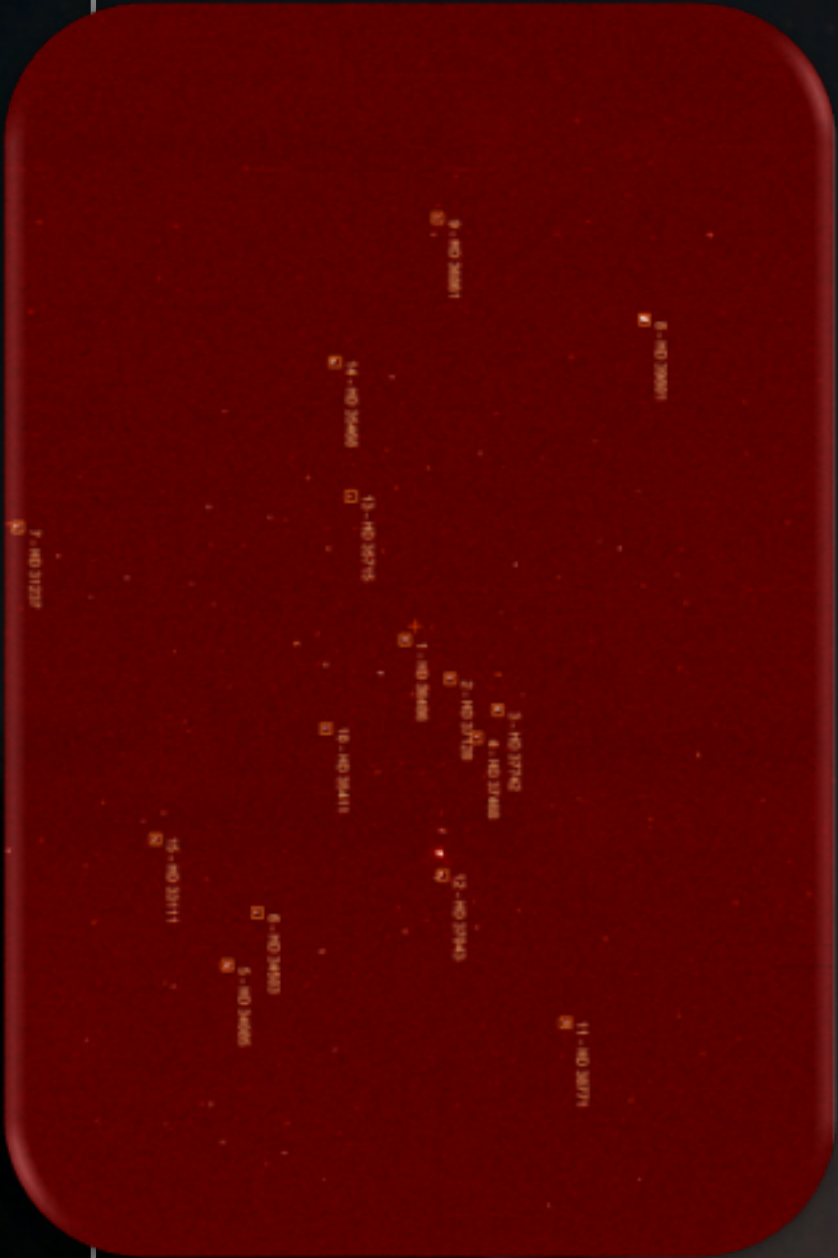
UniBRITE: Red

BLUE 400-450 nm

RED 550-700 nm

funded since 2006 (FFG/TUG, Uni Wien) – launched Feb. 2013 India

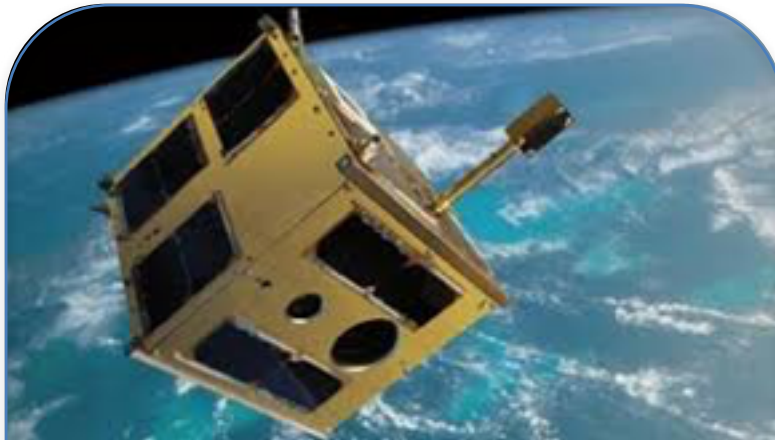
picture: pre-launch lab at ISRO India, courtesy of SFL



Field of View  
~ 24 deg

# Typical BRITE targets

Naked eye stars



- **Filters, Telescope & CCD**
  - 3 cm aperture
  - $V < 4$  mag ( $< 6.5$  mag)
- **Orbit & Attitude**
  - $\sim 100$  min  $\leftrightarrow$  14/day
  - 15 to 30 min obs/orbit
  - Obs. Season:  $< 160$  days
  - Orbit-to-Orbit: 1%

## → Target requirements

- **Photometric Variability**
  - $> 100$  ppm (brightest)
- **Frequency domain**
  - Nyquist:  $> 80 \mu\text{Hz}$   
(1 dpt/orbit)
  - Freq. Res:  $\sim 0.08 \mu\text{Hz}$

## → Optimal Targets

- Heat-driven pulsators
- Long-periodic giants

→ **32 camp.:  $> 550$  stars (2018)**

# Parallel SONG/BRITE Campaigns

Projects suited to exploit the synergies of both instruments



## Giant: Aldebaran (DDT, P7 & P8)

- Direct comparison of photometric and radial velocity amplitudes (Talk GH)

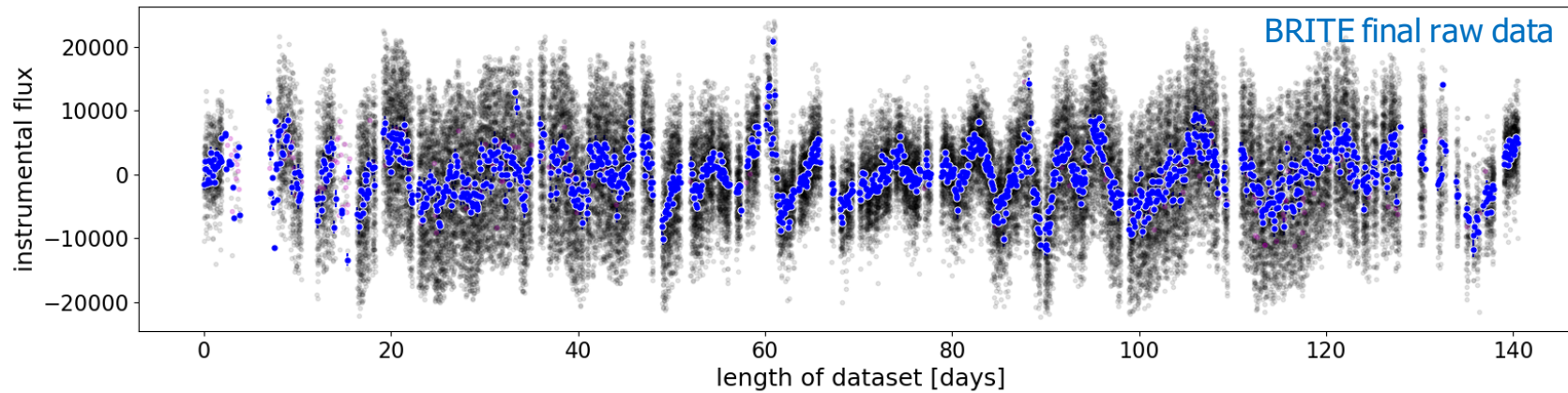


## $\delta$ Scuti: Beta Cas (P8)

- Analysis of absorption line profile variations (Fourier-Parameter-Fit, Moments)

Both programs: Slit #6 (R=90.000)

# $\alpha$ Tau: Aldebaran



## **BRITE Lem: Blue**

Exposure: 3"  
140 days with

50600 individual datapoints  
1560 orbital median points

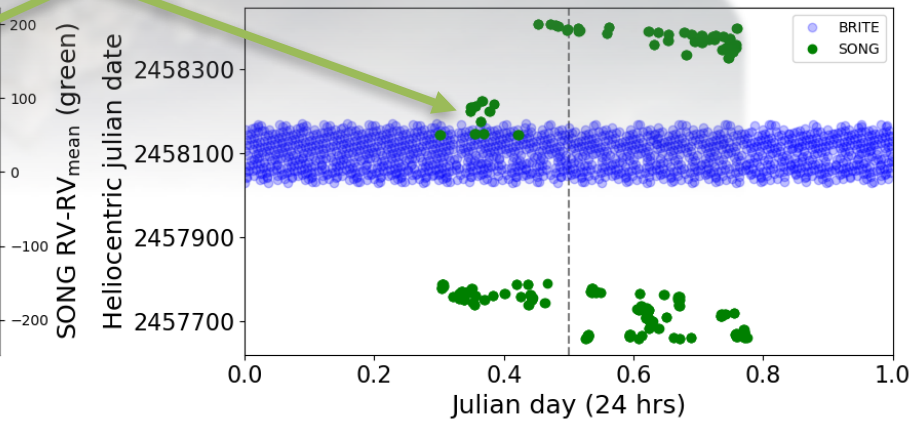
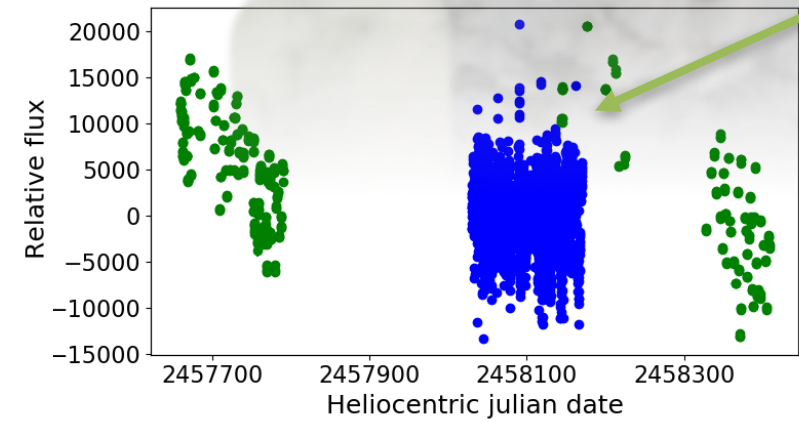
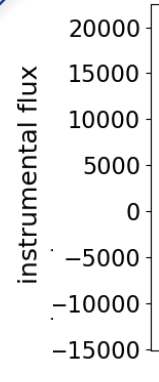
## **BRITE Toronto: Red**

Exposure: 5" (Science driven)  
Saturated

→ Halo photometry

$\alpha$

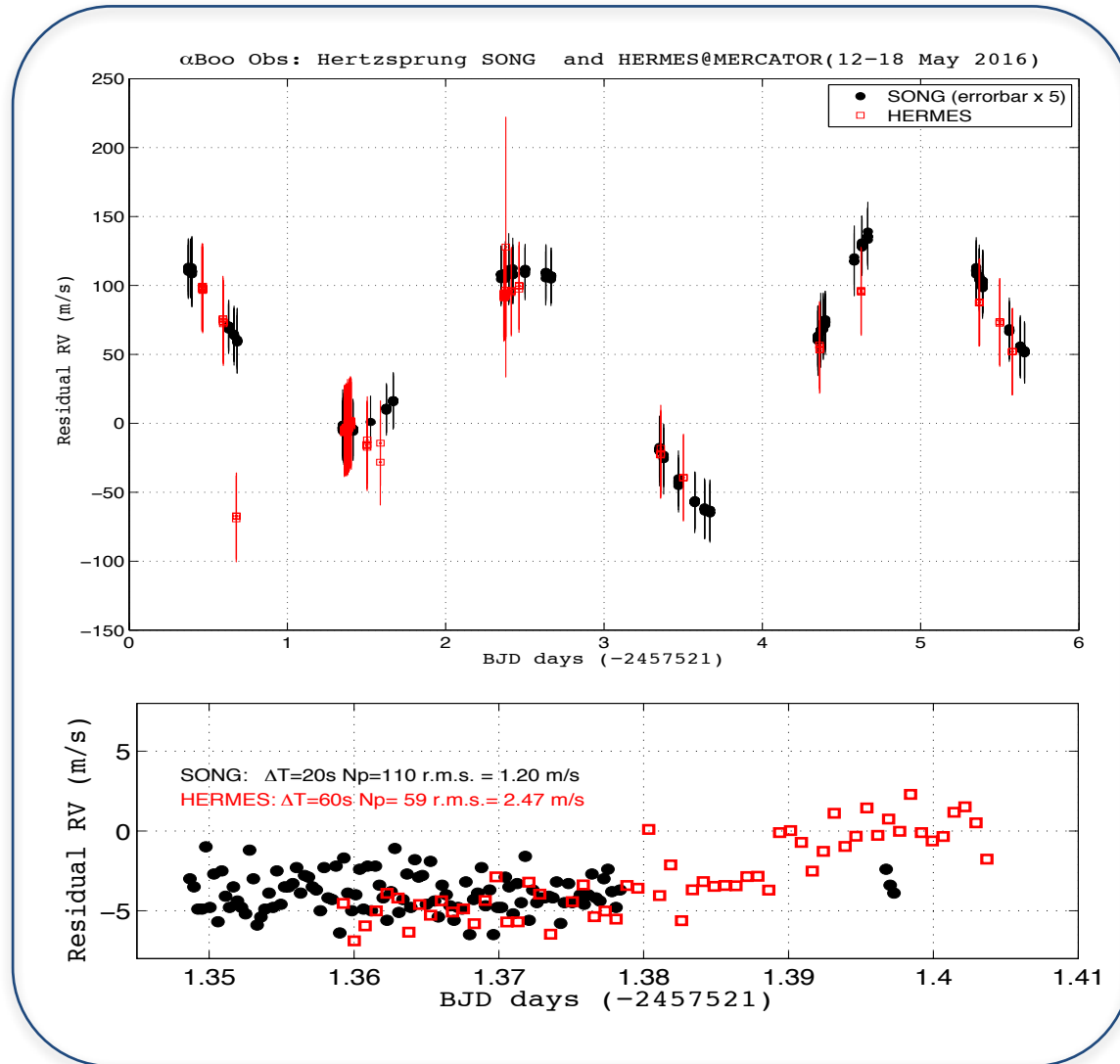
2018, February



*Aldebaran b* (Hatzes+ 2015, Farr+ 2018)  
 $M=6.47 \pm 0.53 M_{\text{jup}}$ ;  $P=629.96 \pm 0.9$  days



# SONG/Hermes: Arcturus



## BRITE

- poorly populated field
- single target
- few guide stars

→ TESS is coming

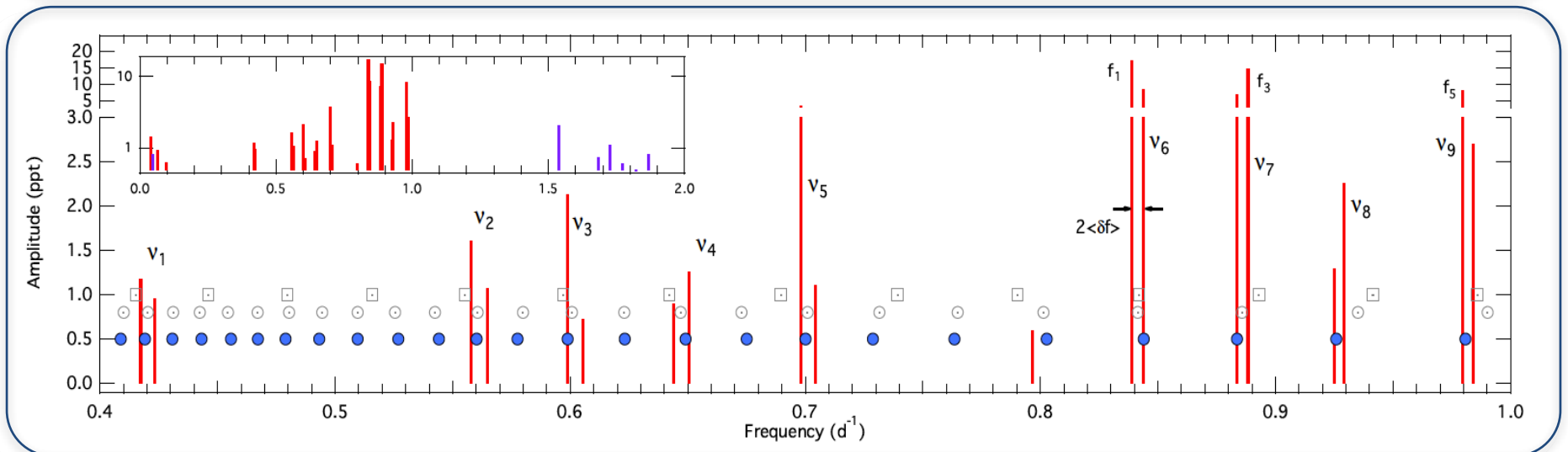
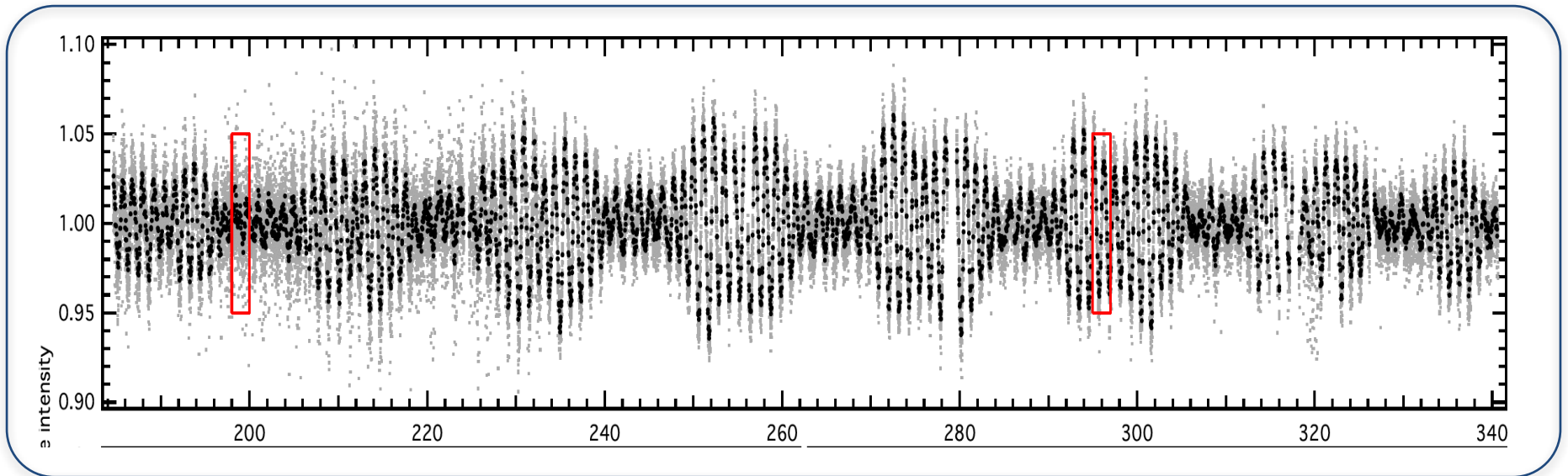
## Comparison with Hermes

- 6 day comparison (May 2016)
  - SONG: 12
  - Hermes: Sim ThAr (LRF)
- ~8 years: Ca H&K (Hermes)
  - inconclusive (back then)

→ TIGRE time series?

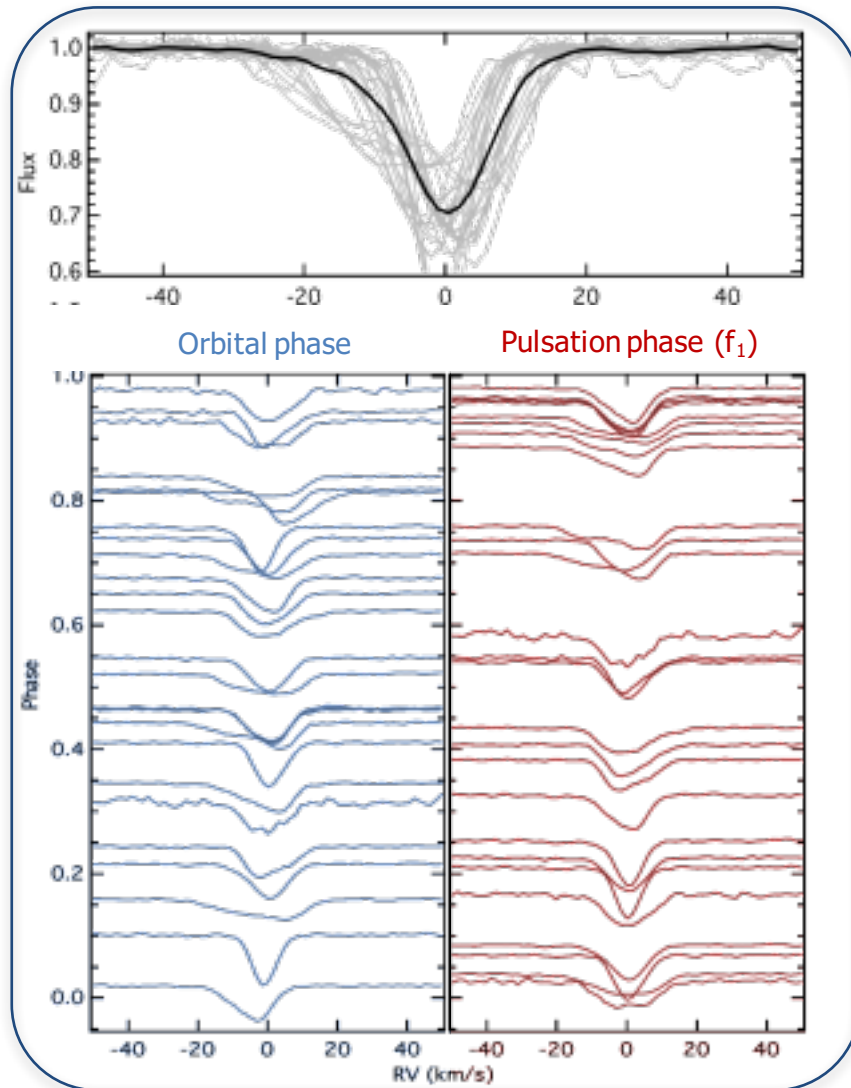
# Previous Campaigns: HD201433 (SPB)

Kallinger, Weiss, Beck et al. (2017, A&A 603, A13)



# BRITE Target: HD201433

Kallinger, Weiss, Beck et al. (2017, A&A 603, A13)



## HD201433 : B9V /SPB in a triple system

- 3.3 & 154 days periods
- Frequencies extracted
- 9 rotational split modes:  
solid body rotator with spun up surface

## Observations:

- BRITE: 140 days
- Hermes: 2017: LPV

→ LPV originate from oscillations  
→ Hints for additional frequencies (spec)

- If done on full scale:
  - Mode identification  
l, m, pro/retrograde

LPV (Fe II 423.3nm) from HERMES spectra in HD 201433

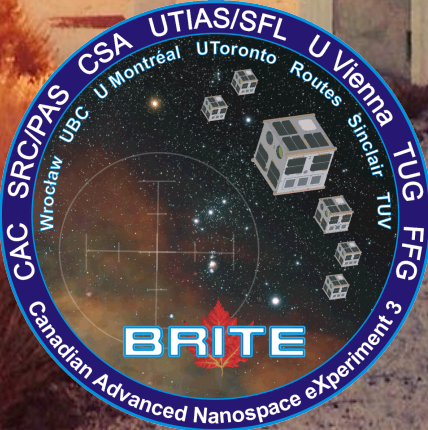


Photo: Paul Beck

The image is a composite. The top left shows a satellite in space. The background is a landscape with a sunset over a mountain range. The bottom left features a globe with several telescopes positioned around it, and the text 'SONG' and 'STELLAR OBSERVATIONS NETWORK GROUP' overlaid.

# Summary & Outlook

- **BRITE/TESS & SONG science cases are as diverse** as the number of stellar classes/objects.
  - Flexibility with SONG:  
observing mode & data products
- **Optimal combination with SONG + BRITE & K2/TESS**
  - Freq. resolution: Long-time base photometry
  - Simultaneous RV and PV
  - Diff. sensitivity to spherical harmonics  
& Potential mode Identification
- **Remaining BRITE-Constellation Mission & Archive**
  - Estimated Mission lifetime: ~2 years
  - Predefined target fields are revisited
  - Proprietary time : 1 year
  - Vicky: 'cross-branding' in archive: SONG / BRITE