

The Transiting Exoplanet Community

Early Release Science Program

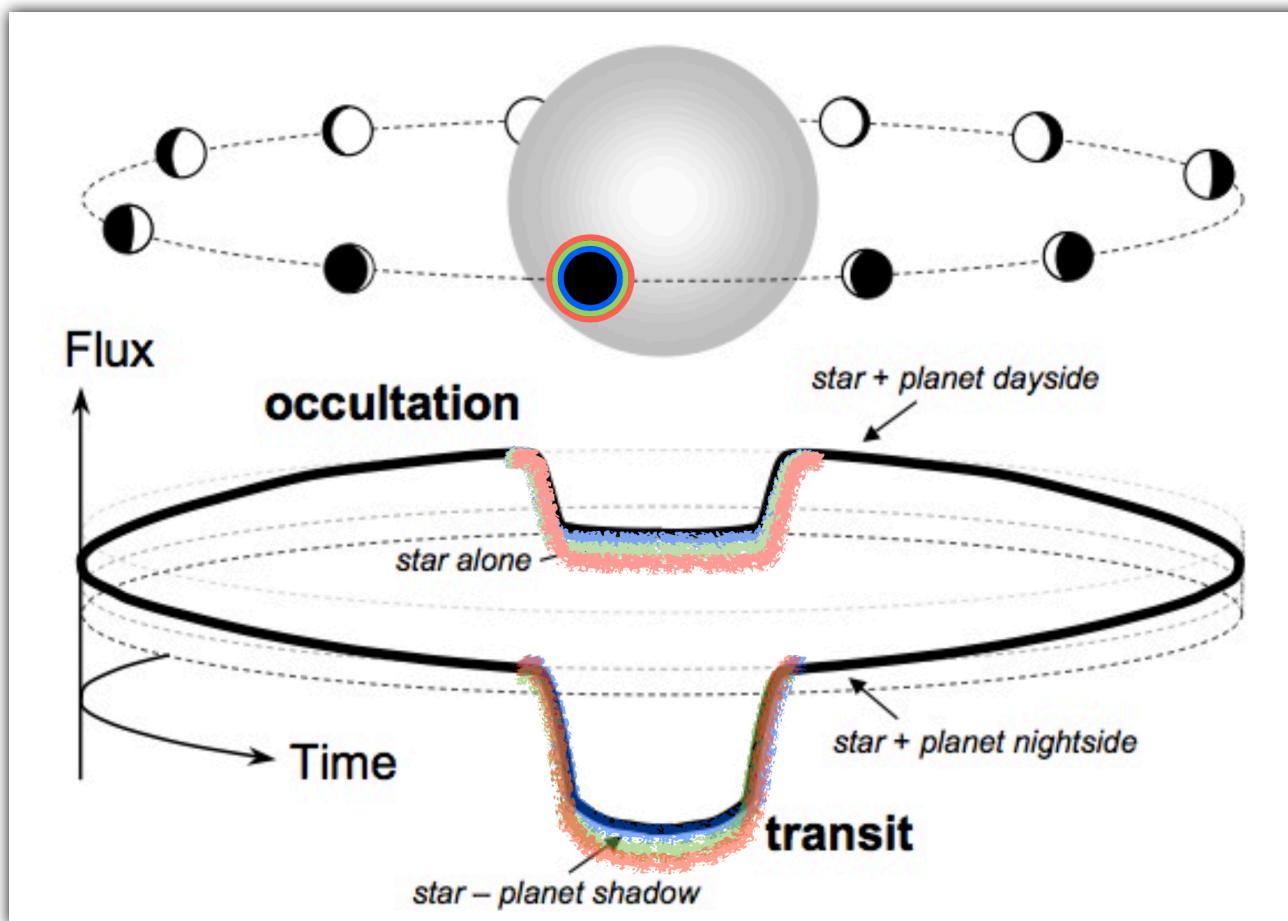
Nicolas Crouzet

And the Transiting Exoplanet ERS team

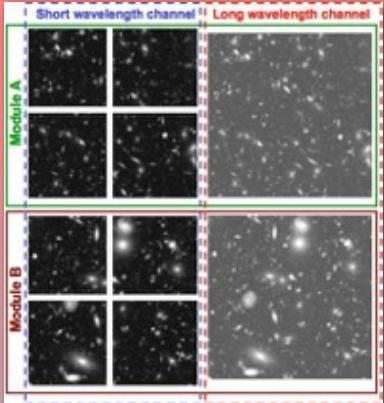
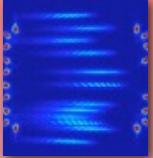
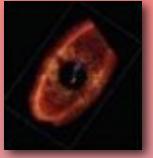
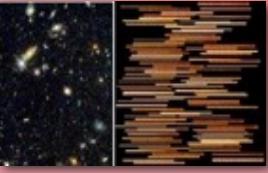
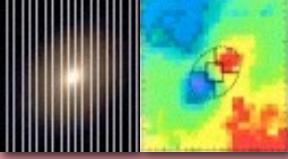
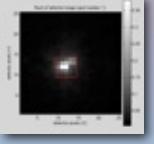
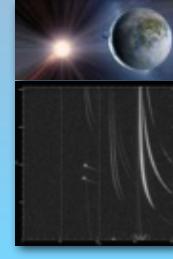
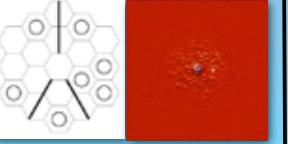
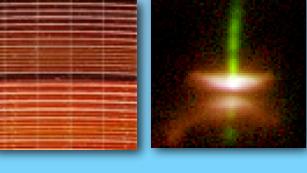


Spectroscopy of transiting exoplanets

Goal: Observe and study their atmospheres

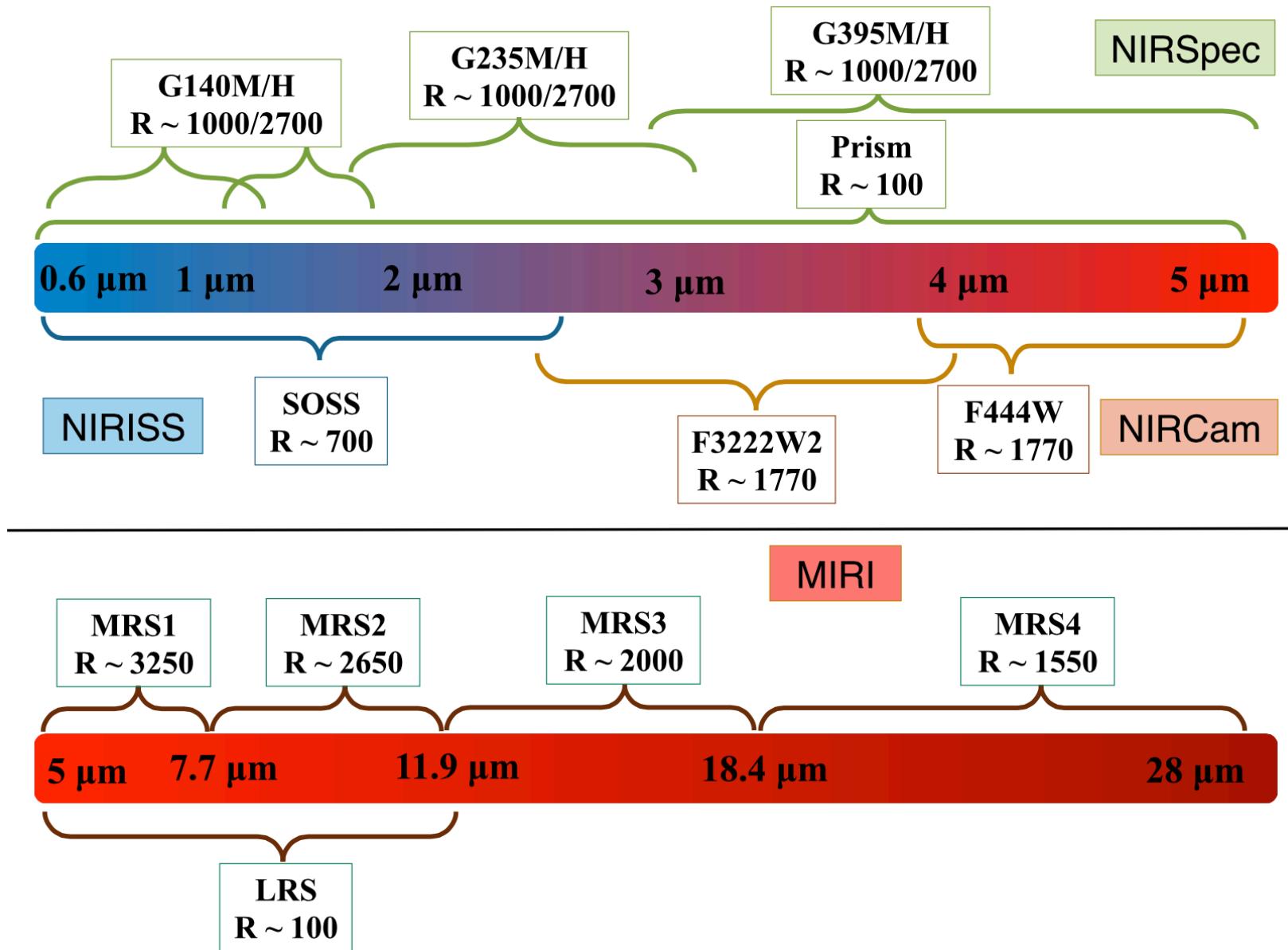


The JWST Instruments

NIRCam		NIRSpec	
 <p>Deep, wide field imaging</p>		<p>WFSC</p>  <p>Coronagraphic Imaging</p> 	<p>Multi-Object, IR spectroscopy</p>  <p>IFU spectroscopy</p> 
<p>Fine Guidance Sensor</p>  <p>Moving Target Support</p> 	<p>Slitless Spectroscopy</p>  <p>Near-IR imaging</p> 	<p>High Contrast Imaging</p>  <p>Mid-IR Coronagraphy</p> 	<p>Mid-IR, wide-field</p>  <p>IFU spectroscopy</p> 
FGS/NIRISS	MIRI		

Slide from Mark Clampin (GSFC, JWST Observatory Project Scientist)

Spectroscopic modes for transiting exoplanets



The Transiting Exoplanet ERS program

Overview:

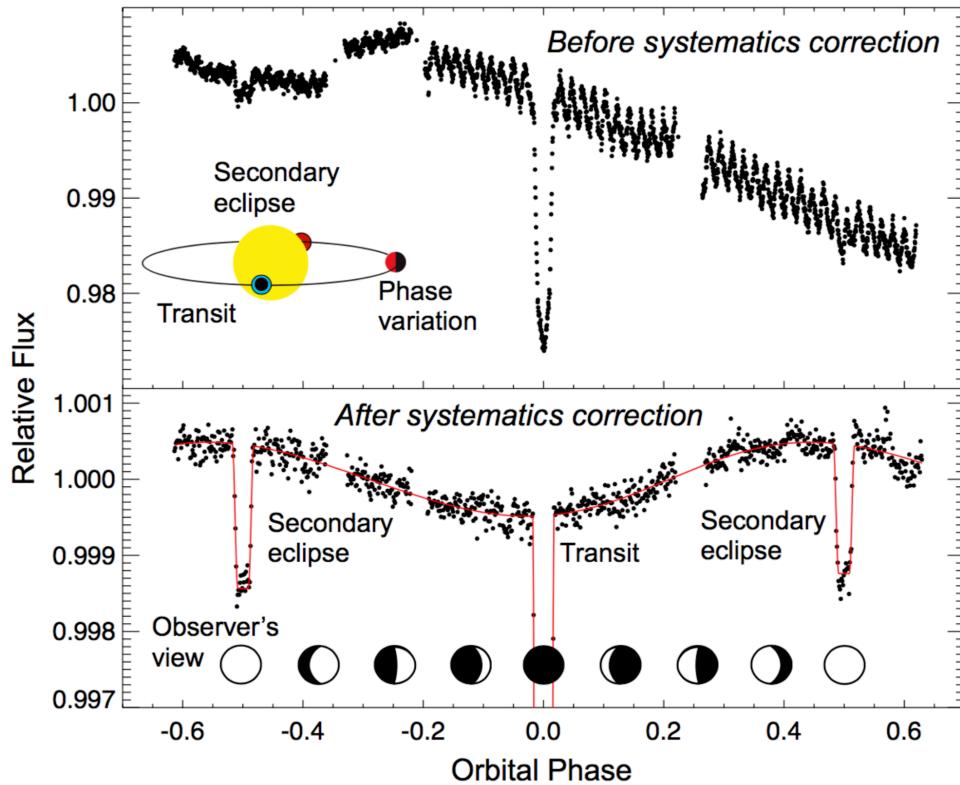
- Spectroscopy of **transiting exoplanet atmospheres**
- A **community program**: ~100 people
- PI: Natalie Batalha, co-PIs: Jacob Bean & Kevin Stevenson,
Science council: led by David Sing
- First round of ideas: *Stevenson et al. 2016, PASP 128, 4401*
Program summary: *paper in prep.*
- Divided into **four sub-programs**, all in one proposal
- **78 hours** allocated

The Team

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Berta-Thompson Zachory K.	Garland Ryan	Lines Stefan	Shkolnik Evgenya L.
Blecic Jasmina	Gibson Neale P.	Lopez-Morales Mercedes	Showman Adam P.
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Crossfield Ian J.M.	Hong Yucian	Marchis Franck	Todorov Kamen O.
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Cubillos Patricio E.	Ingalls James G.	May Erin M.	Tsiaras Angelos
Decin Leen	Iro Nicolas	Mayne Nathan	Tucker Gregory S.
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Desert Jean-Michel	Kataria Tiffany	Morello Giuseppe	Waalkes William C.
de Val-Borro Miguel	Kendrew Sarah	Morley Caroline V.	Wakeford Hannah R.
de Wit Julien	Kempton Eliza M.-R.	Moses Julianne I.	Waldmann Ingo P.
Dragomir Diana	Kilpatrick Brian M.	Nikolov Nikolay	Weaver Ian
Drummond Benjamin	Knutson Heather A.	Palle Enric	Wheatley Peter J.
Endl Michael	Kreidberg Laura	Parmentier Vivien	Zellem Robert T.

Context

Spitzer/IRAC 4.5 μm phase curve for the hot Jupiter HD 189733b (*Knutson et al. 2012*)



Evaluating and correcting for systematic effects is essential

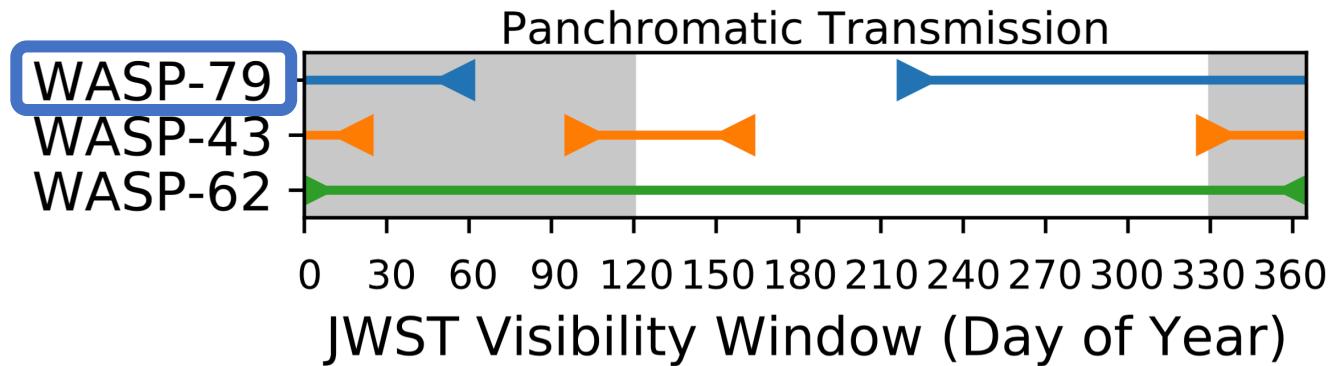
Our program will test transits, eclipses, and phase curves

Program 1: “Panchromatic Transmission”

- Chairs: Hannah Wakeford, David Sing, Kevin Stevenson (42 members)
- Goals: - **Compare and validate the observing modes** available for transmission spectroscopy in the 1 – 5 μm range
 - Extract a hot Jupiter transmission spectrum from 1 to 5 μm , measure its **atmospheric composition, metallicity, C/O ratio**
- Observations: Four transits of one hot Jupiter using different modes

Program 1: “Panchromatic Transmission”

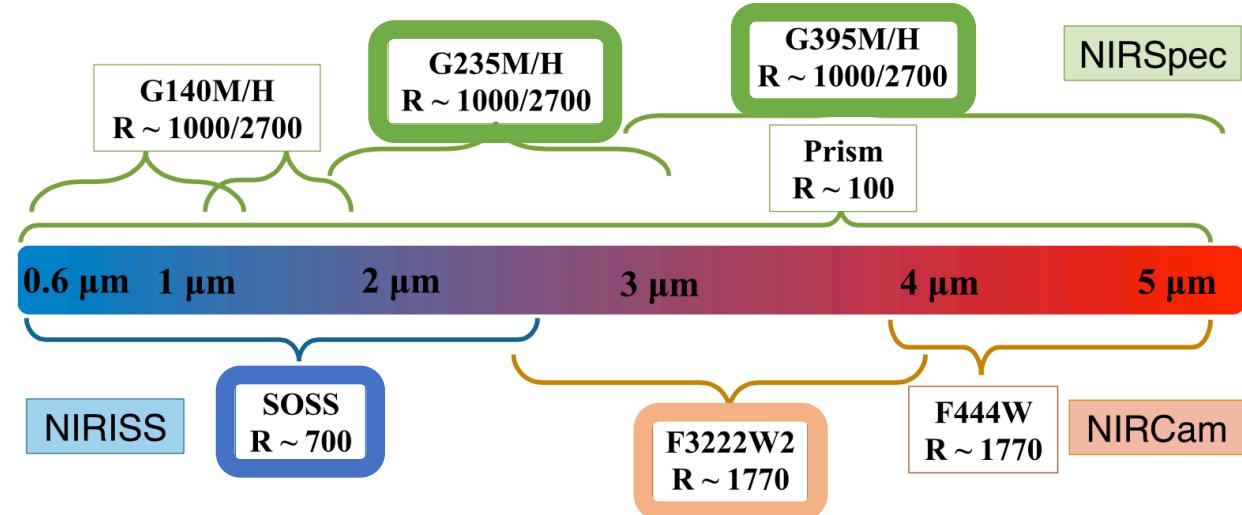
Which target?



- WASP-79b: - A highly bloated hot Jupiter ($0.9 M_{Jup}$, $1.7 R_{Jup}$)
- Water vapor absorption detected with HST WFC3

Program 1: “Panchromatic Transmission”

Which modes?

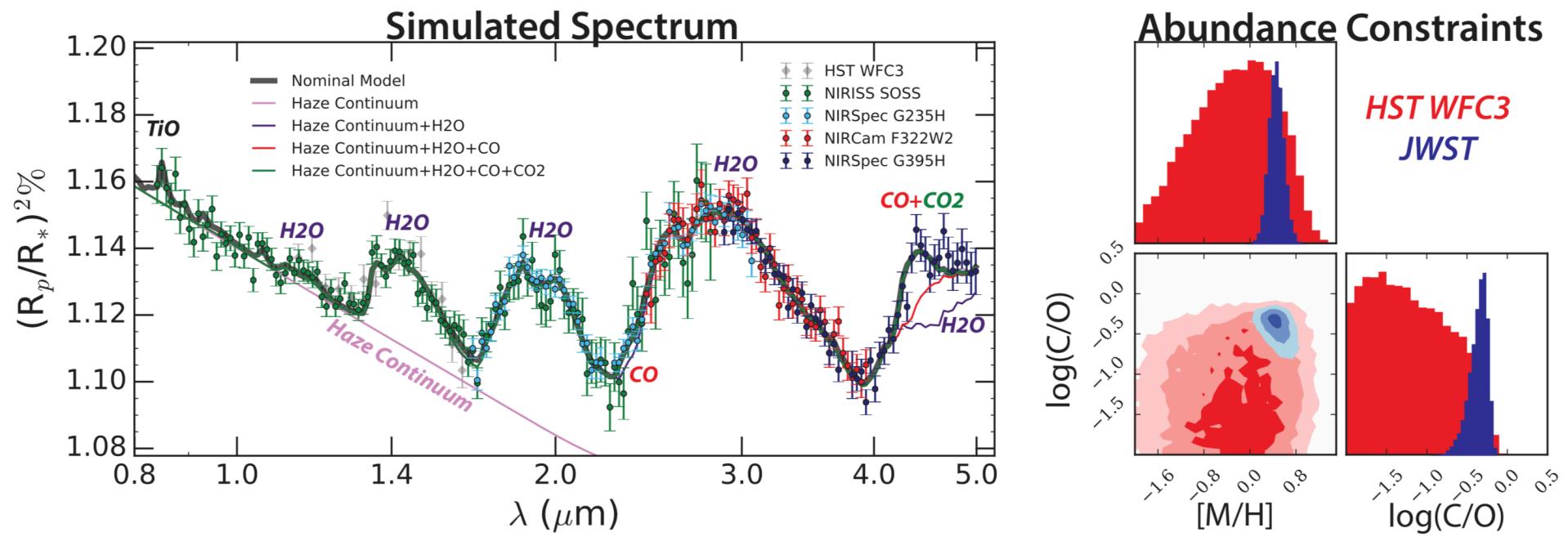


NIRISS/SOSS + NIRSpec/G235H + NIRCam/F322W2 + NIRSpec/G395H

Program 1: “Panchromatic Transmission”

Expected results

- Pandexo simulator (*Batalha et al. 2017*)
- Model transmission spectrum from the CHIMERA code (*Line et al. 2014; Line & Parmentier 2016; Batalha & Line 2017*)
- Based on HST WFC3 measurements (*K. Showalter et al. in prep*)

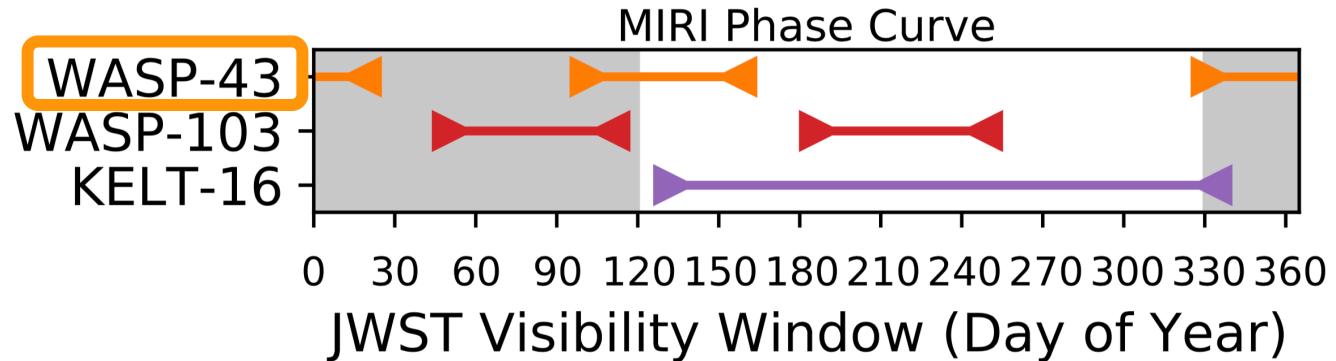


Program 2: “MIRI Phase Curve”

- Chairs: Laura Kreidberg, Nicolas Crouzet, Julianne Moses (48 members)
- Goals:
 - Test the **stability of JWST and MIRI/LRS over day-long timescales**
 - Extract **spectroscopic phase curves** of a hot Jupiter in the 5 – 12 μm range
 - Determine its **3-D atmospheric composition, temperature structure, heat transport, chemistry, cloud properties**
- Observations: **Full-orbit phase curve** (including 2 eclipses and 1 transit) of a short period hot Jupiter with **MIRI/LRS**

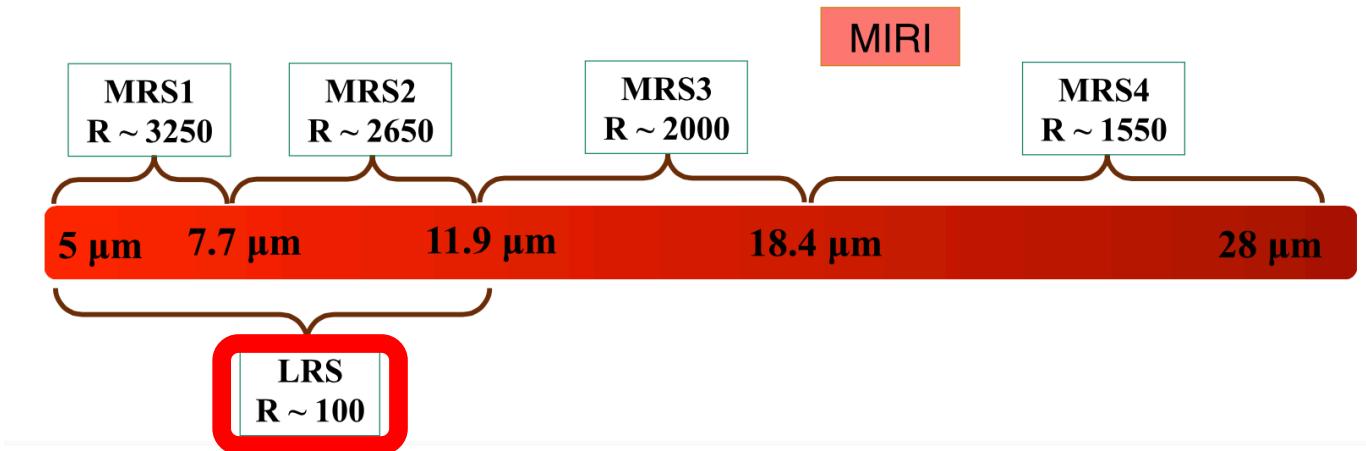
Program 2: “MIRI Phase Curve”

Which target?



WASP-43b: - A highly irradiated hot Jupiter ($T_{eq} \sim 2100$ K, $\delta_e \sim 0.5\%$)
- Spectroscopic phase curves obtained with HST WFC3
(Stevenson et al. 2014)

Program 2: “MIRI Phase Curve”



MIRI:

- Only instrument at $\lambda > 5 \mu m$
- Si:As detectors, different systematics

LRS:

- Only MIRI mode available for time-series observations

Program 2: “MIRI Phase Curve”

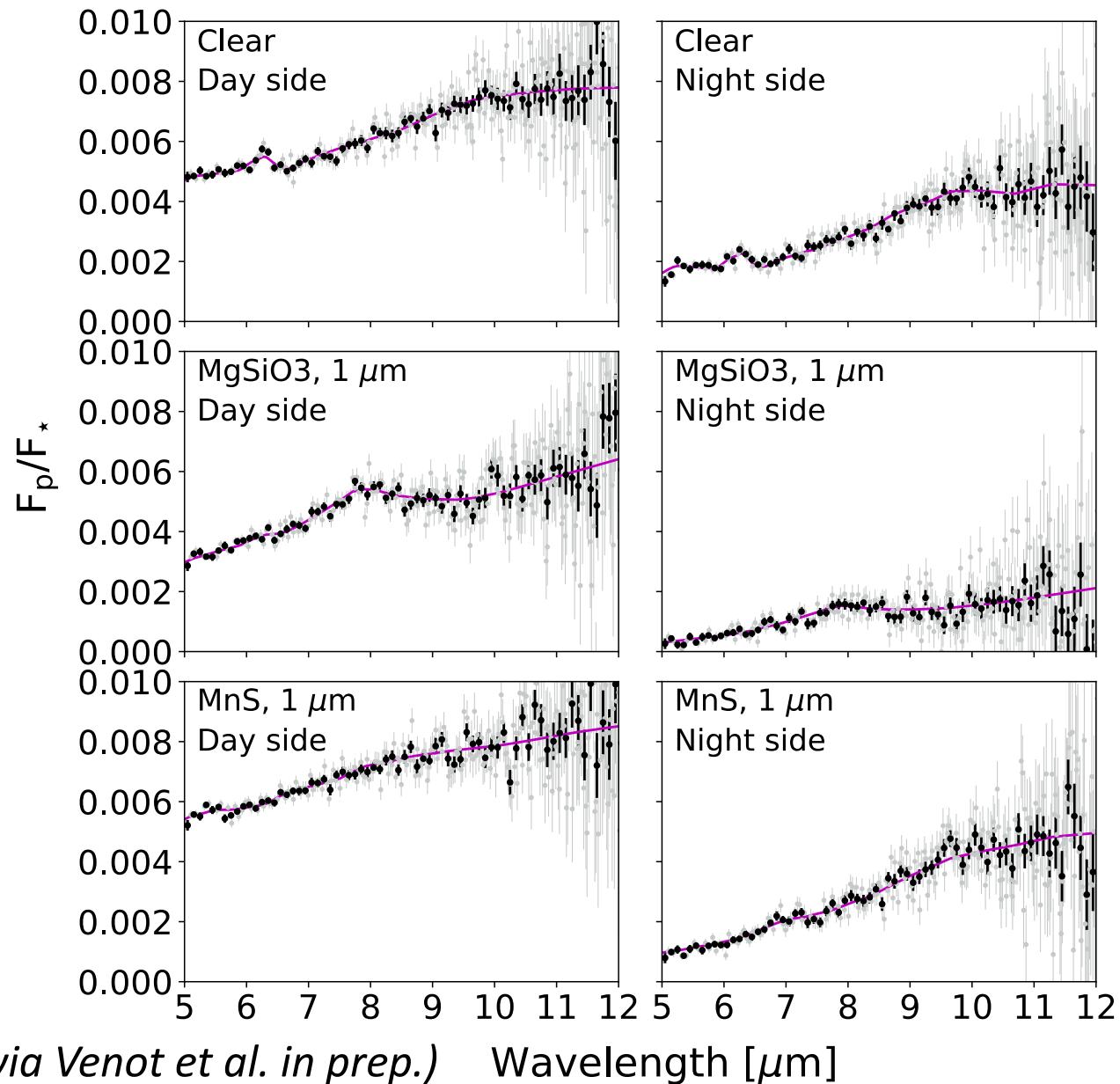
Simulations

System: WASP-43

GCM models
(*Vivien Parmentier*)

Clear or cloudy
atmospheres
(Al_2O_3 , CaTiO_3 , Cr, Fe,
 MgSiO_3 , MnS)

Pandexo simulations,
MIRI LRS



Program 2: “MIRI Phase Curve”

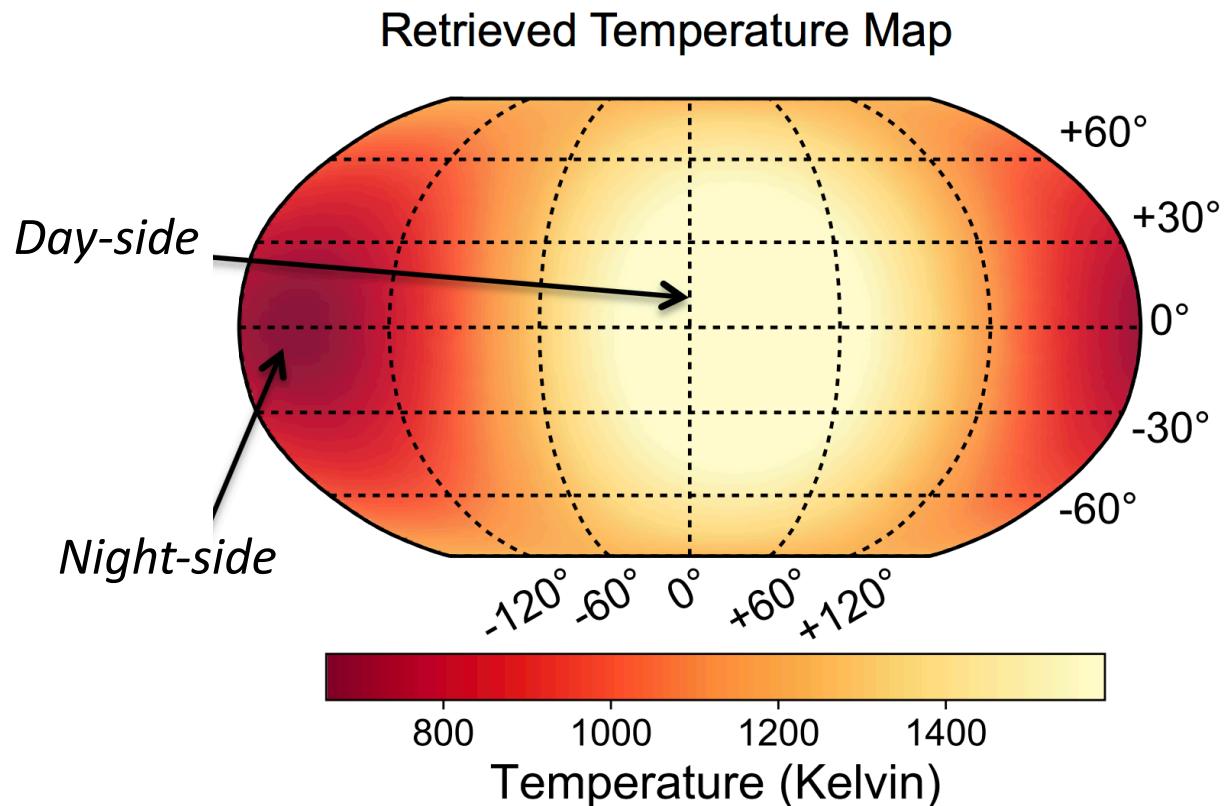
Simulations

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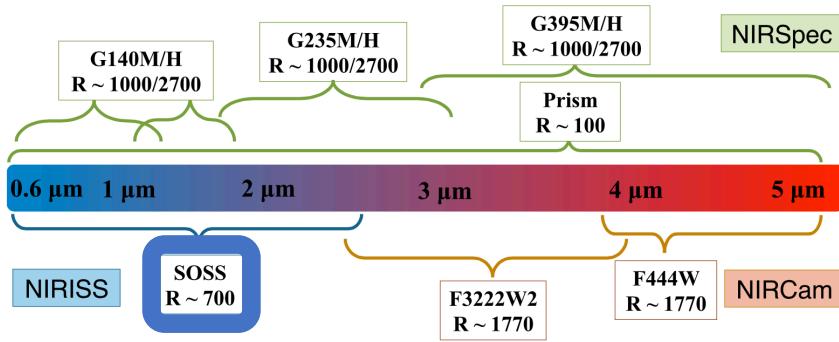
Pandexo simulations,
MIRI LRS



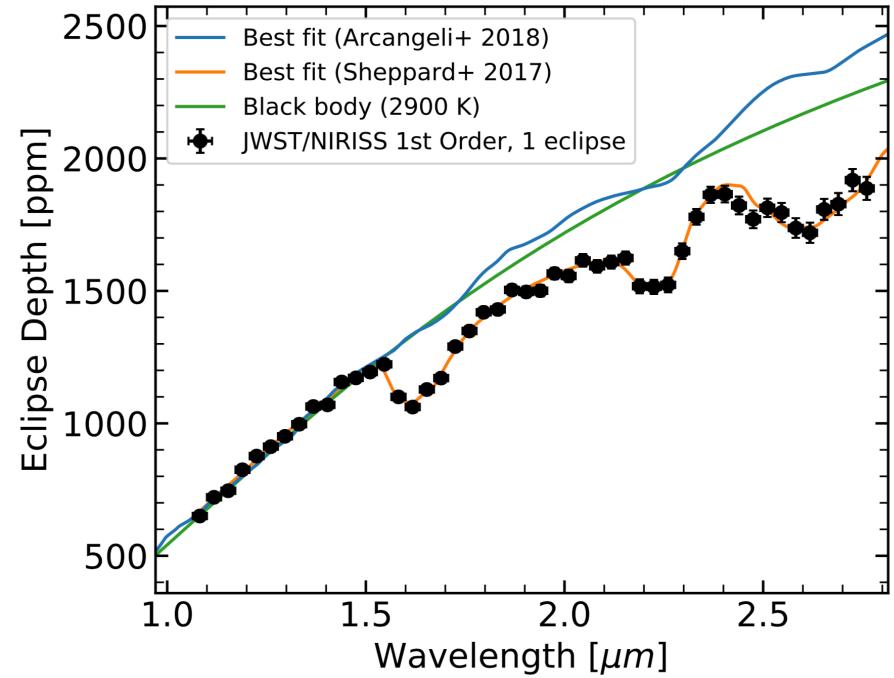
Program 3: “Bright star secondary eclipse”

- Chairs: Björn Benneke, Jacob Bean, Eliza Kempton (14 members)
- Goals:
 - Test JWST’s behavior at the **limit** of its achievable precision
 - Prepare for follow-up of **TESS** exoplanets
 - Obtain the **emission spectrum** of a highly-irradiated hot Jupiter in the NIR, measure its **energy budget** and **thermal structure**
- Observations: **One secondary eclipse** of a hot Jupiter orbiting a bright star with **NIRISS/SOSS**

Program 3: “Bright star secondary eclipse”



A single eclipse of the hot Jupiter
WASP-18b with NIRISS SOSS
(Single Object Slitless Spectroscopy,
0.85 – 2.8 μm)

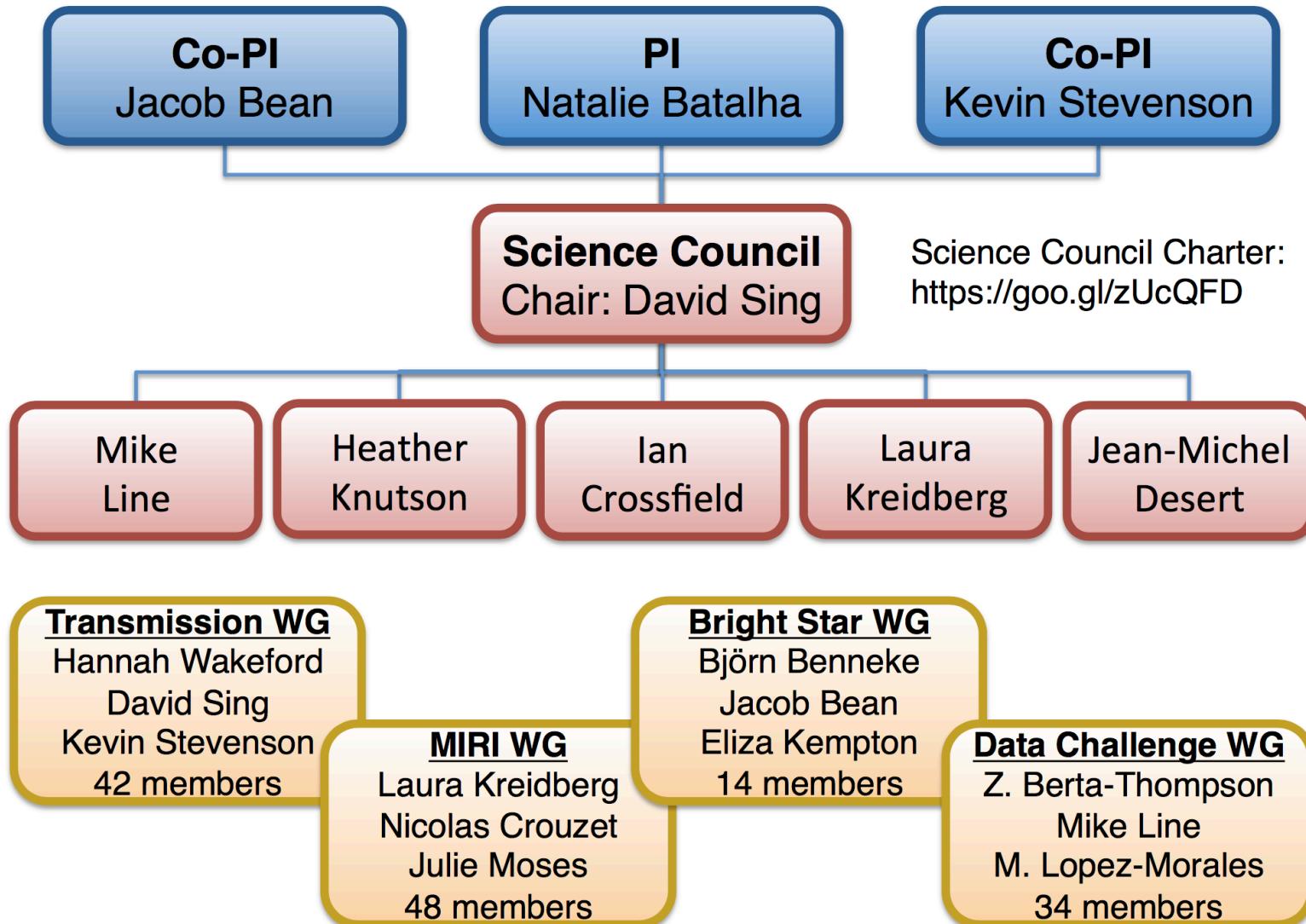


Program 4: “Data challenge”

Chairs: Zach Berta-Thomson, Mike Line, Mercedes Lopez-Morales
(34 members)

- **Prepare tools** to enable quick and accurate analysis and interpretation of the data, make them available to the community.
- Assess the achieved precision, **identify the major systematic noise** sources, their quantitative impact, and potential avenues for mitigating them.
- Internally **validate the scientific conclusions** drawn from JWST data through comparison of results by different team members, and **determine best practices** and required ingredients for JWST analyses.
- **Inform** the planning and selection of future JWST exoplanet programs.
- Deliverables: Data Analysis Toolkits, Instrumental Performance Reports, Field guides

Management Team



Allocated time

Full program: **78.1 hours** allocated

- Panchromatic Transmission: 39.6 hours
- MIRI Phase Curve: 29.4 hours
- Bright star secondary eclipse: 9.1 hours

Requested: 78.1 hours

Science time: 52.1 hours

JWST Exoplanet Tools

 PandExo

Search docs

Getting Started

Pre-installation Data Download

Installation for Users With Conda

Installation for Users With PIP

Troubleshooting-Common Errors

The Importance of Upgrading PandExo

JWST Tutorial

Possible Instrument Input Params

Py Dict Structure of JWST Output

HST Tutorial

The Code

[Docs](#) » PandExo: An Exoplanet ETC

[View page source](#)

PandExo: An Exoplanet ETC

Tools to help the community with planning exoplanet observations.



The online PandExo module has moved to [STScI's ExoCTK](#).

PandExo is both an online tool and a python package for generating instrument simulations of JWST's NIRSpec, NIRCam, NIRISS and NIRCam and HST WFC3

Contents:

- [Getting Started](#)
 - [Should I install PandExo or use the online interface?](#)
 - [Requires](#)
- [Pre-installation Data Download](#)
 - [JWST Reference Data](#)
 - [Stellar SEDs](#)

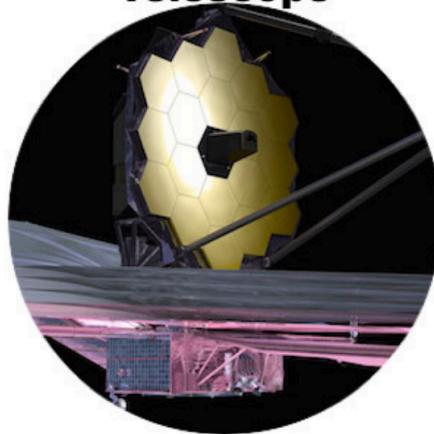
PandExo: The Exoplanet ETC

Tools to help the community with planning exoplanet observations.



On-line version

James Webb Space Telescope



Hubble Space Telescope



JWST Exoplanet Tools

Pandexo:

<https://natashabatalha.github.io/PandExo/index.html>

<https://github.com/natashabatalha/PandExo>

<https://exoctk.stsci.edu/pandexo/>

+ APT (Astronomer's Proposal Tool)

+ ETC (Exposure Time Calculator)

+ Visibility Tool, ESA Sky, JWST User Documentation

Conclusion

- JWST **ERS program**: engage the community to use and evaluate JWST
- JWST will offer **unique capabilities** for transiting exoplanet spectroscopy
- The transiting exoplanet community developed a **collaborative proposal**
- **Four sub-programs**:
“Panchromatic Transmission”, “MIRI Phase Curve”,
“Bright star secondary eclipse”, “Data challenge”
- Proposal accepted, **78.1 hours allocated**
- Data and tools will be **available to the community**
- Everyone is welcome to contribute!!

