

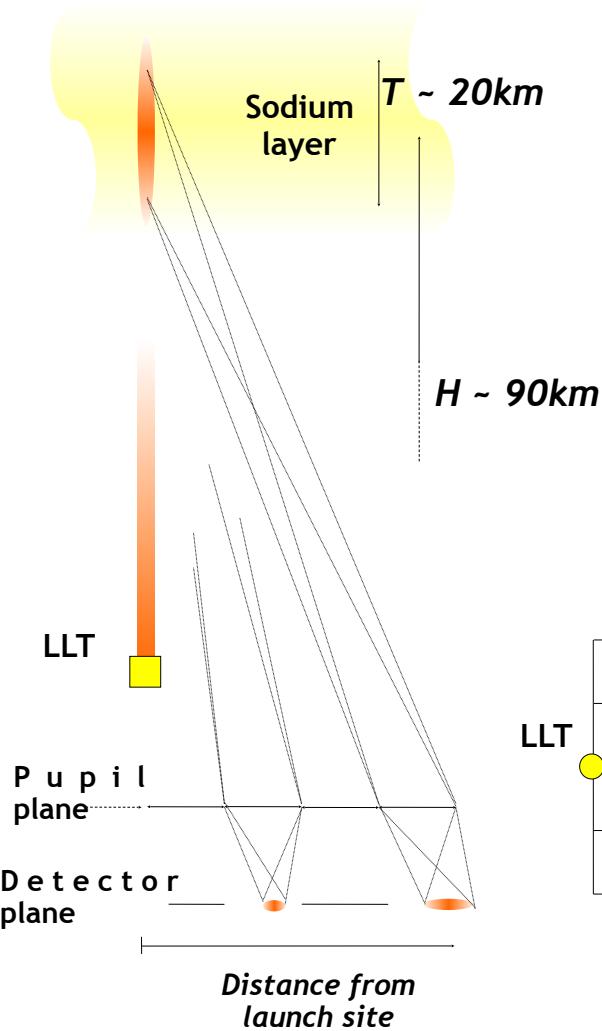
LGS spot truncation mitigation in ELTs: optimizing the pixel usage

L.Blanco¹, C. Correia¹, T. Fusco², B. Neichel¹, Y. Ono¹

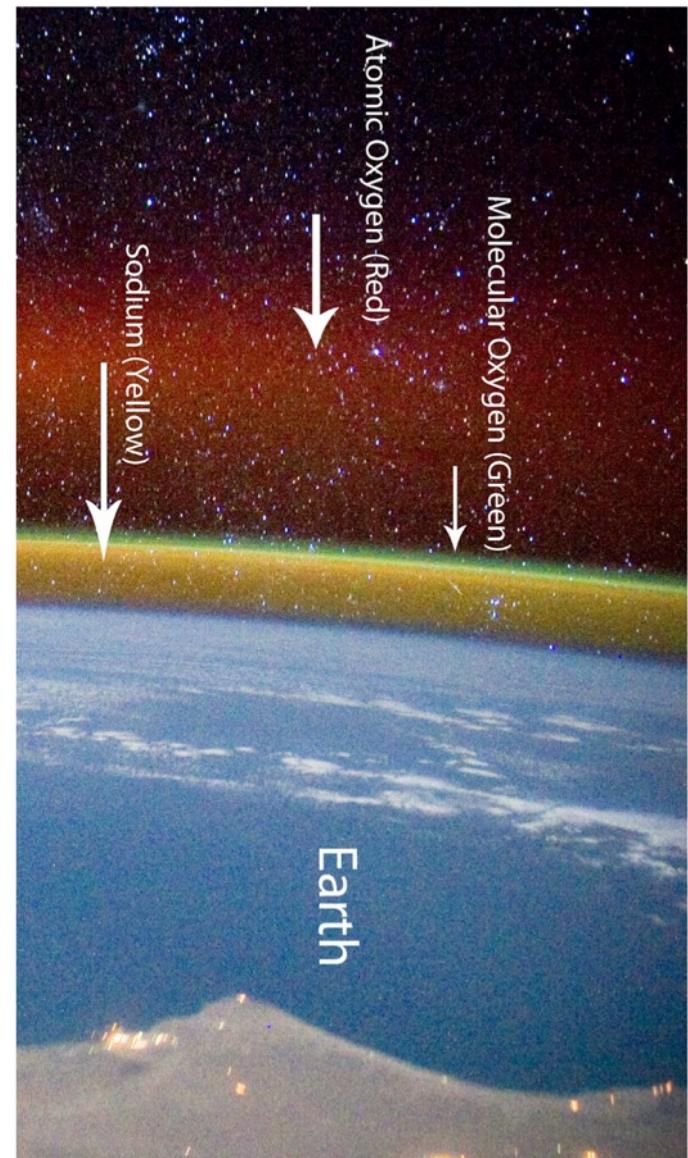
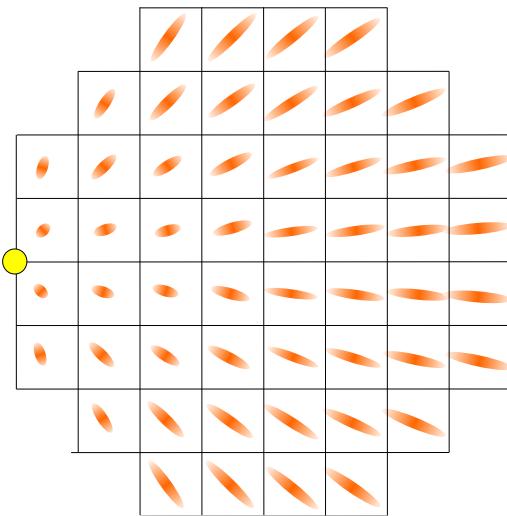
1 Laboratoire d'Astrophysique de Marseille (UMR7326 - CNRS-INSU,
Université d'Aix-Marseille)

2 ONERA - The French Aerospace Lab

LGS Spot elongation

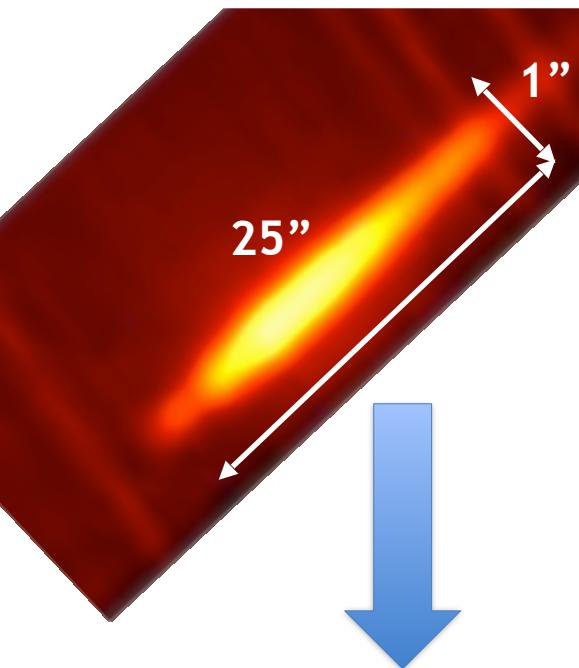


Predicted spot elongation pattern

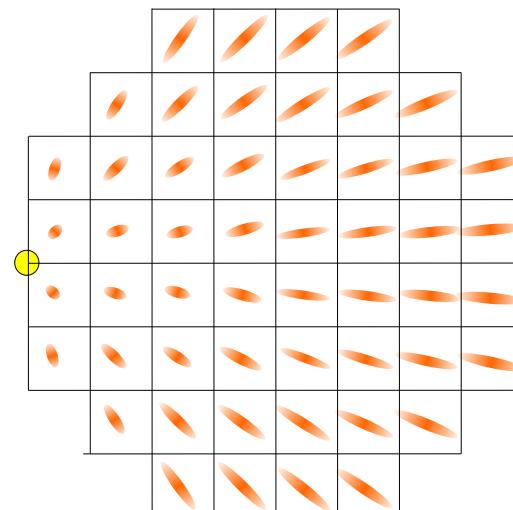


LGS Spot elongation

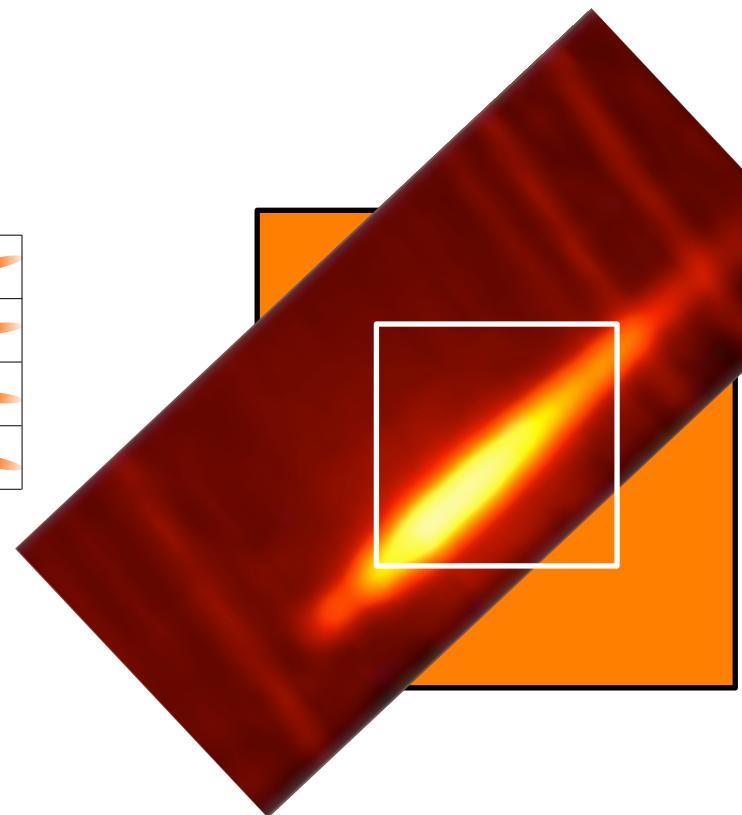
On a ~40m telescope



Ideally, we need subapertures
with 25x25 pixels of ~1"



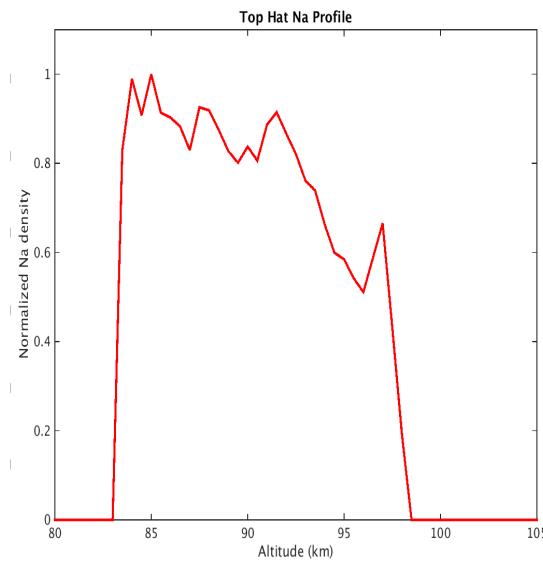
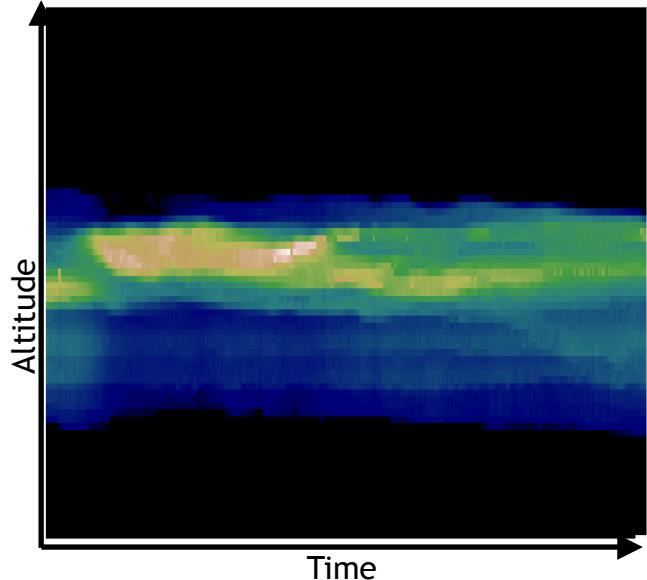
Most likely, we will have no more
than 10x10 pixels



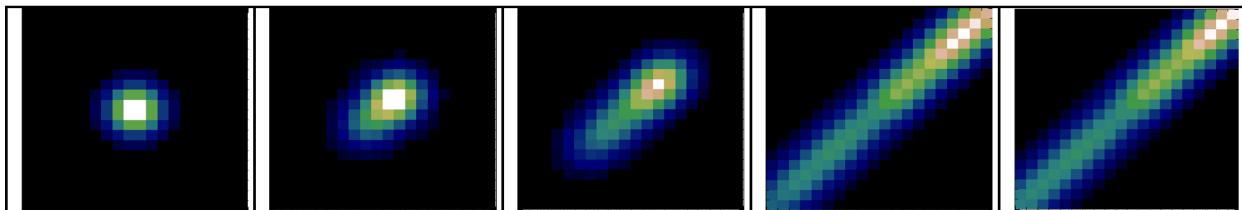
For 80x80 subapertures, we
need 2000 x 2000 pixels

Strong truncation

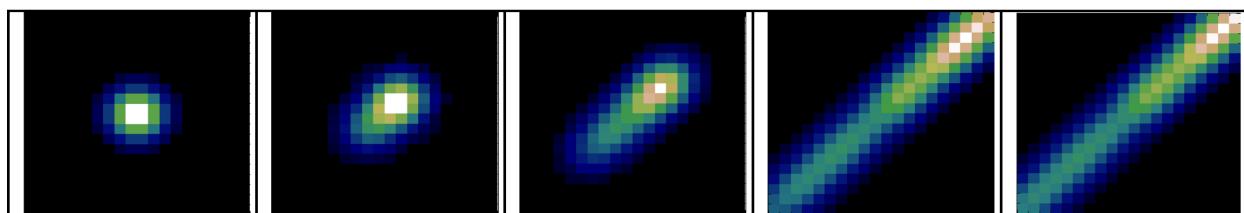
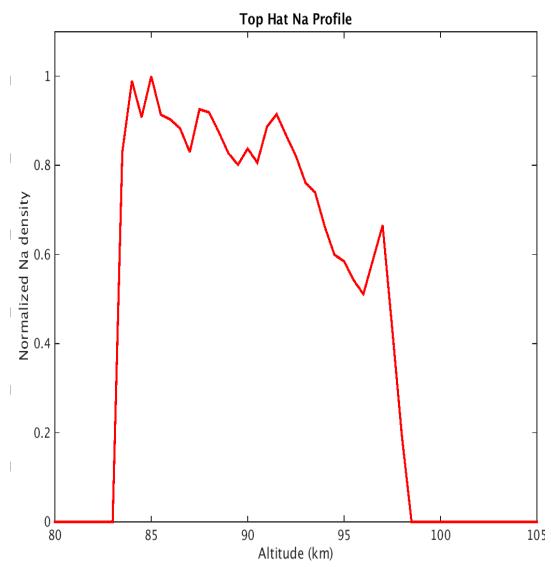
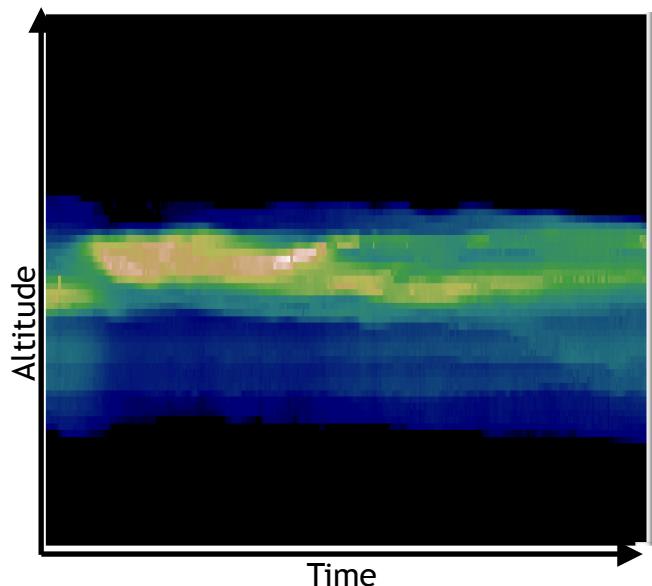
Impact of truncation on a single WFS



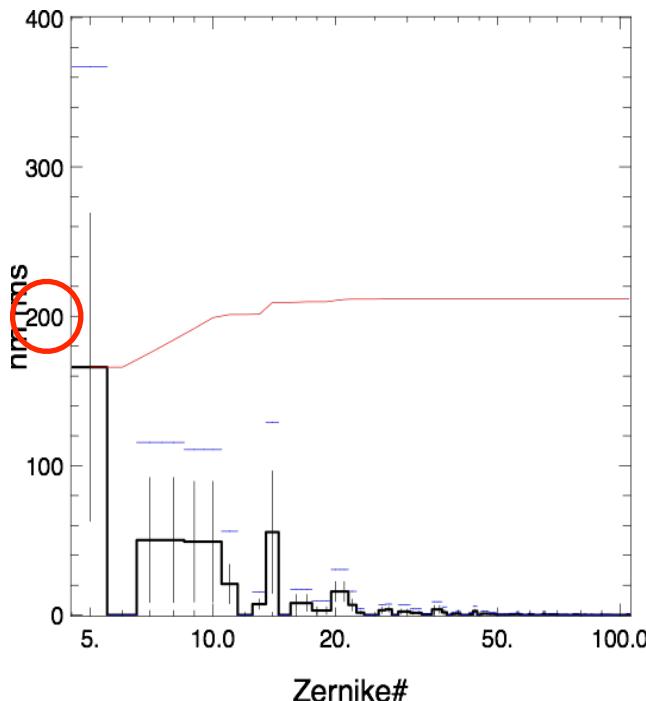
- À la Diolaiti/Schreiber, we simulate elongated spots (normalized by the Na profile intensity)
- Compute the centroid of each sub-aperture, and project the resulting wave-front on Zernikes
- No elongation: no aberration
- No truncation: Tip/Tilt/Focus only



Impact of truncation on a single WFS



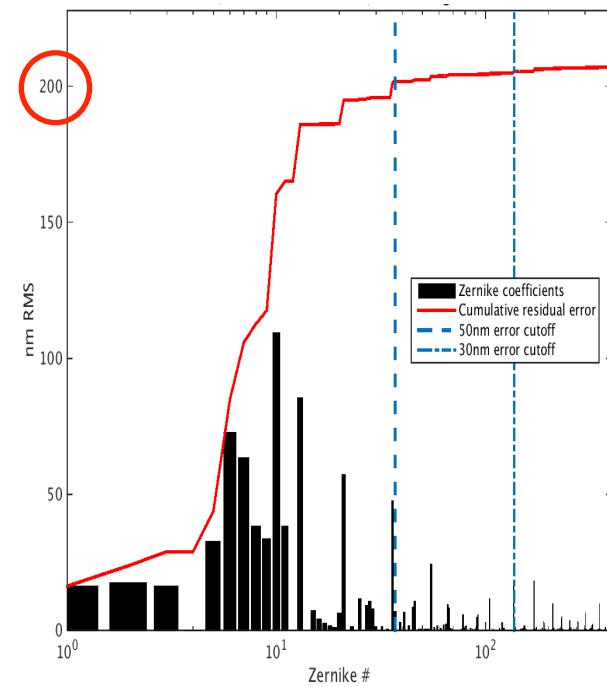
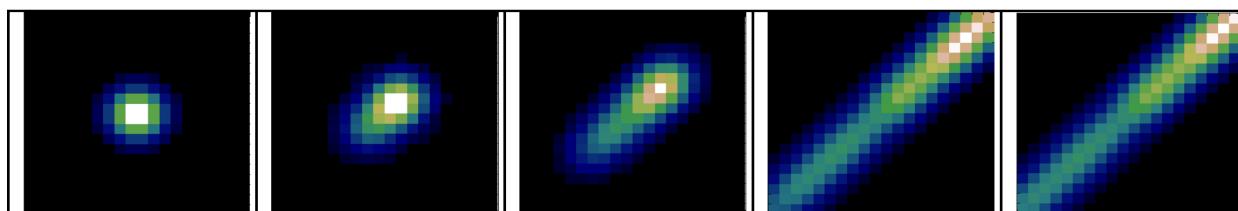
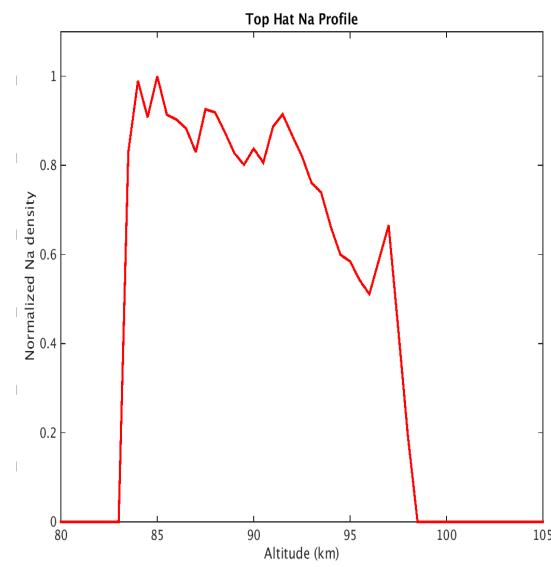
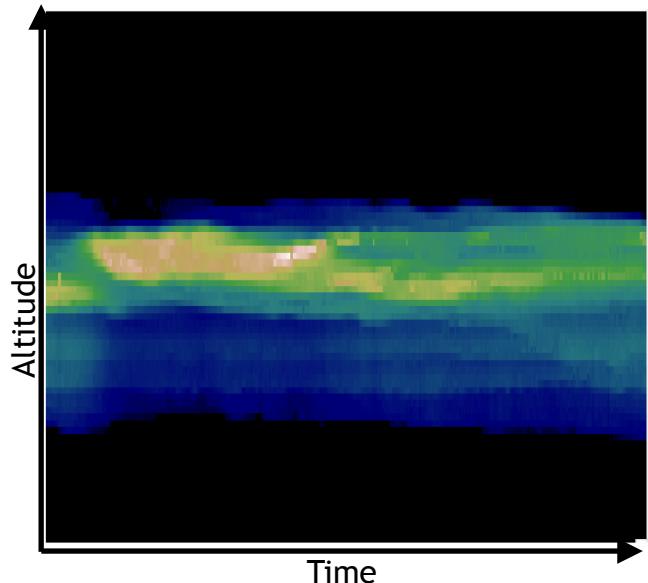
- Select 11 Na profiles showing a large variety of configurations
- For each profile, and each time, compute the residual aberrations (for a single WFS)



(blue = peak; error bar = rms ; histogram = avg ; red = cumulative of avg)

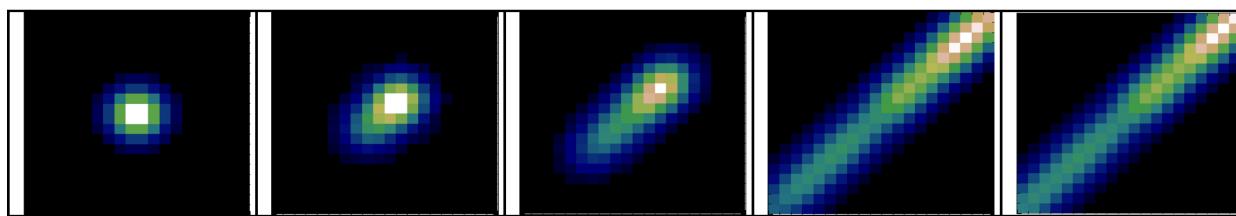
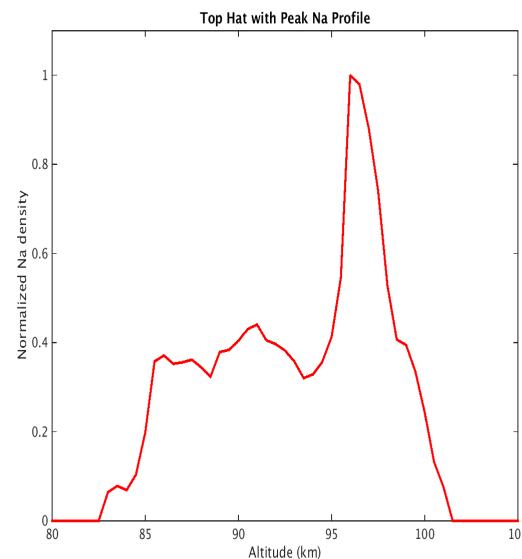
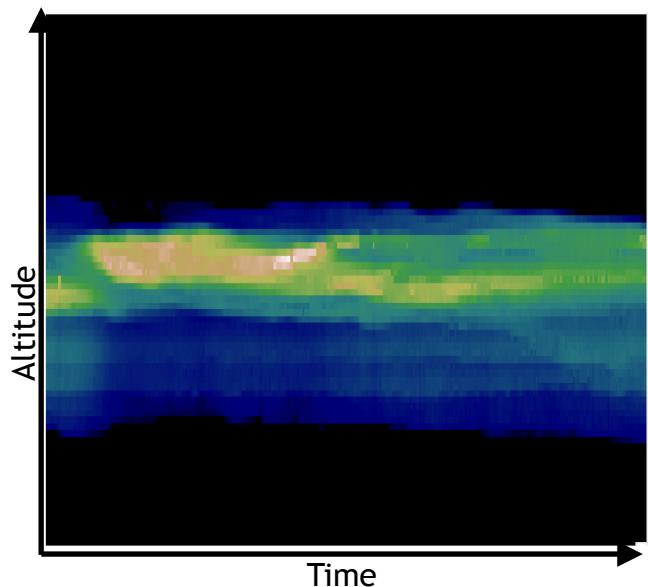
Impact of truncation on LTAO

- 6 LGS
- No turbulence
- MMSE reconstructor
- 42" radius asterism

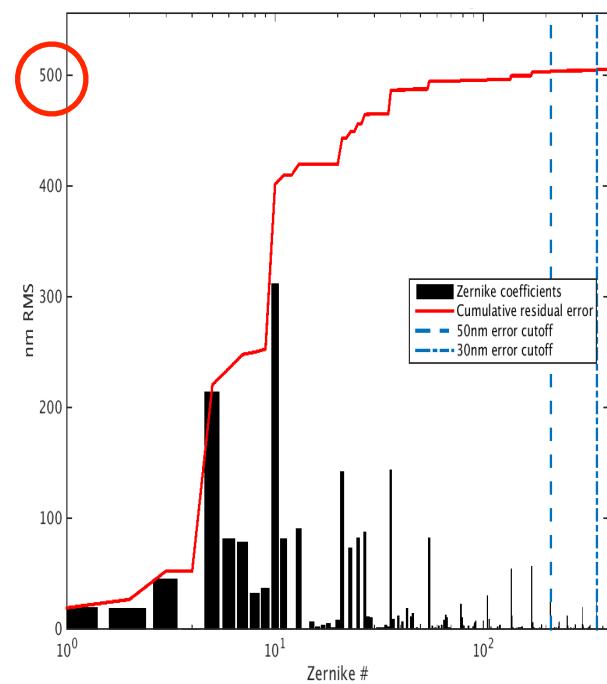


(histogram = avg ; red = cumulative of avg ;
blue = 30nm and 50nm cutoff)

Impact of truncation on LTAO



- 6 LGS
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- 42" radius asterism



(histogram = avg ; red = cumulative of avg ;
blue = 30nm and 50nm cutoff)

Need for a Truth Sensor but aberrations are mostly low order
=> 10x10 subapertures on a truth sensor should be enough on average but on very perturbed profiles...

(more on tomographic errors in T. Fusco's talk this afternoon)

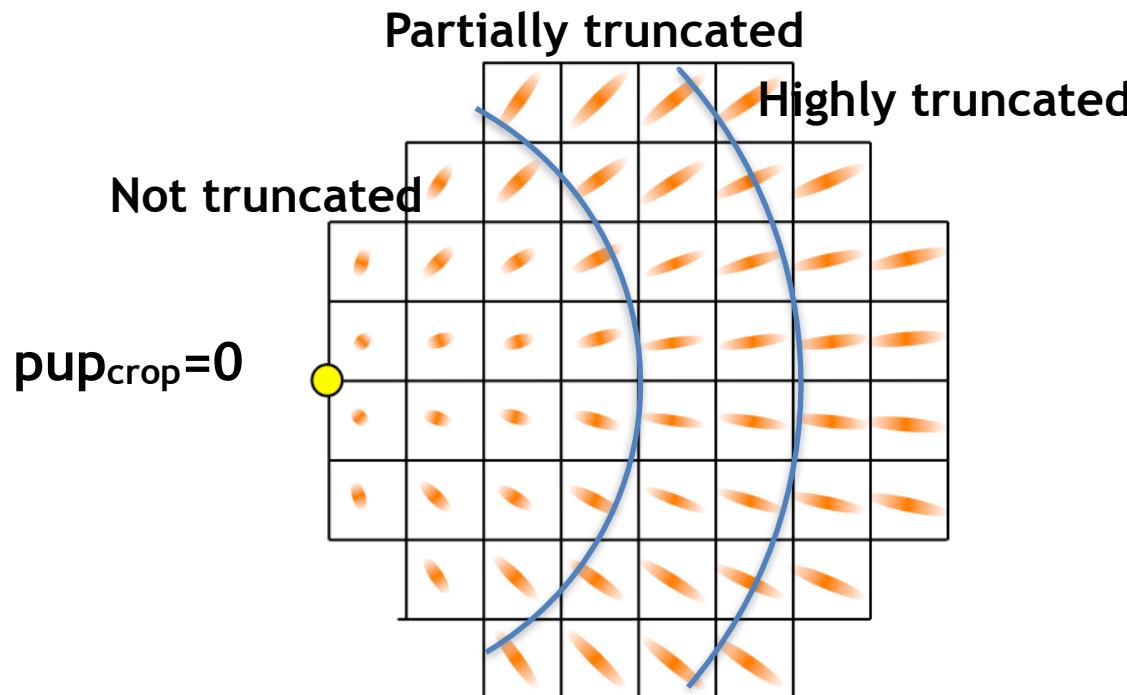
Elongated spots cropping

What if we completely discarded the biased measurements from the subapertures with the most truncated spots ?

Elongated spots cropping

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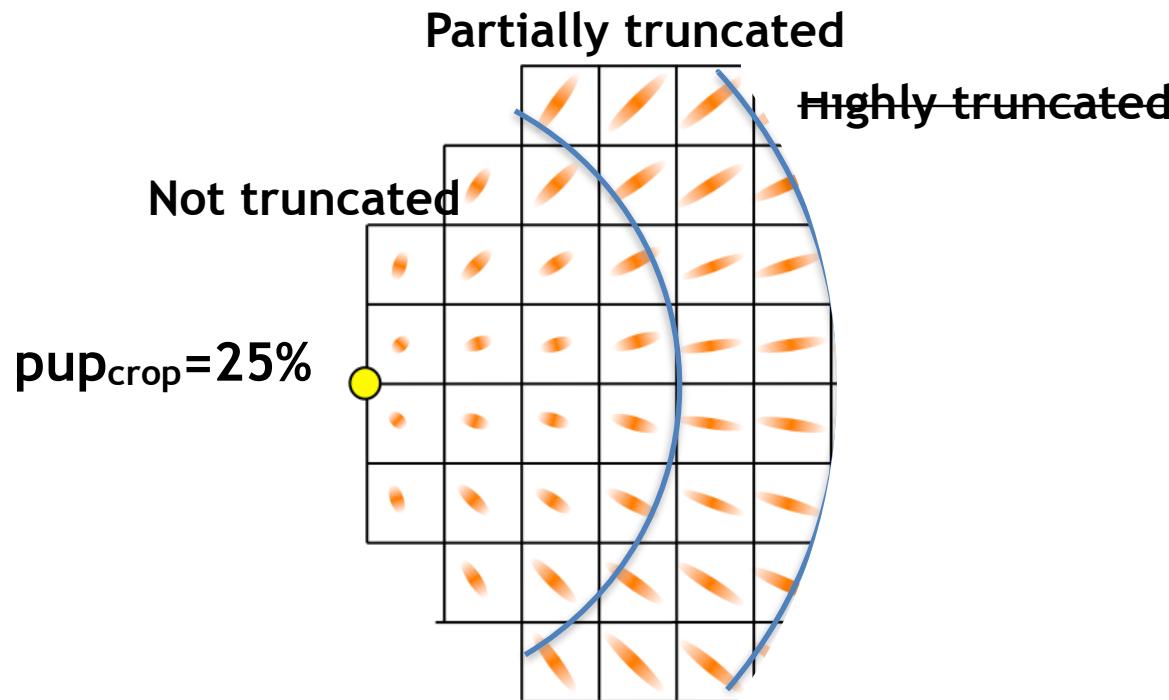
Spot elongation \propto distance from laser launch telescope D_{LLT}
->we ‘crop’ the subapertures at $D_{LLT} > D \times (1-p_{\text{pup}_{\text{crop}}})$



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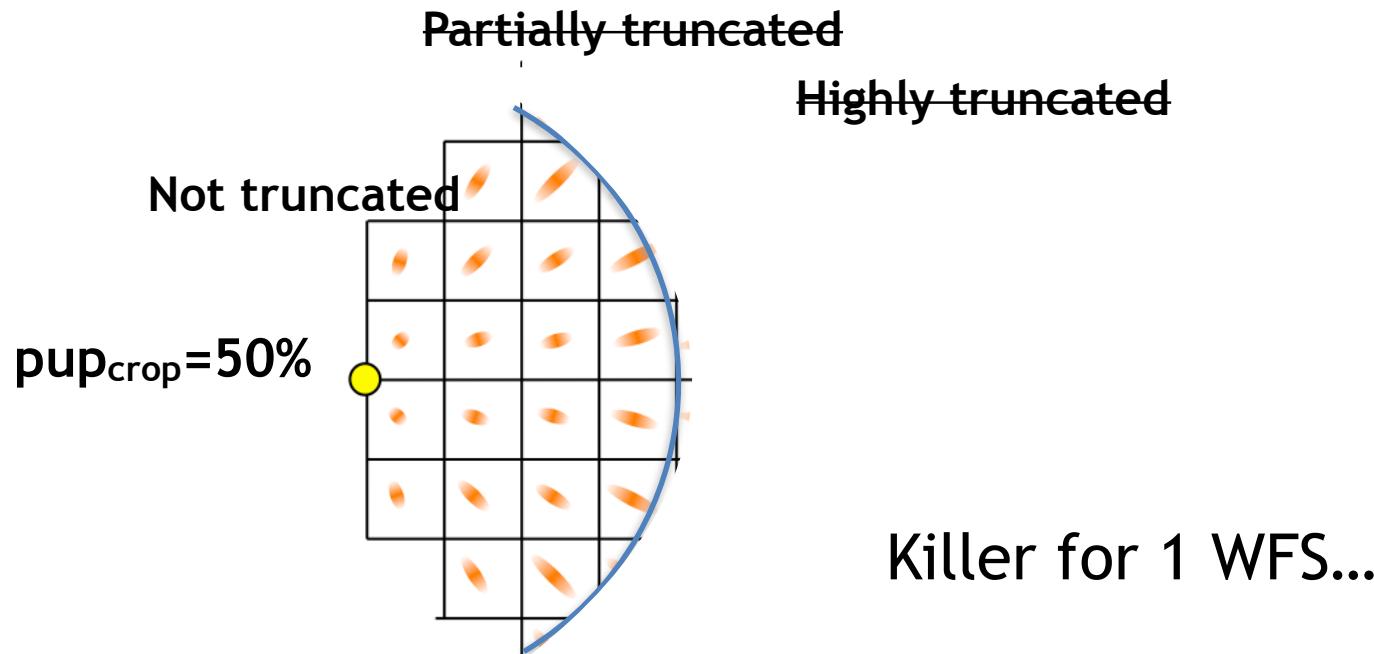
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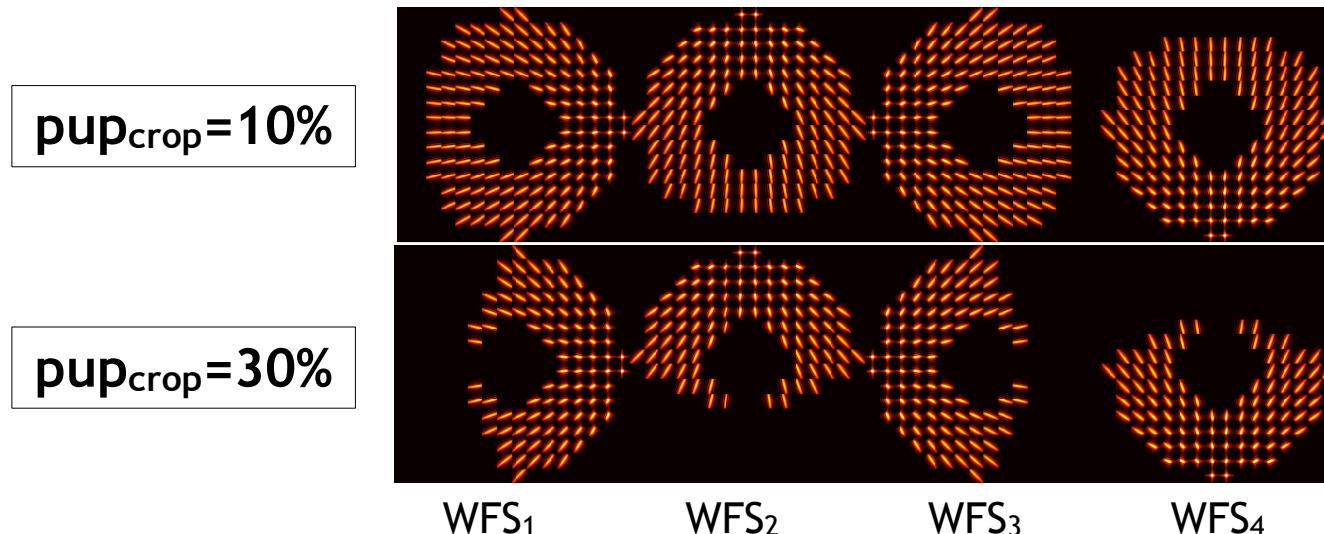
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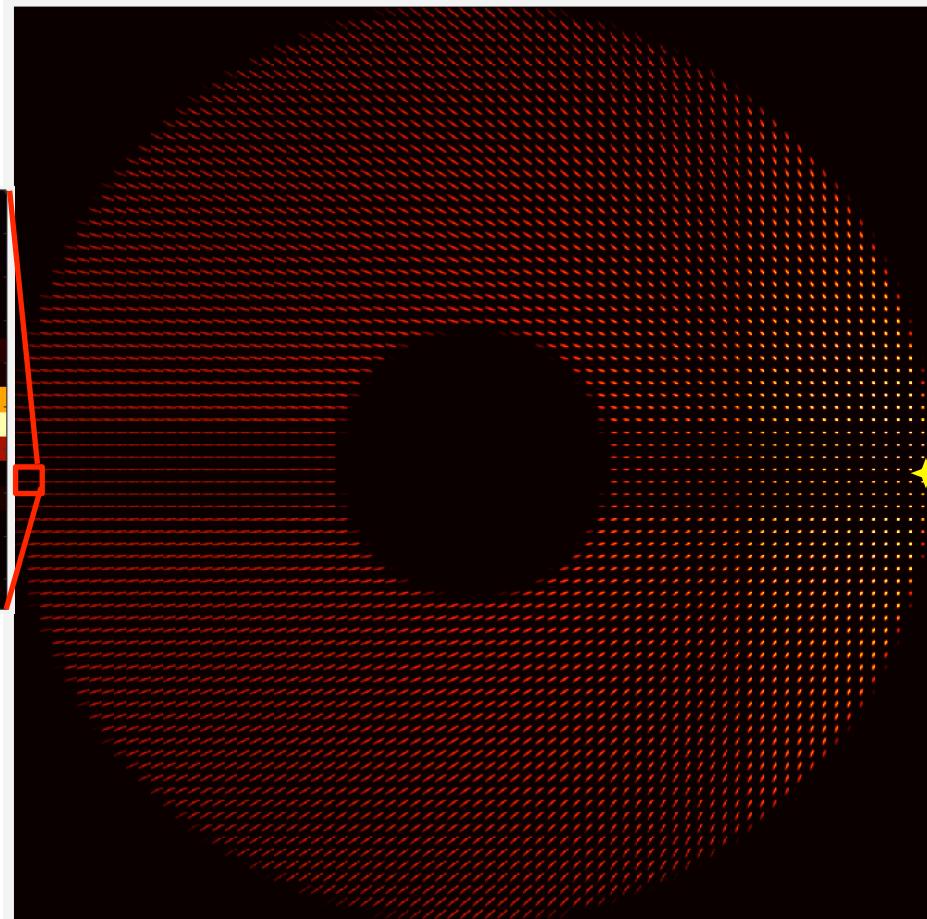
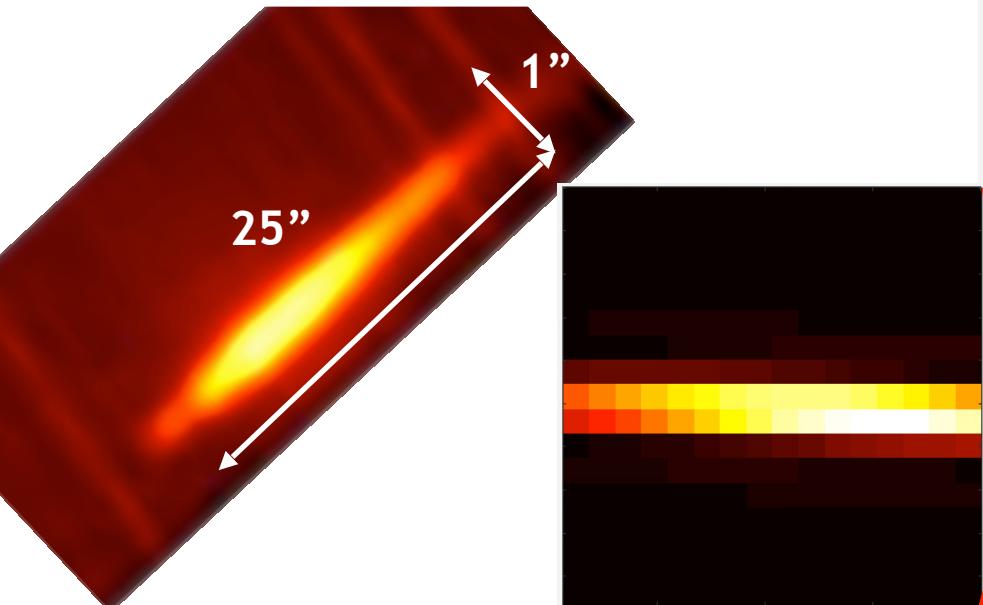
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But we have redundancy



E2E simulations (OOMAO)

On a ~40m telescope



Full E-ELT simulation:

Pupil diameter : 37m

74x74 SH LGS WFS

75x75 actuator DM (Fried
geometry)

Subaperture pitch : 50cm

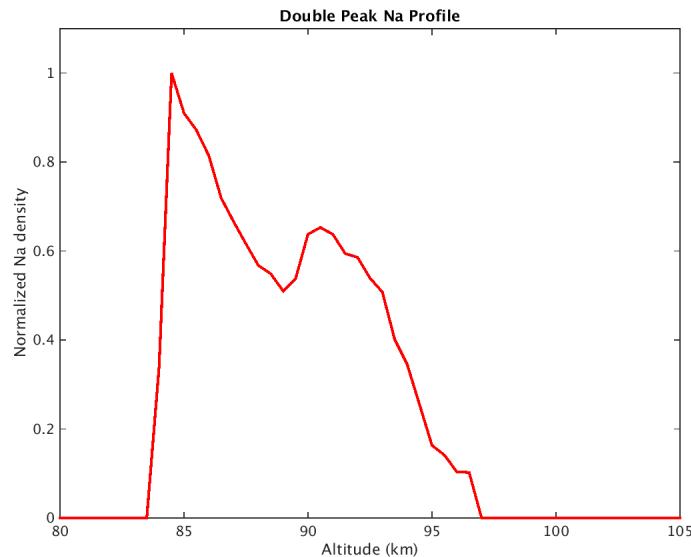
Subaperture FOV : 10'' (1'' pixels)

6 LGS-Asterism diameter : 42''

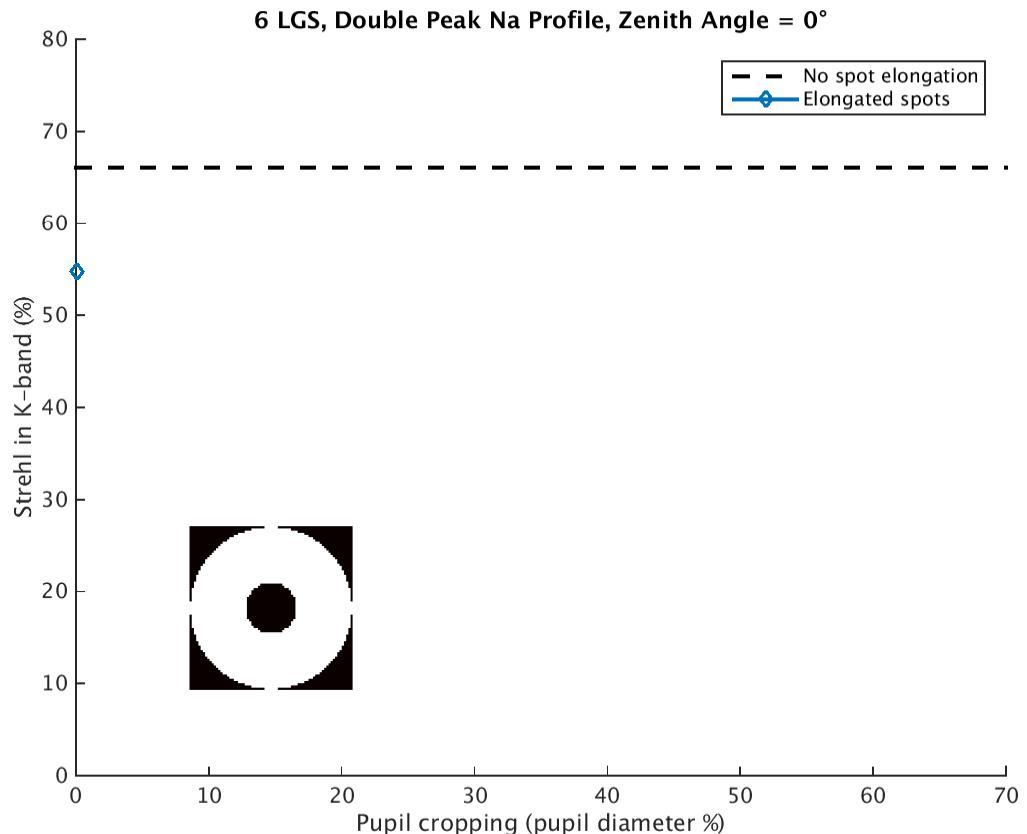
30% central obscuration

(Y. Ono poster P3059, Thursday)

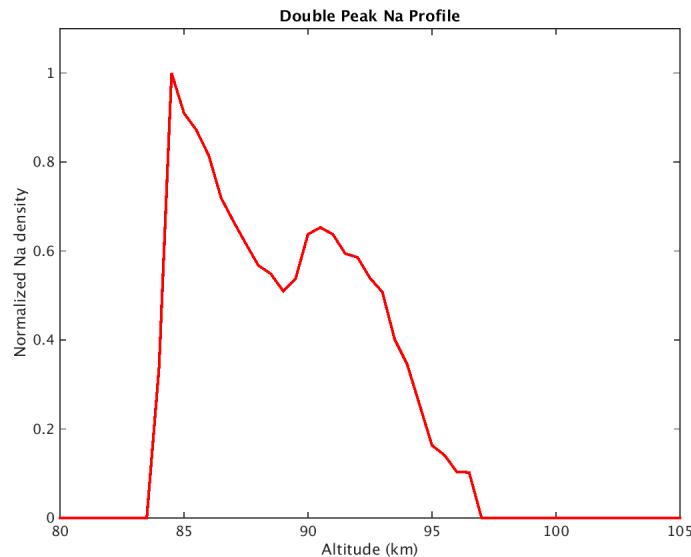
Elongated spots cropping



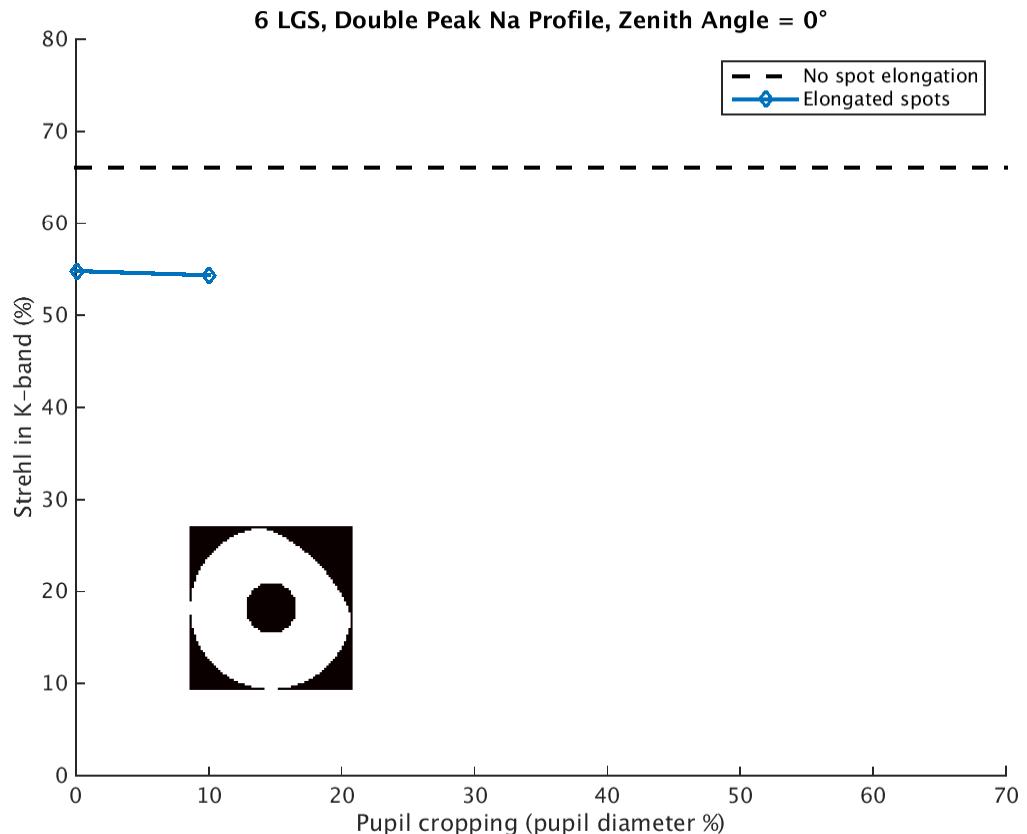
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30% central obscuration
High flux



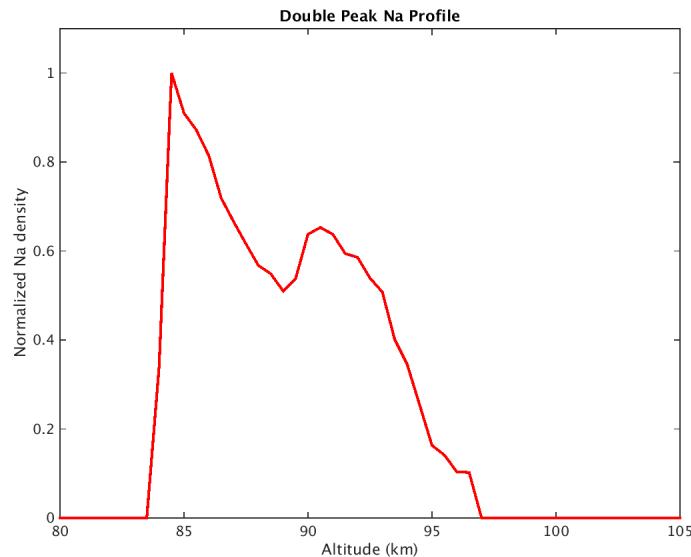
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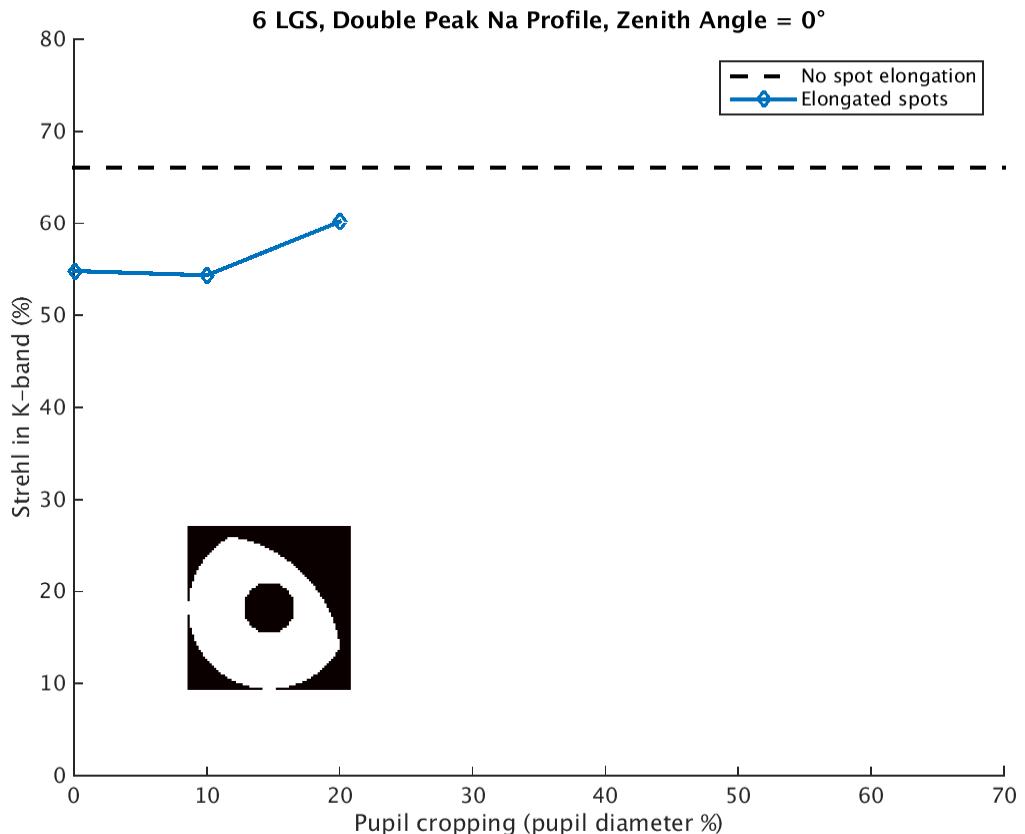
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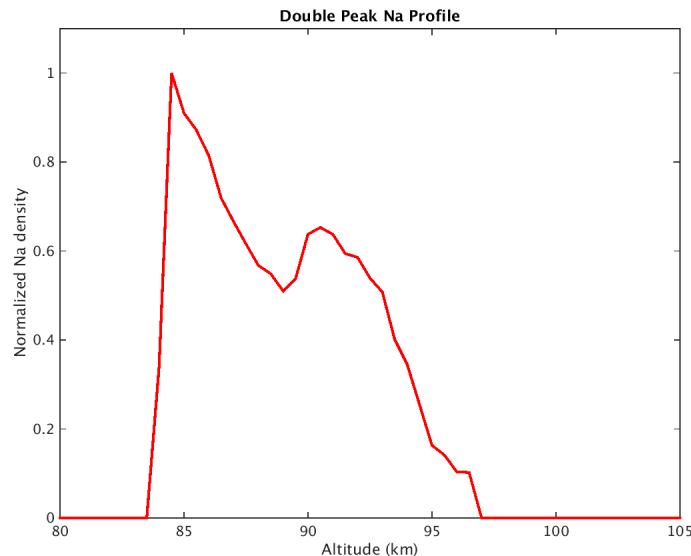
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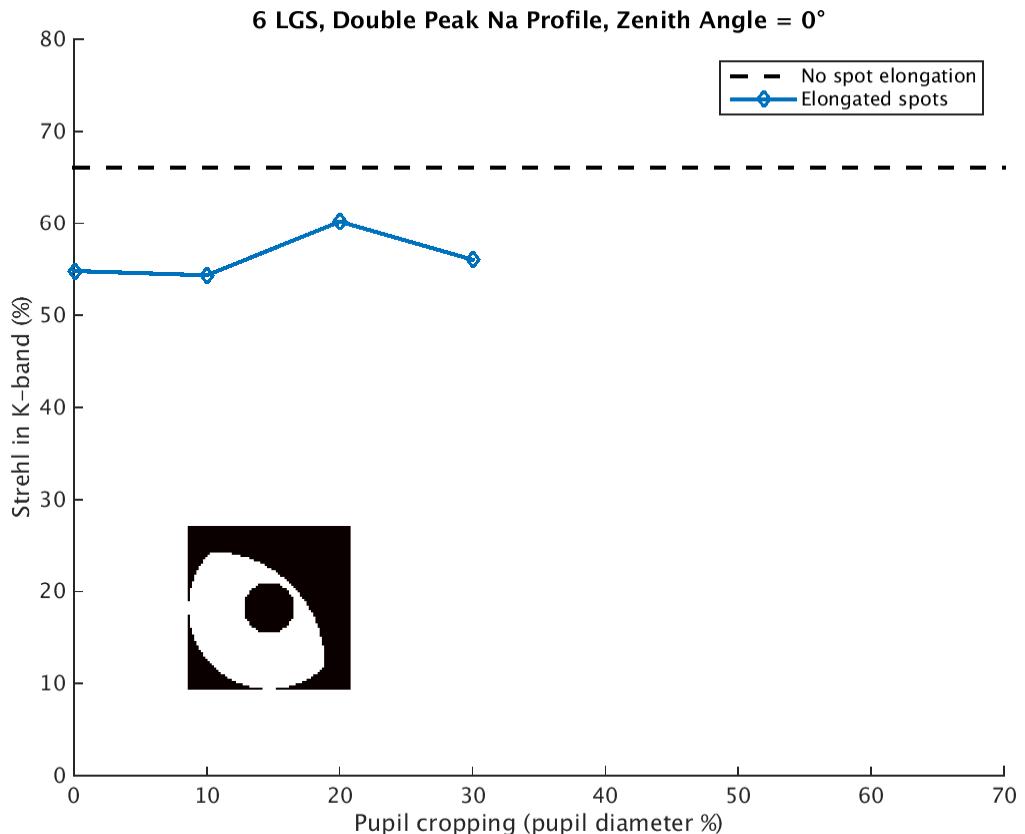
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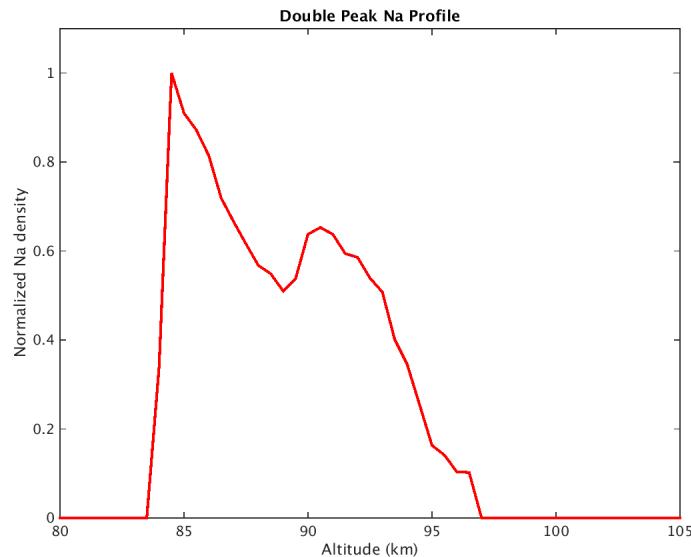
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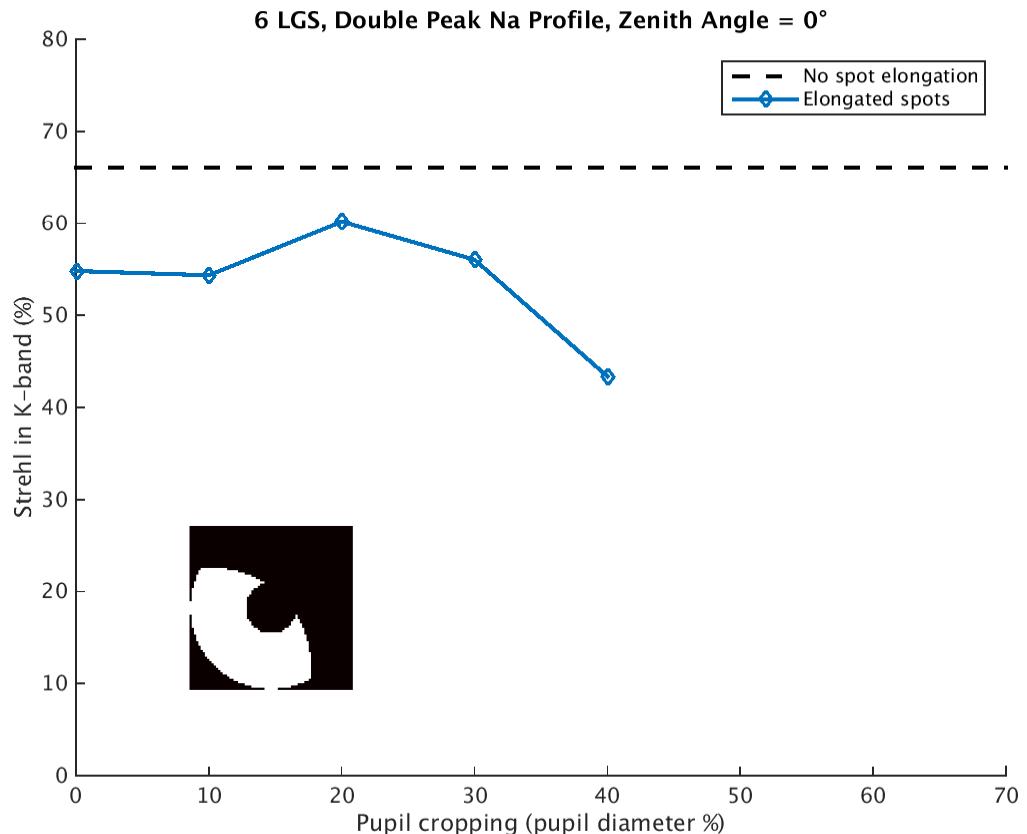
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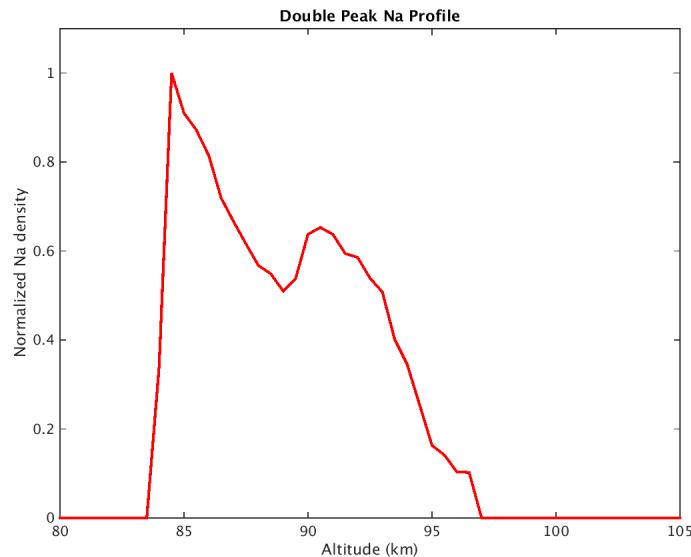
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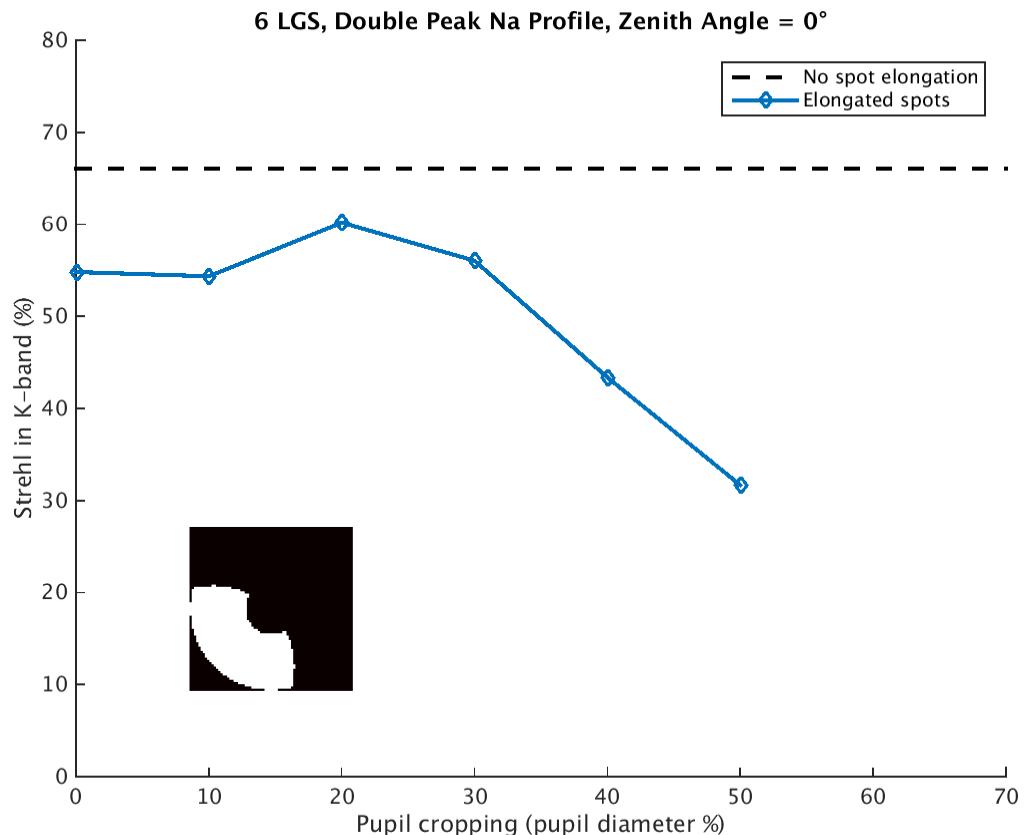
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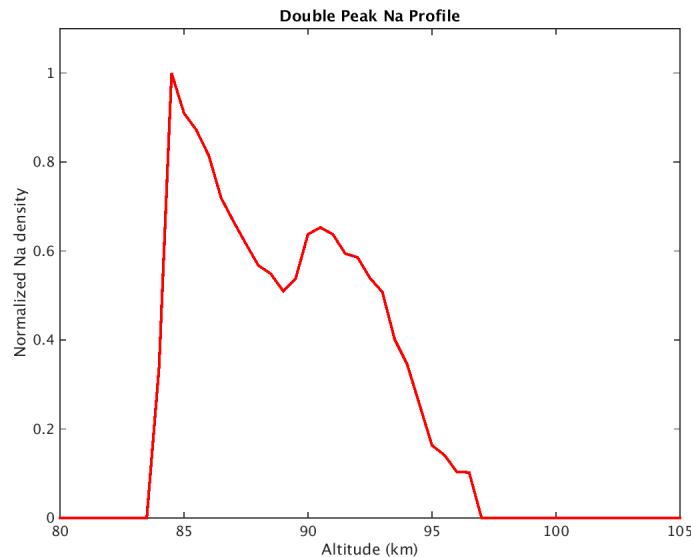
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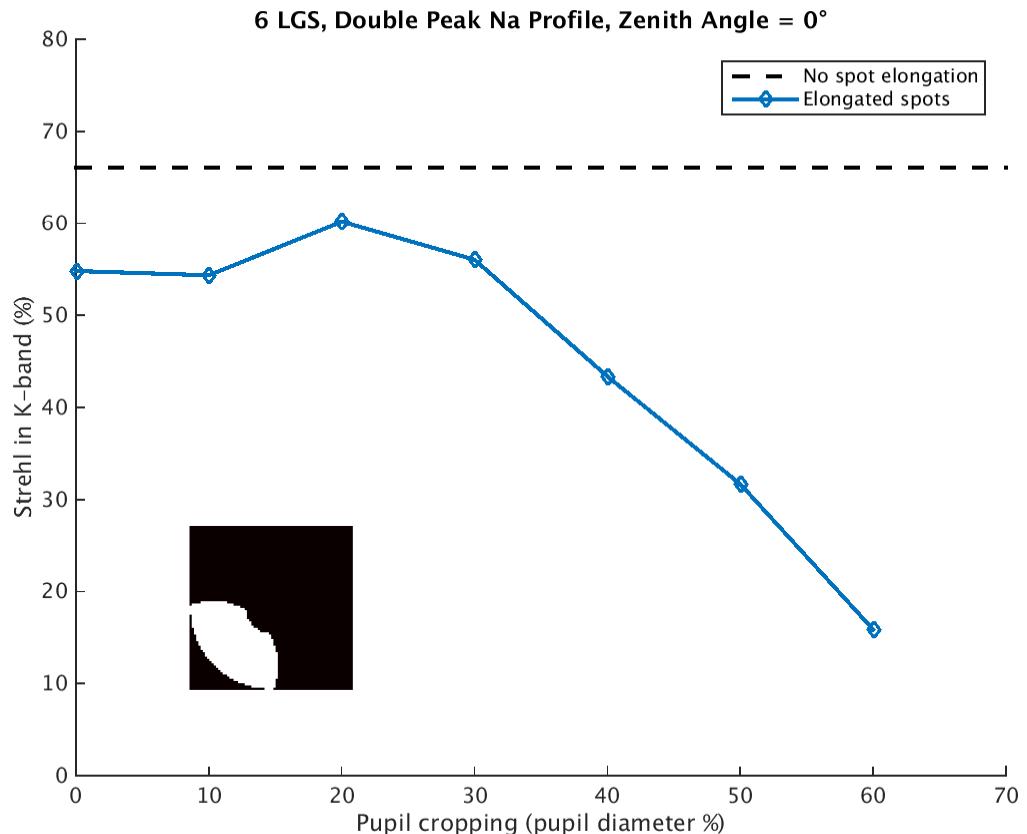
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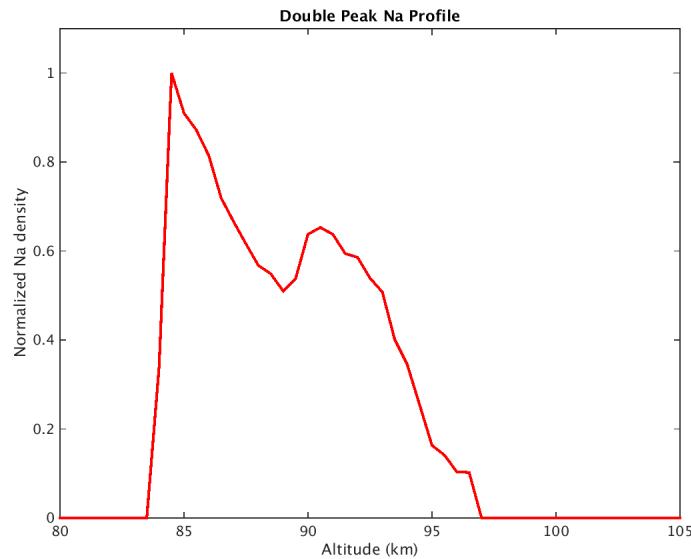
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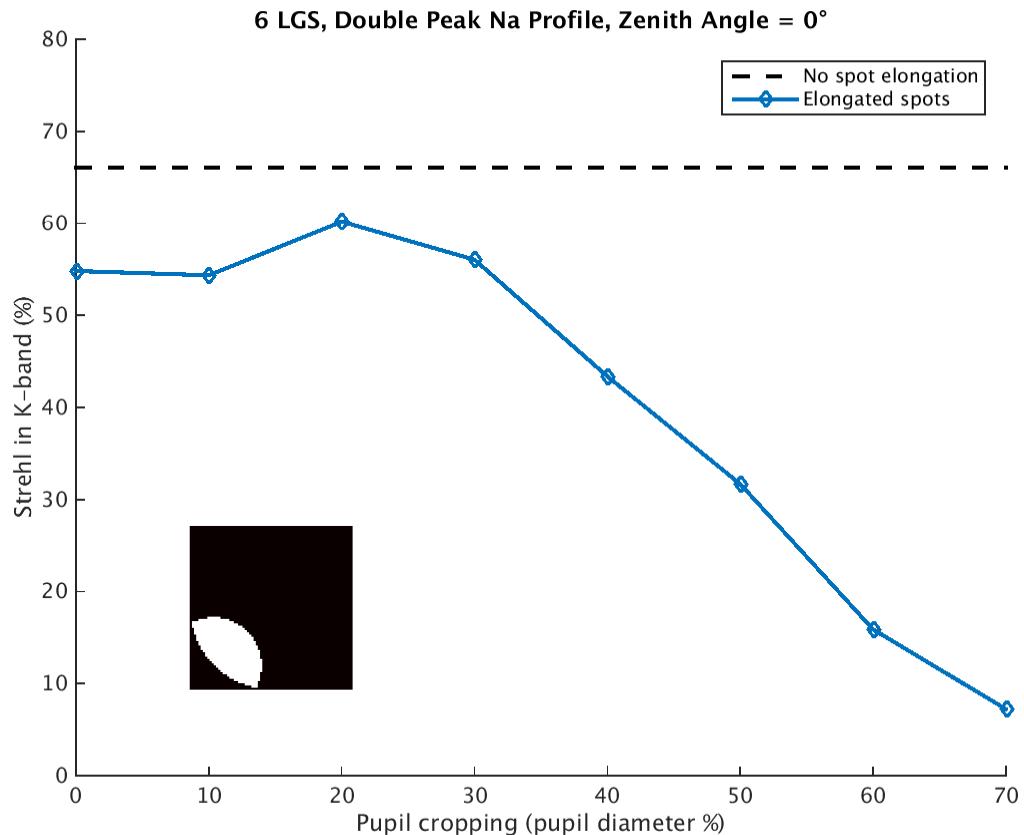
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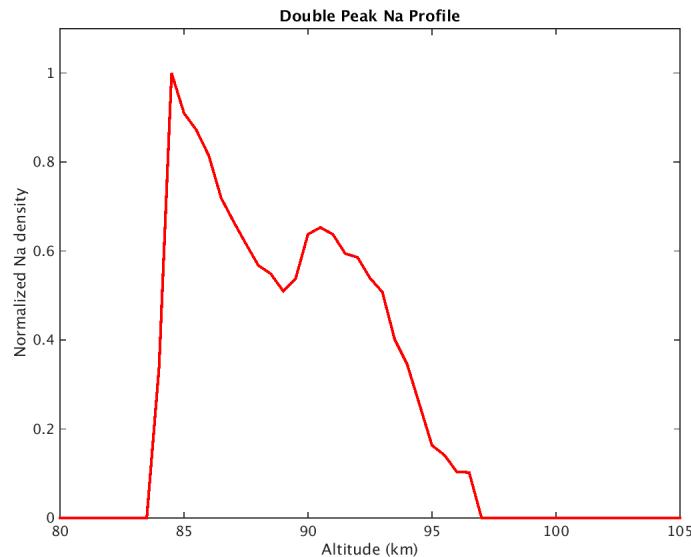
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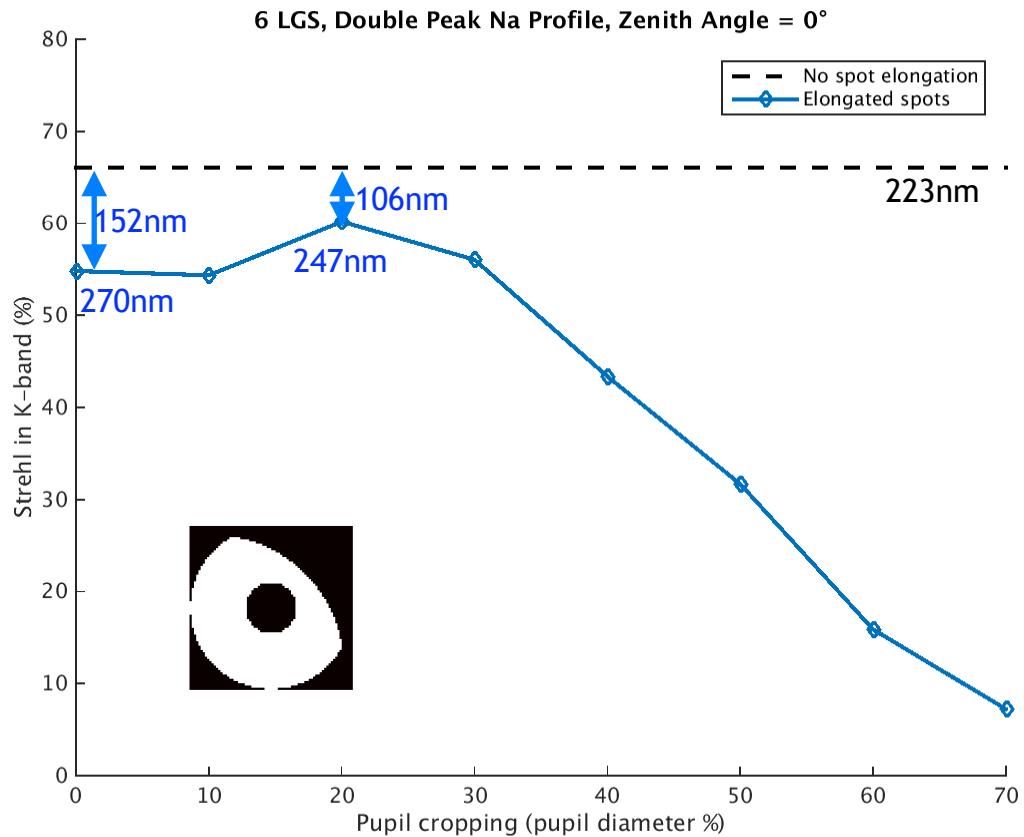
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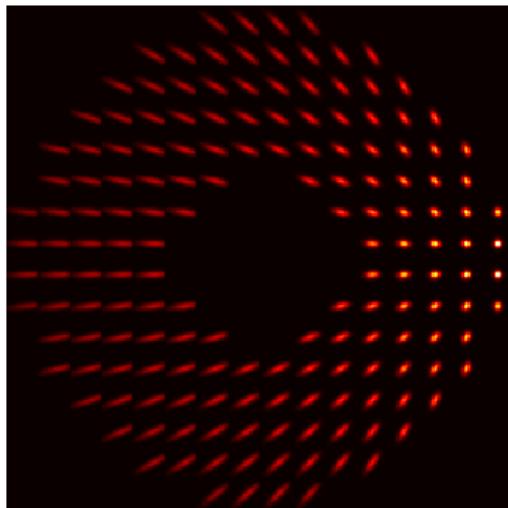


Performance increases as we discard subapertures with too large an elongation/truncation

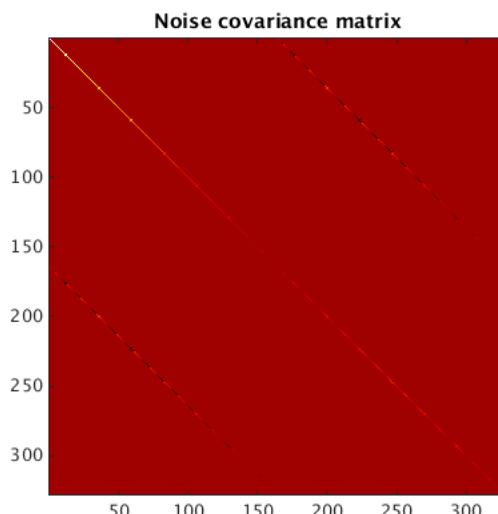
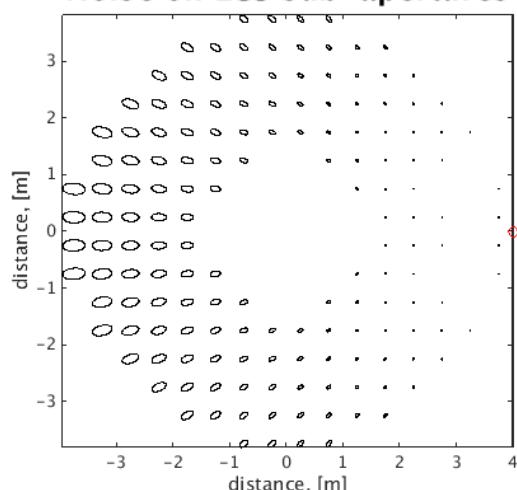
WFE due to elongation reduced from 152nm to 106nm RMS

Differential subaperture weighting

- à la Tallon 2008/Robert 2010
- Takes into account the elongation in the noise covariance matrix
- Takes into account cross-coupling between X and Y slopes

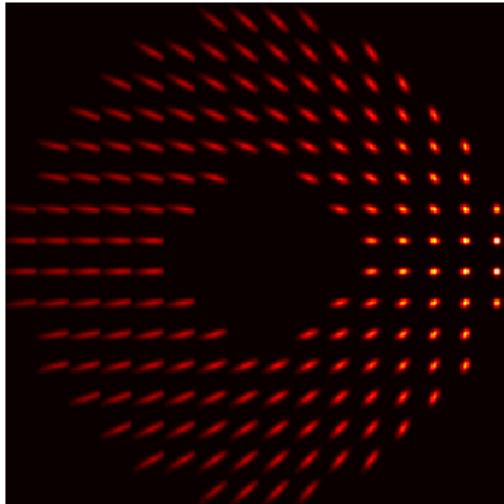


Noise on LGS sub-apertures

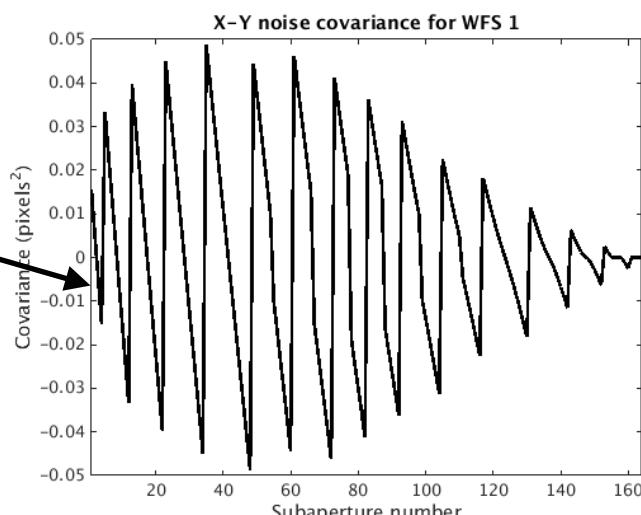
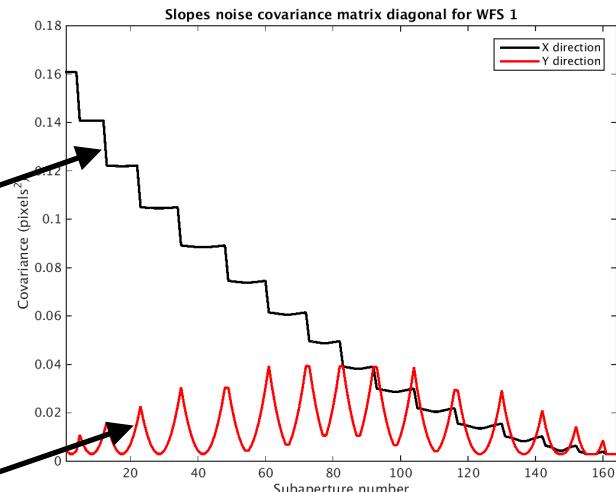
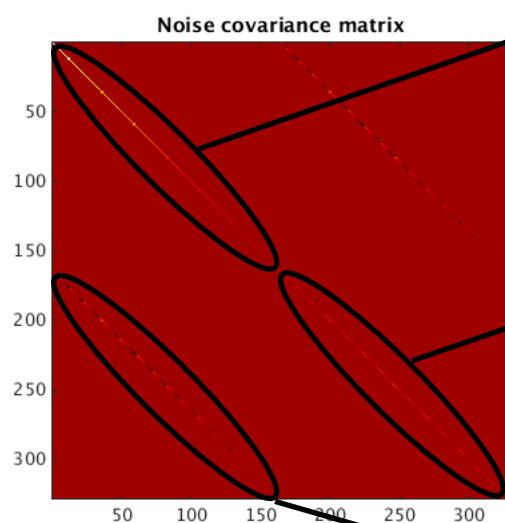
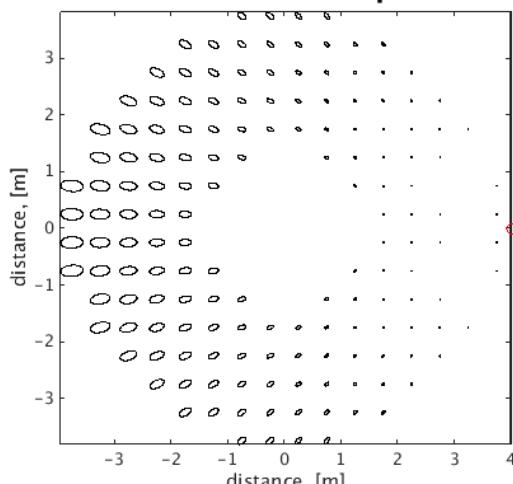


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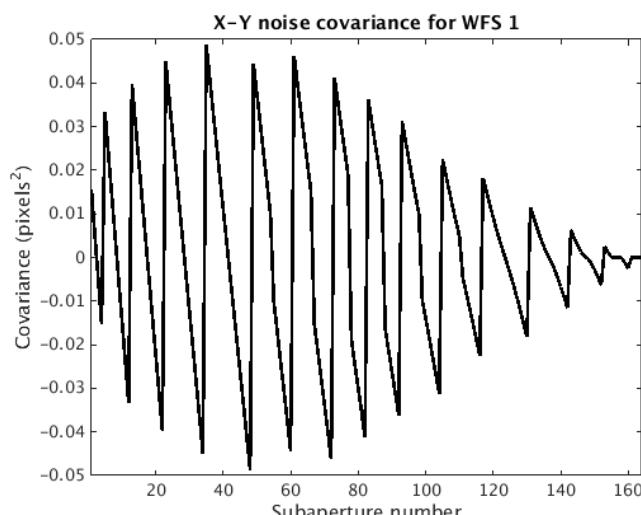
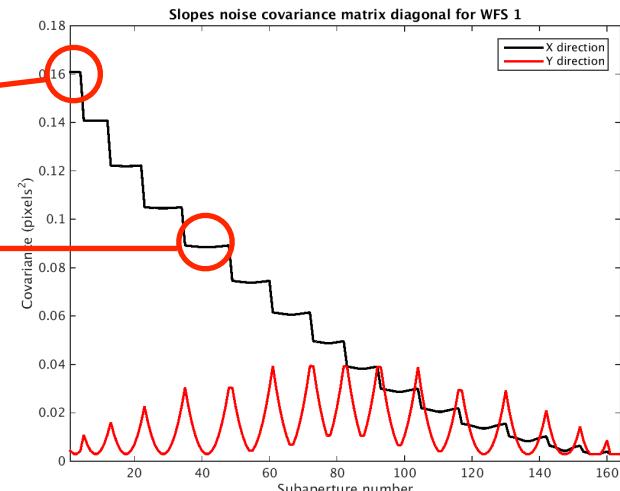
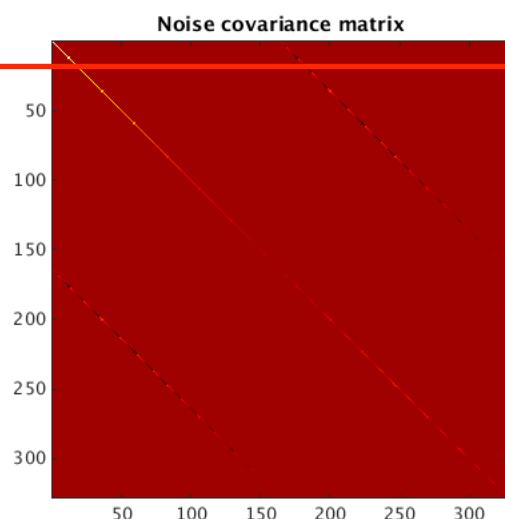
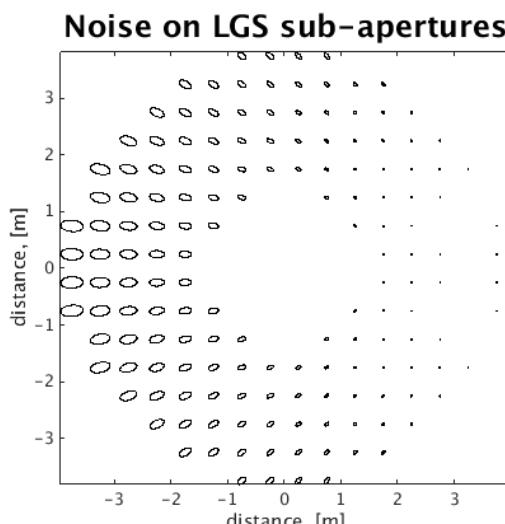
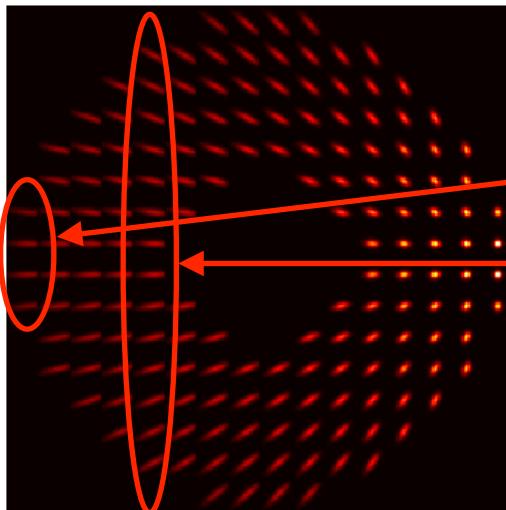


Noise on LGS sub-apertures



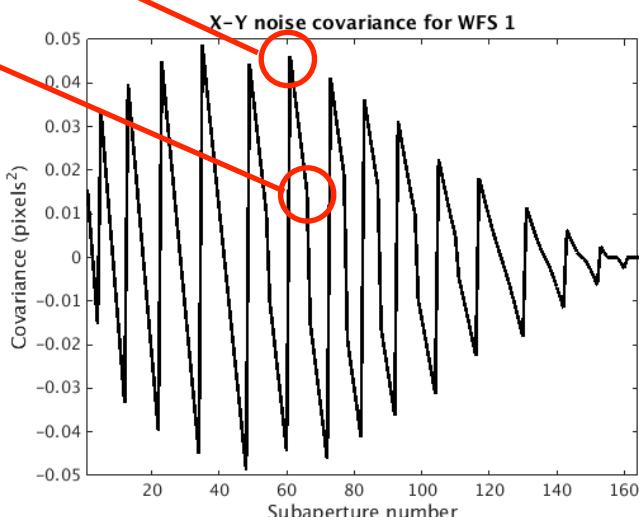
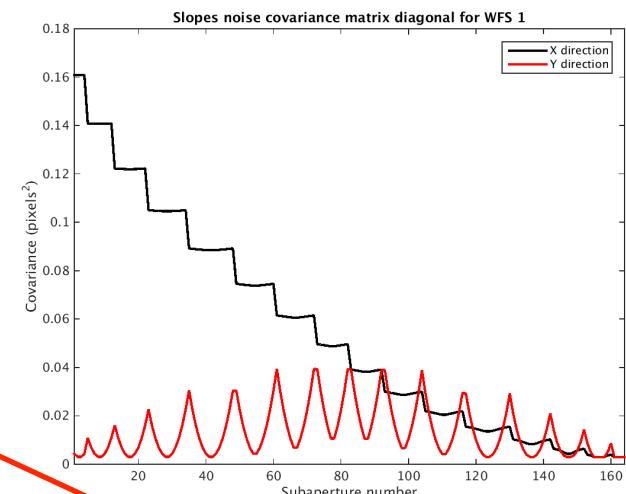
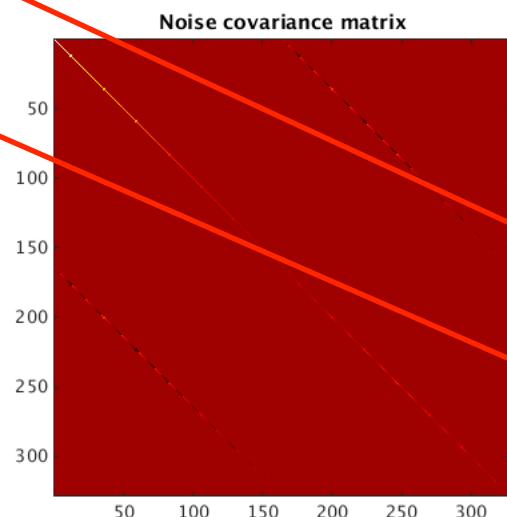
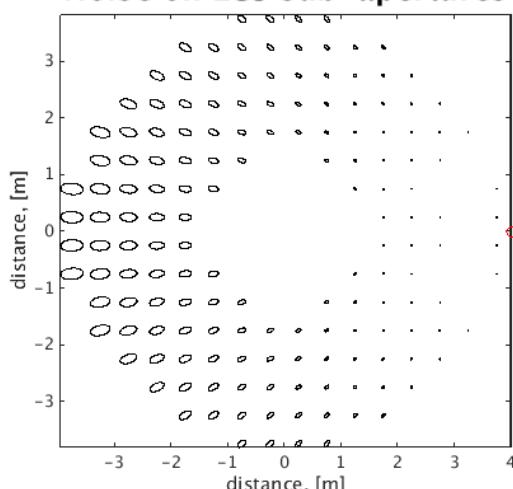
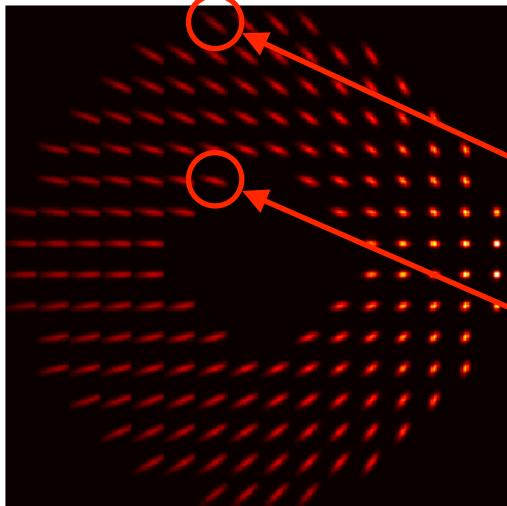
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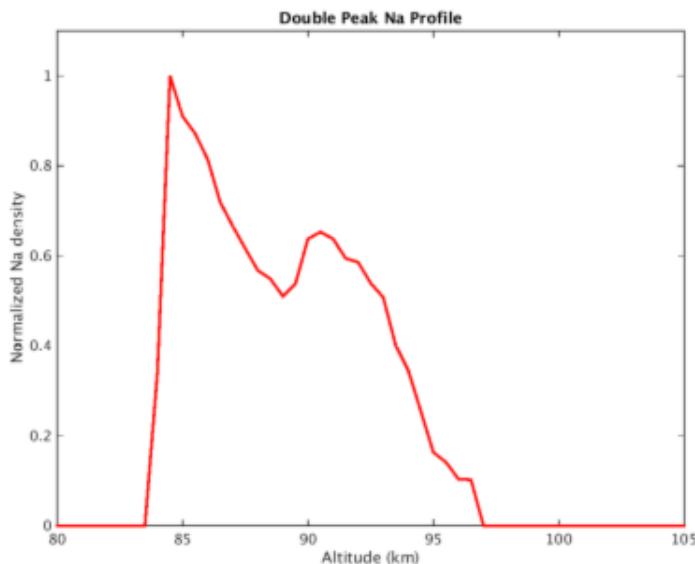


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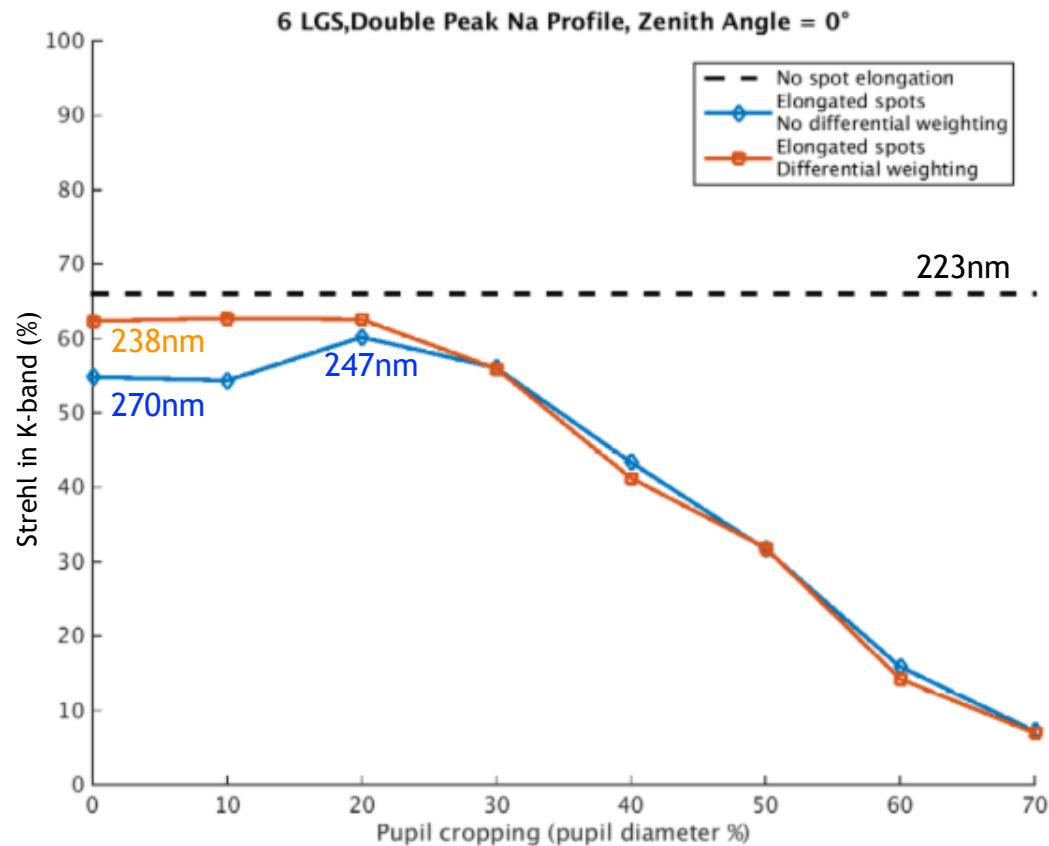
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Subaperture weighting+cropping



Full E-ELT simulation:
Pupil diameter : 37m
74x74 SH LGS WFS
75x75 actuator DM
Subaperture pitch : 50cm
Subaperture FOV : 10" (1" pixels)
6 LGS-Asterism diameter : 42"
30% central obscuration
High flux

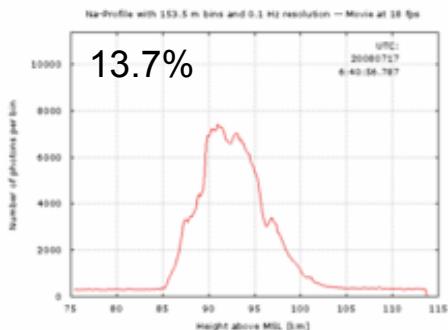


WFE due to elongation reduced from 152nm to 83nm RMS when using subaperture differential weighting

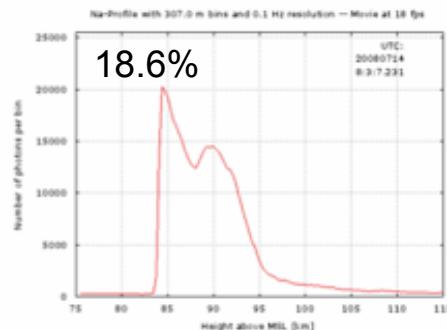
but Na profile not always bimodal !

Subaperture weighting+cropping

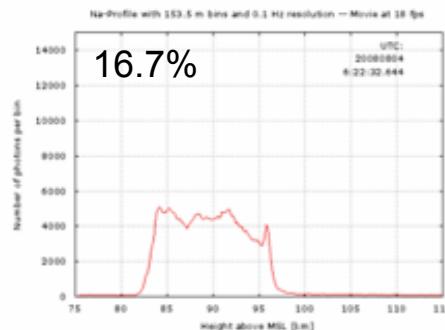
Single peak



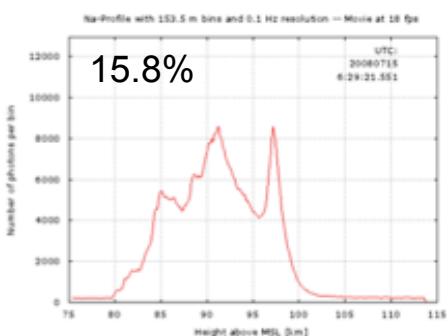
Double peak



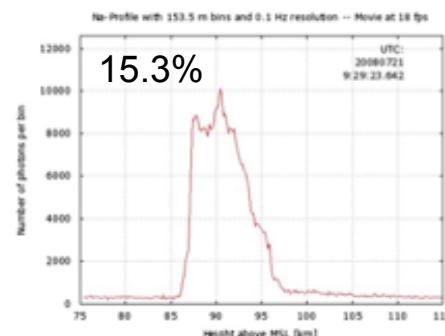
Top hat



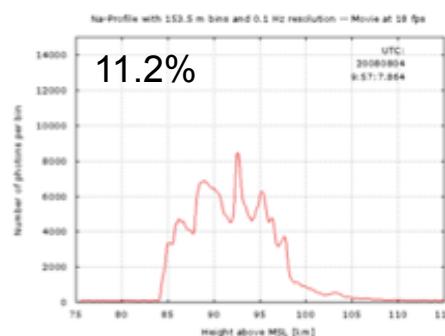
Very wide



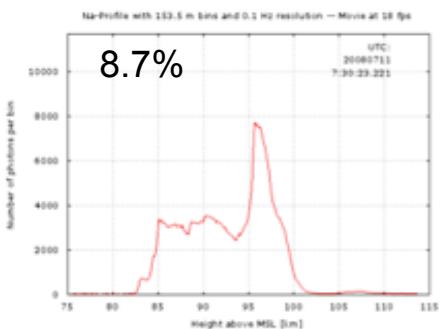
Very narrow



Multi peak

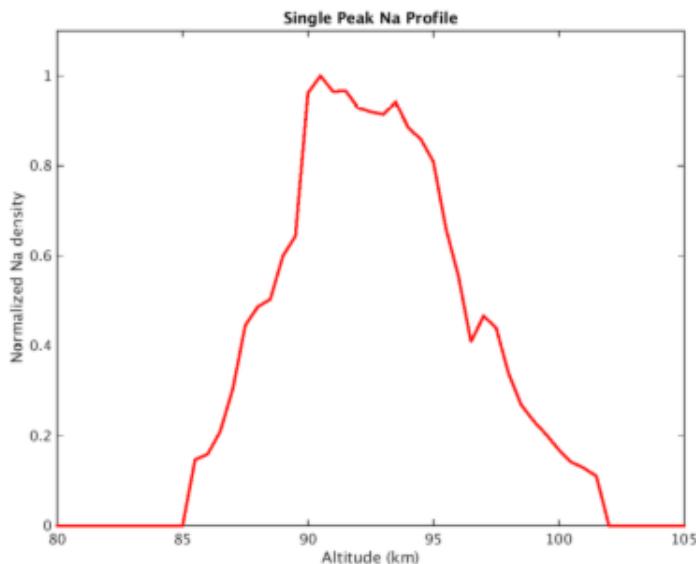


Top hat with pk.

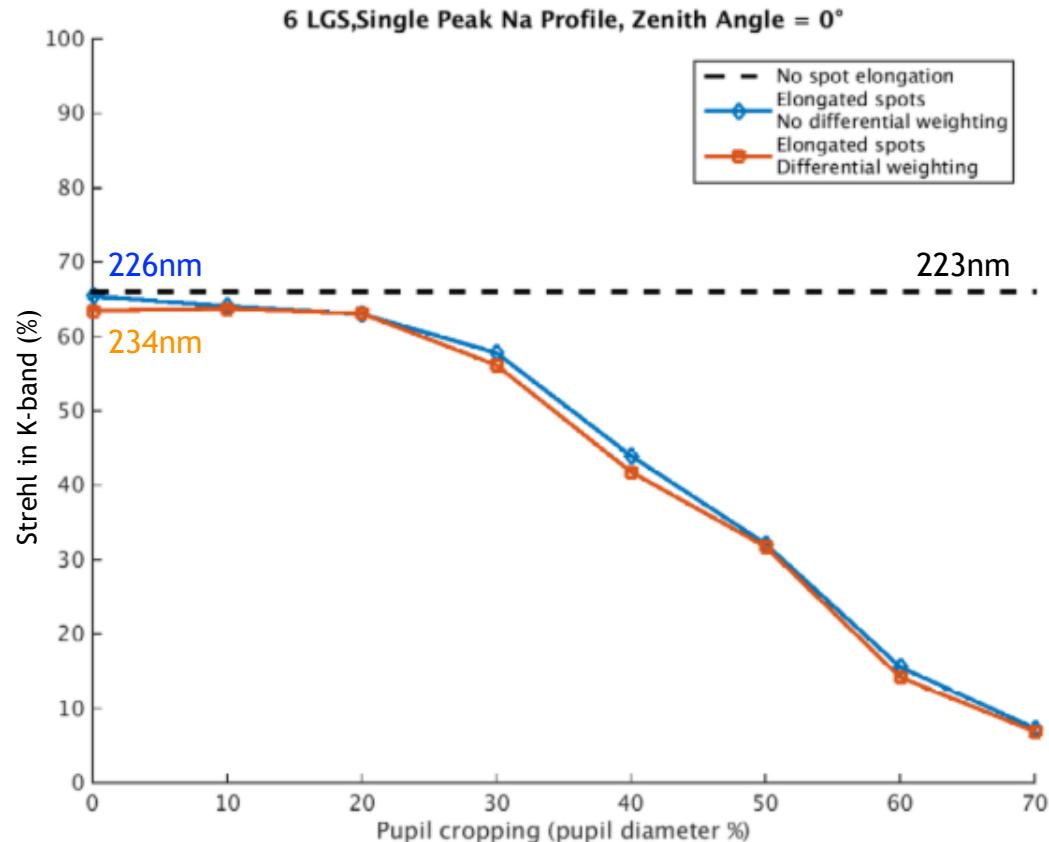


Pfrommer and Hickson 2011

Subaperture weighting+cropping



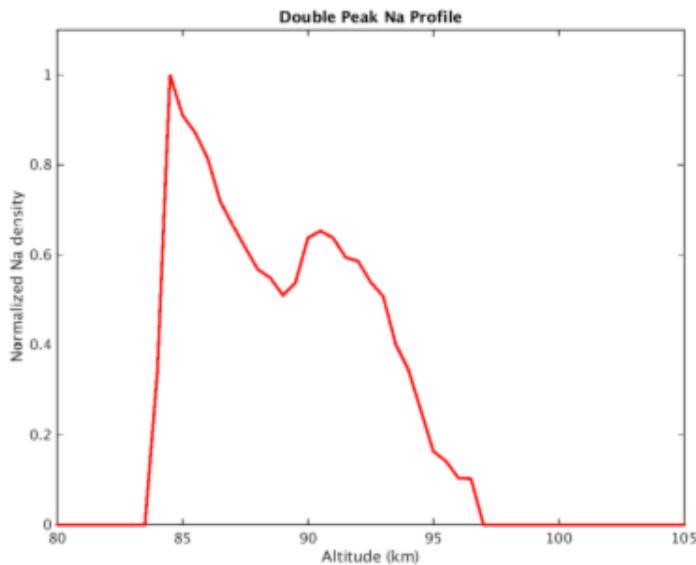
Full E-ELT simulation:
Pupil diameter : 37m
74x74 SH LGS WFS
75x75 actuator DM
Subaperture pitch : 50cm
Subaperture FOV : 10" (1" pixels)
6 LGS-Asterism diameter : 42"
30% central obscuration
High flux



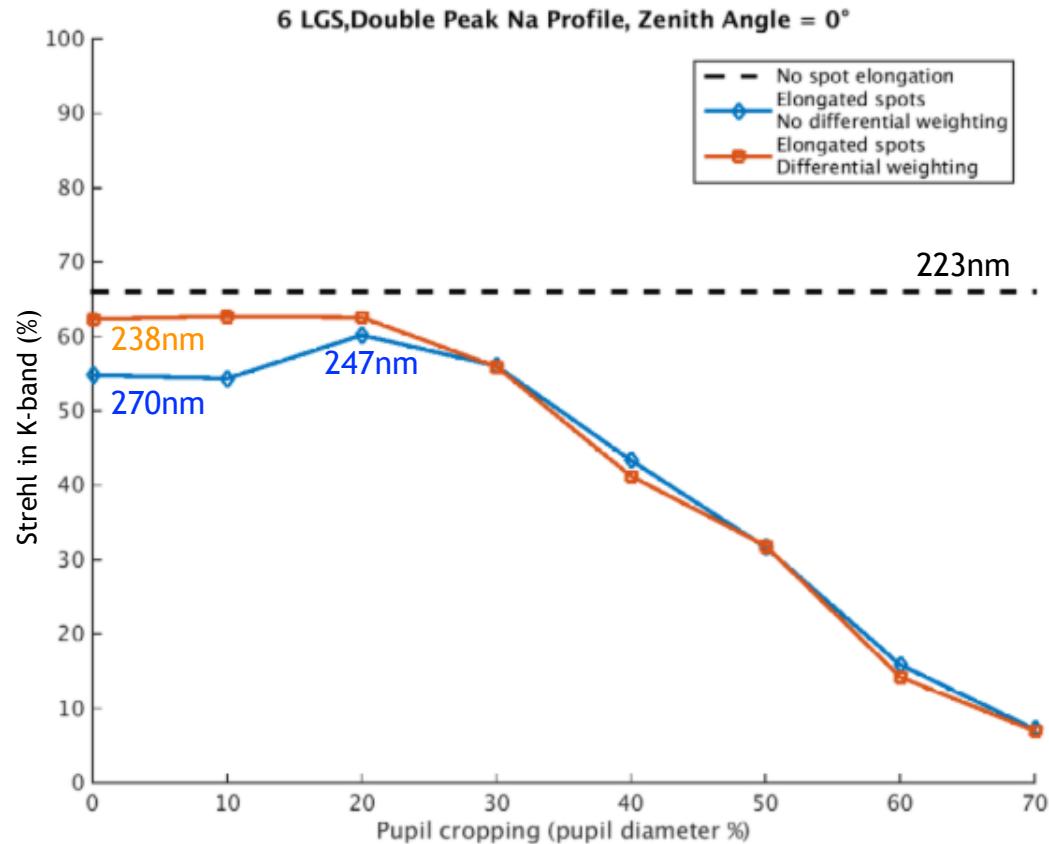
Narrow spot mostly within the subaperture even opposite from LLT

Weighting over-regularizes already good information

Subaperture weighting+cropping

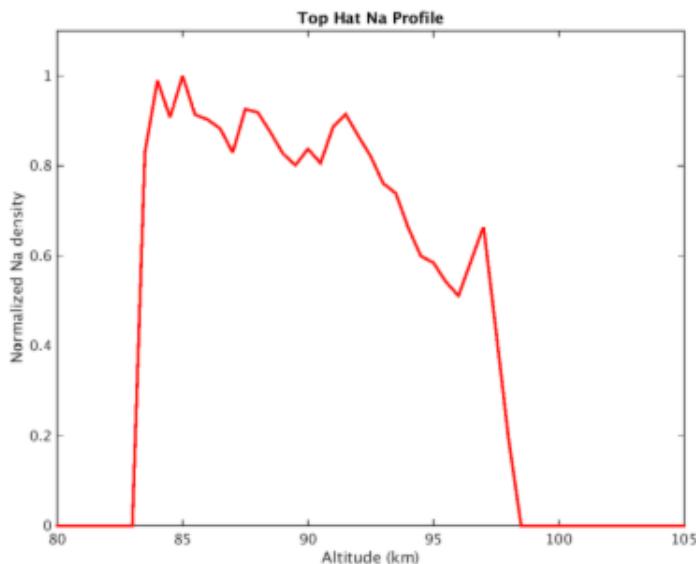


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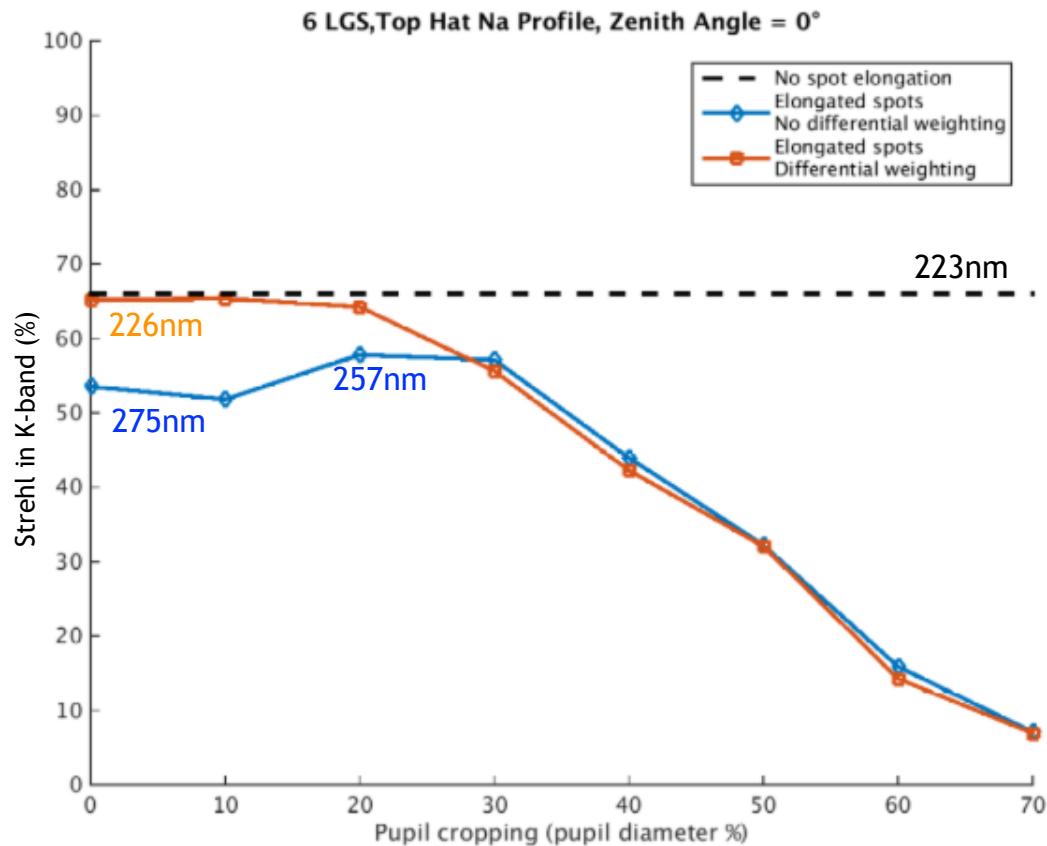


WFE due to elongation reduced from 152nm to 83nm RMS when using subaperture differential weighting

Subaperture weighting+cropping

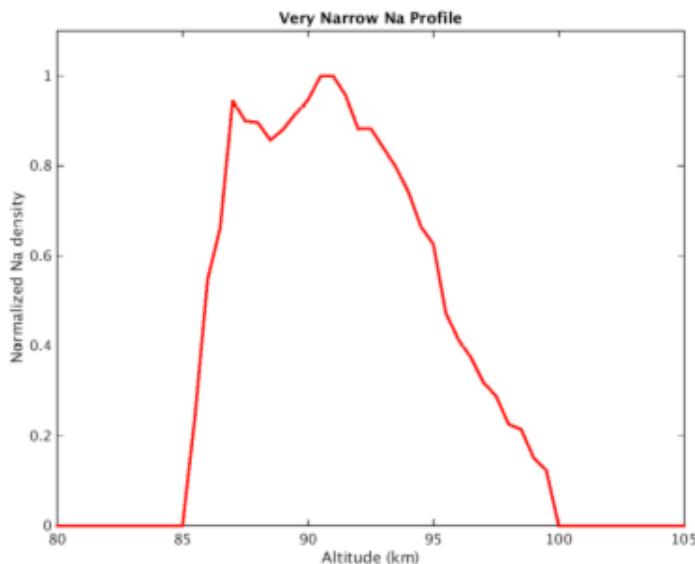


Full E-ELT simulation:
Pupil diameter : 37m
74x74 SH LGS WFS
75x75 actuator DM
Subaperture pitch : 50cm
Subaperture FOV : 10" (1" pixels)
6 LGS-Asterism diameter : 42"
30% central obscuration
High flux

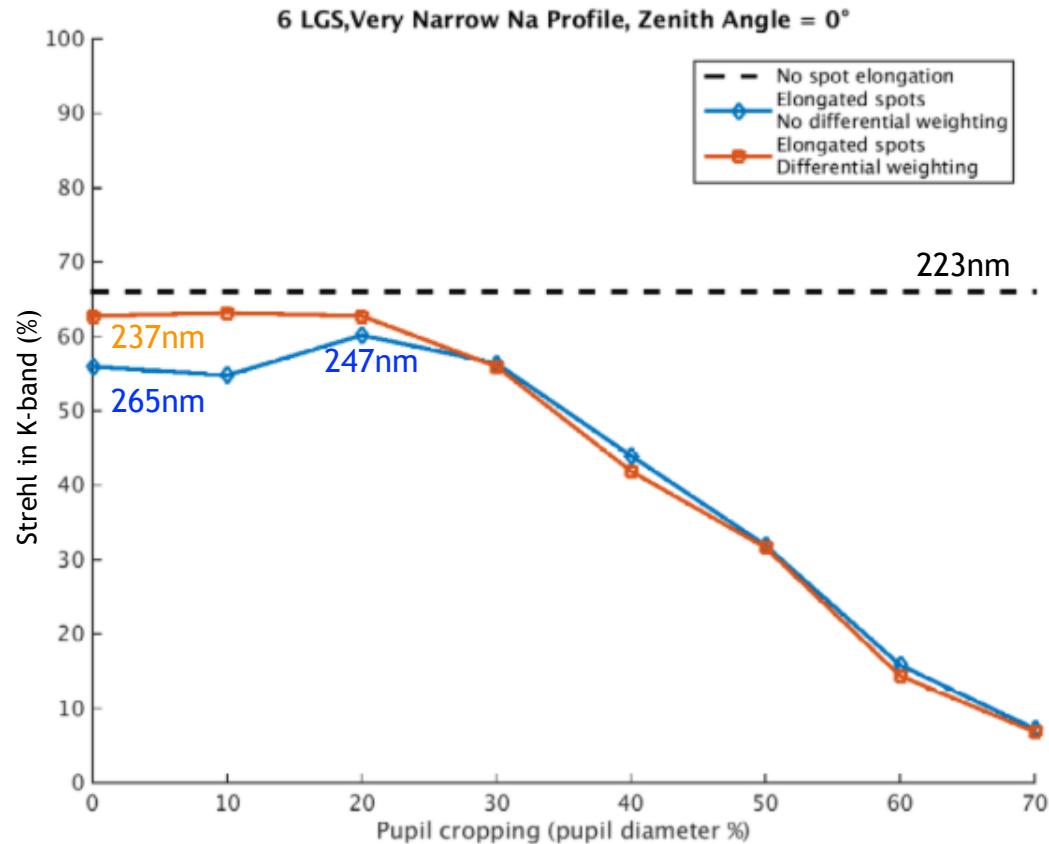


Differential weighting gives almost ideal performance

Subaperture weighting+cropping



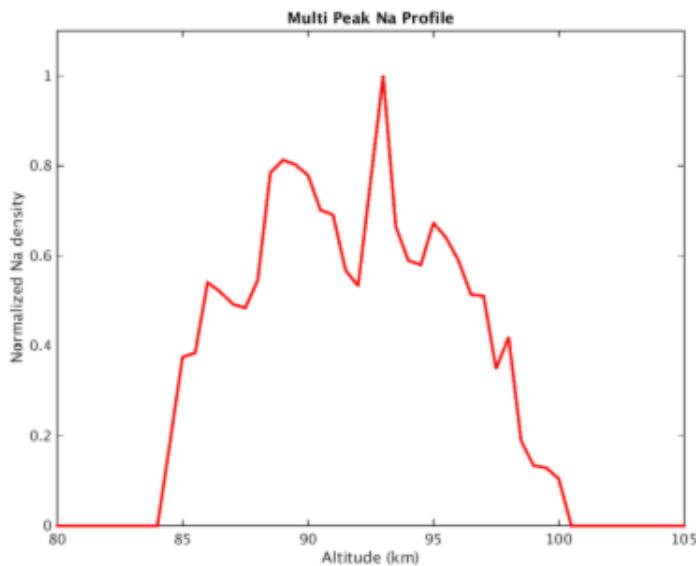
Full E-ELT simulation:
Pupil diameter : 37m
74x74 SH LGS WFS
75x75 actuator DM
Subaperture pitch : 50cm
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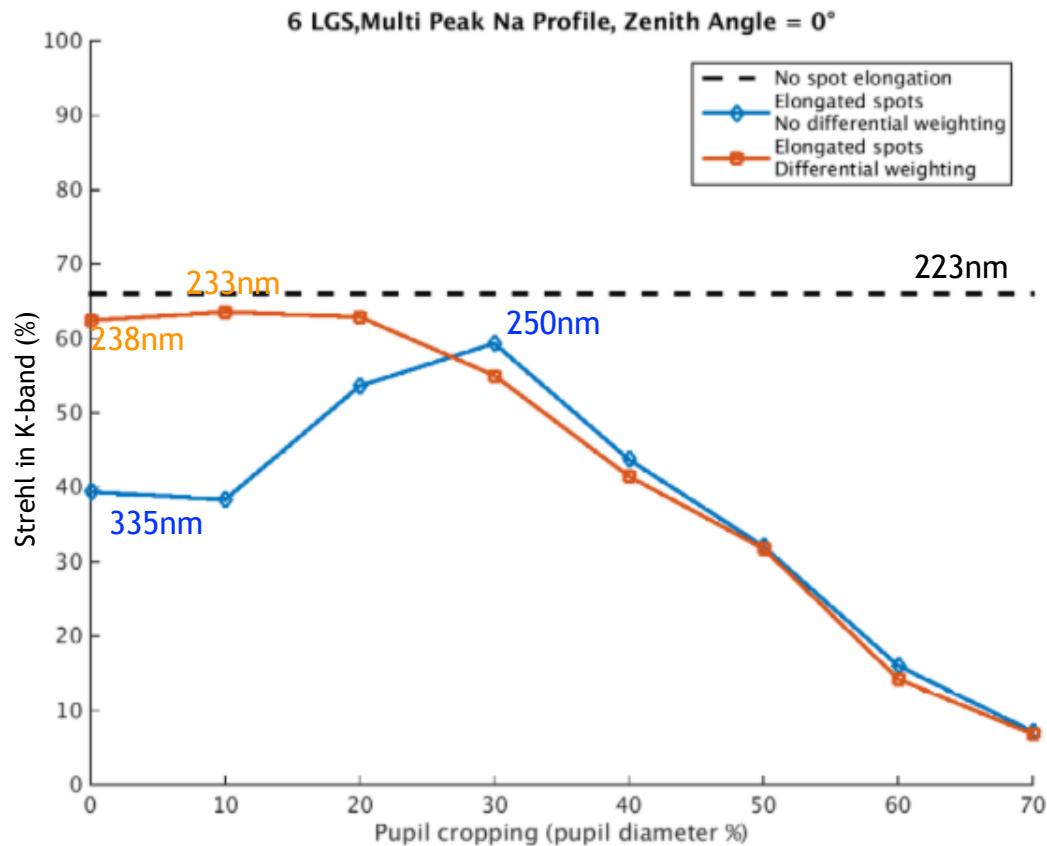
WFE due to elongation reduced from 143nm to 80nm RMS when using subaperture differential weighting

Best ‘cropping’ (20%)
WFE due to elongation = 106nm

Subaperture weighting+cropping



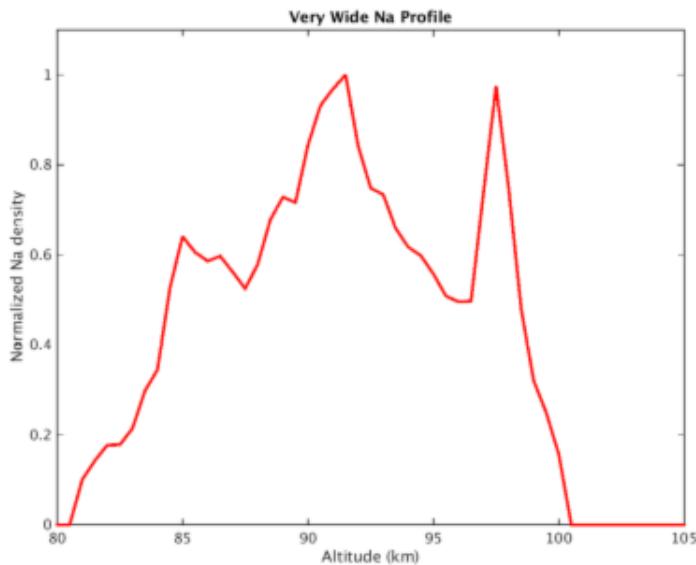
Full E-ELT simulation:
Pupil diameter : 37m
74x74 SH LGS WFS
75x75 actuator DM
Subaperture pitch : 50cm
Subaperture FOV : 10" (1" pixels)
6 LGS-Asterism diameter : 42"
30% central obscuration
High flux



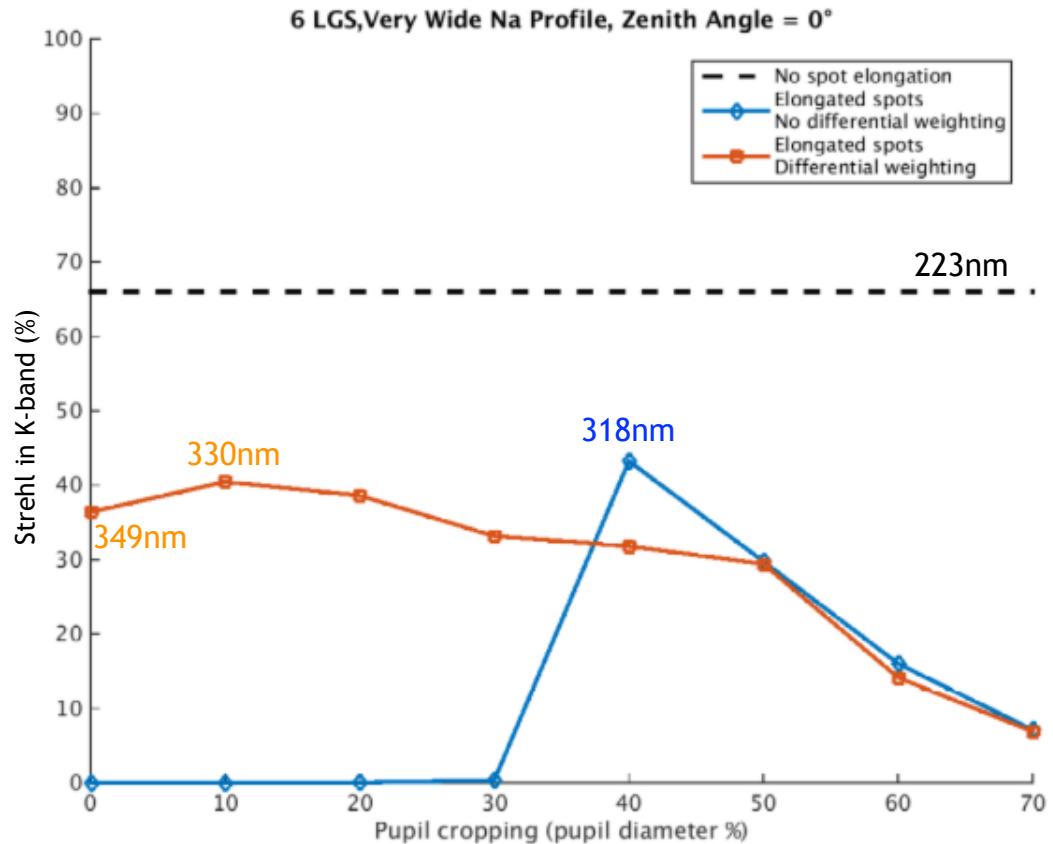
WFE due to elongation reduced from 250nm to 83nm RMS when using subaperture differential weighting

WFE reduced to 67nm RMS when combining weighting and cropping

Subaperture weighting+cropping

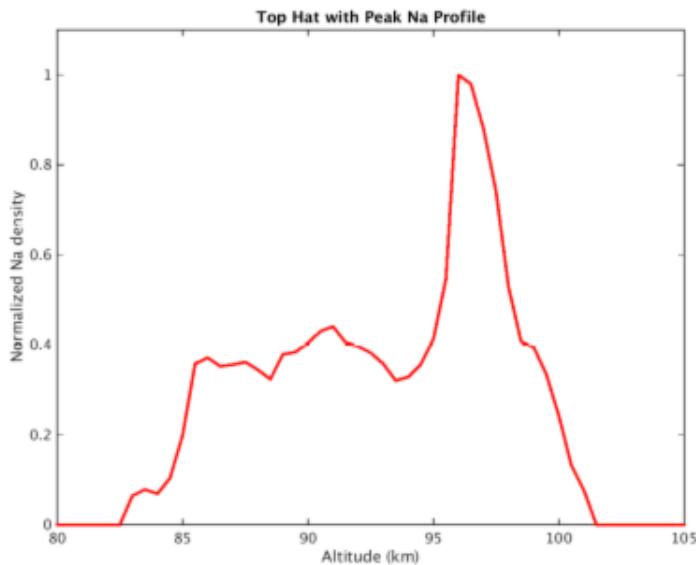


Full E-ELT simulation:
 Pupil diameter : 37m
 74x74 SH LGS WFS
 75x75 actuator DM
 Subaperture pitch : 50cm
 Subaperture FOV : 10" (1" pixels)
 6 LGS-Asterism diameter : 42"
 30% central obscuration
 High flux

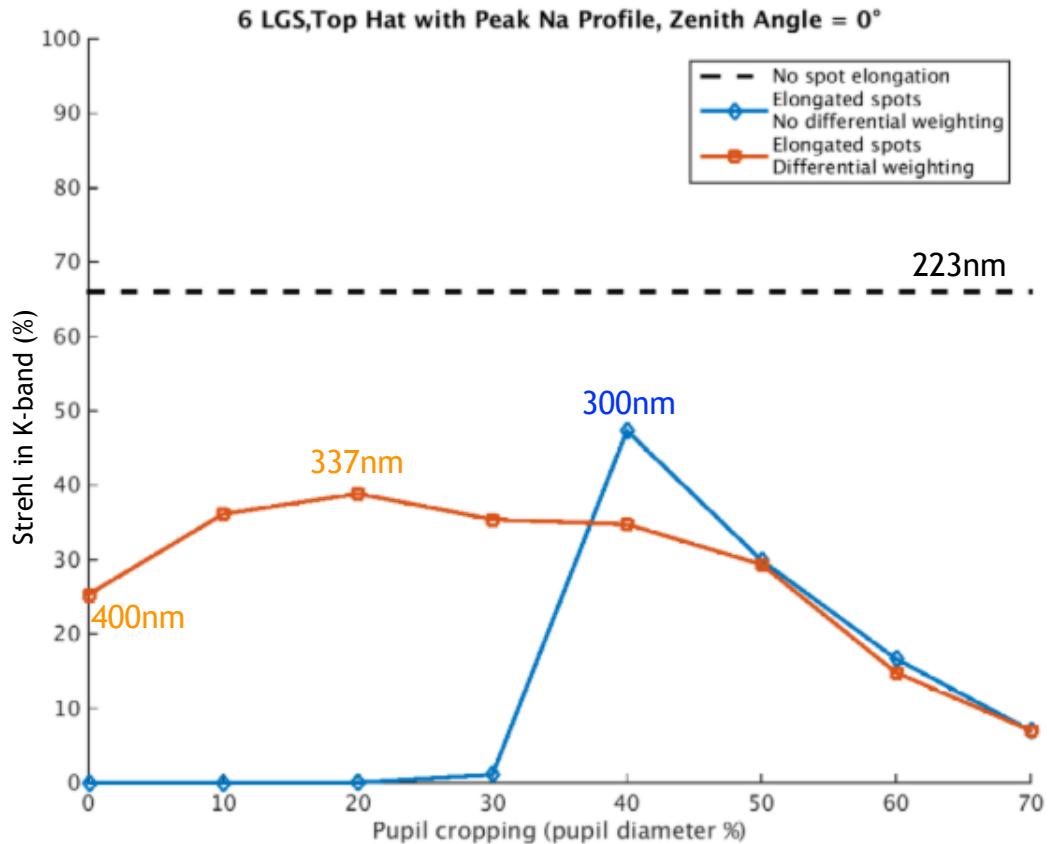


Very perturbed profile with sharp peak at high altitude
 Unstable loop when taking into account all subapertures
 Weighting mandatory for full pupil
 High cropping (40%) improves performance over weighting but far from ideal

Subaperture weighting+cropping

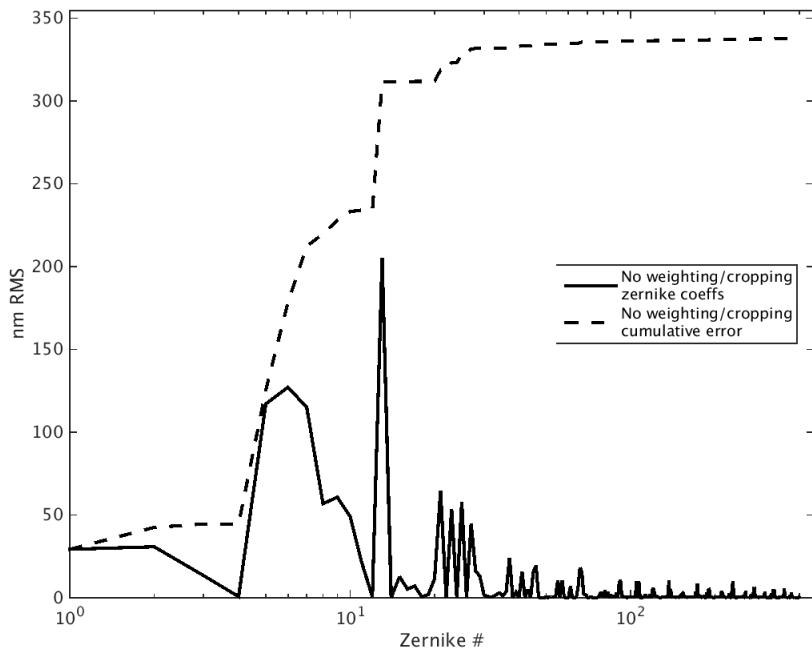


Full E-ELT simulation:
Pupil diameter : 37m
74x74 SH LGS WFS
75x75 actuator DM
Subaperture pitch : 50cm
Subaperture FOV : 10'' (1'' pixels)
6 LGS-Asterism diameter : 42''
30% central obscuration
High flux

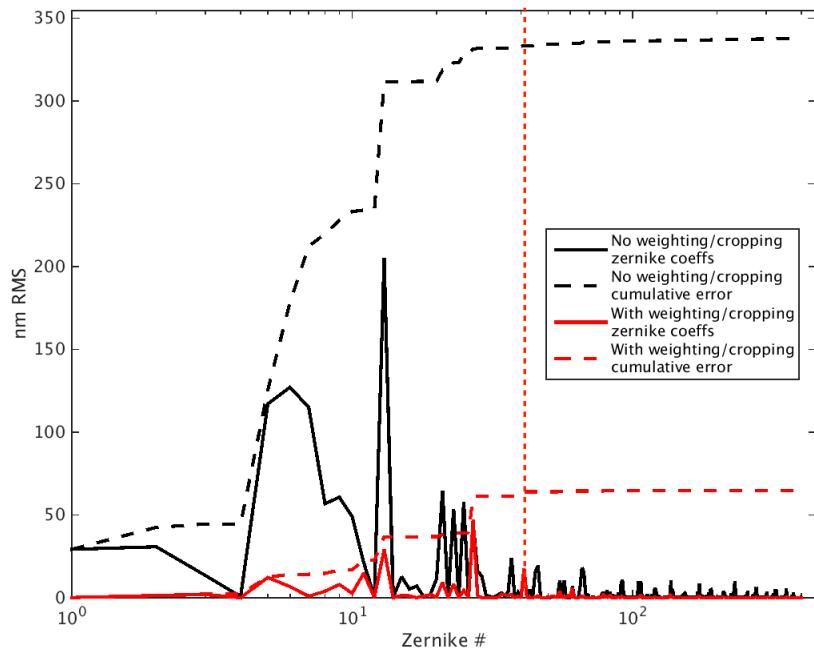


Very perturbed profile with sharp peak at high altitude
Unstable loop when taking into account all subapertures
Weighting mandatory for full pupil
High cropping (40%) improves performance over weighting but far from ideal

How many modes do we need to measure to remain below a prescribed WFE due to spot elongation ?

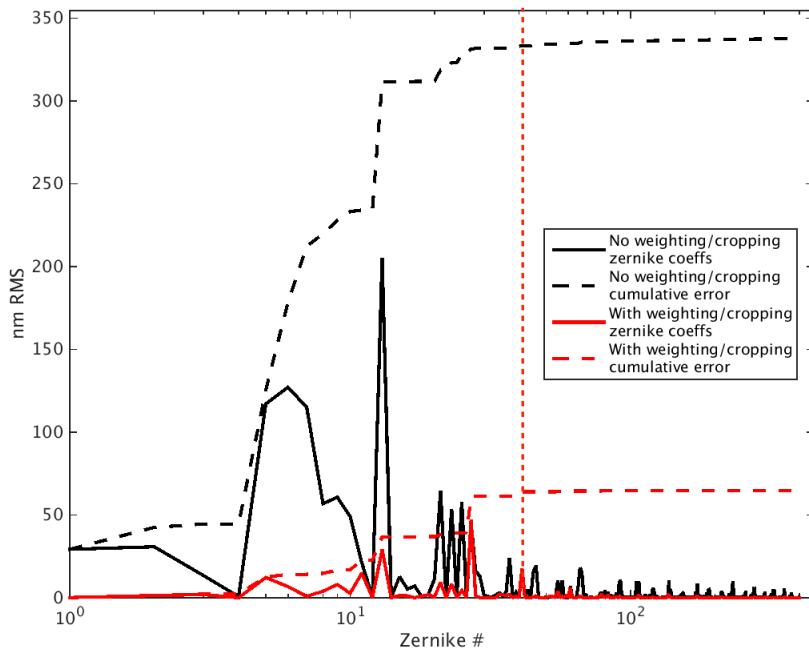


How many modes do we need to measure to remain below a prescribed WFE due to spot elongation ?



Truth sensor

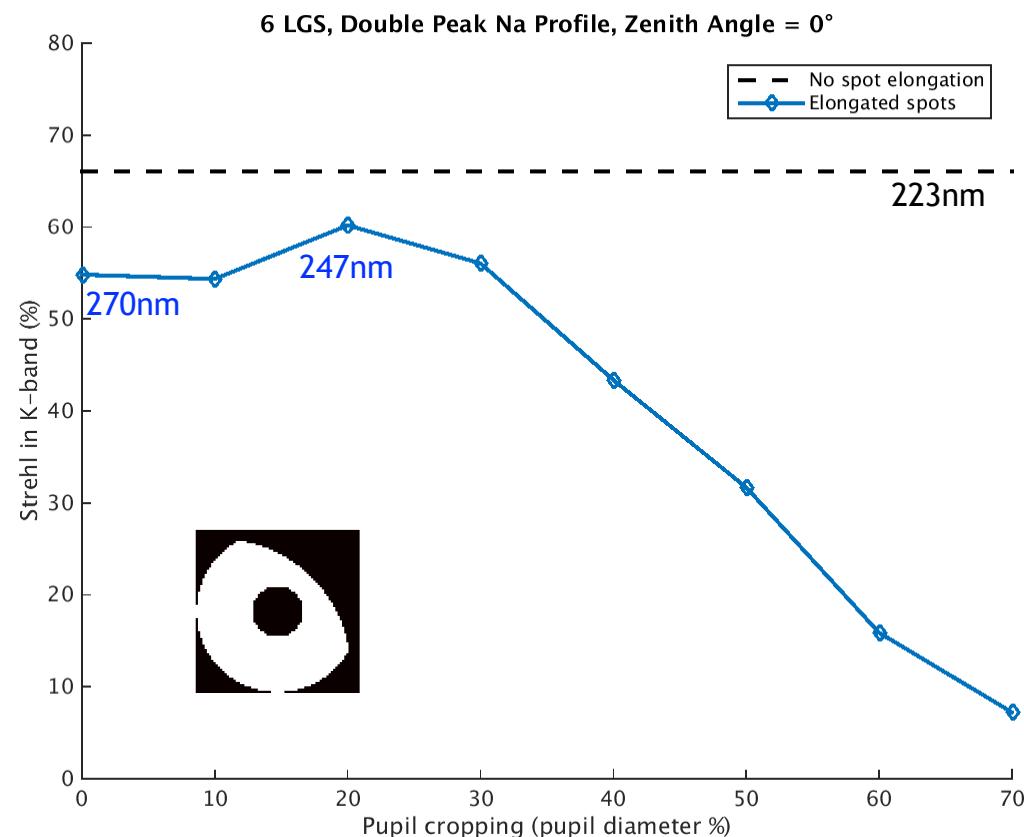
How many modes do we need to measure to remain below a prescribed WFE due to spot elongation ?



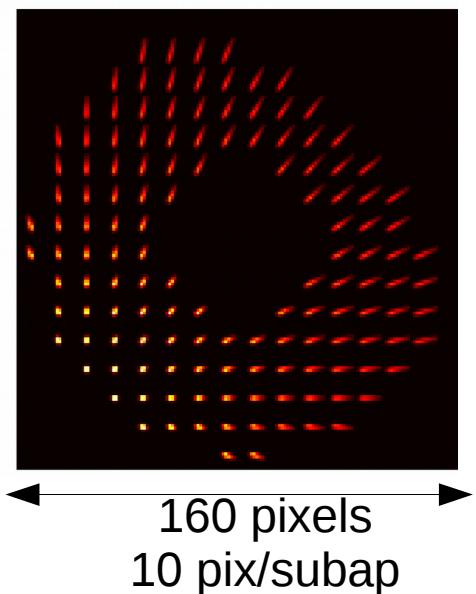
	Single	Top hat	Double	Very narrow	Multi peak	Very wide	Top hat +peak
10nm	z1	z10	z14	z14	z28	z42	z56
20nm	z1	z1	z1	z1	z1	z42	z42
30nm	z1	z1	z1	z1	z1	z42	z42
40nm	z1	z1	z1	z1	z1	z28	z28
50nm	z1	z1	z1	z1	z1	z28	z28

8x8 truth sensor sufficient to remain below 10nm at all times

Corner Shack-Hartmann

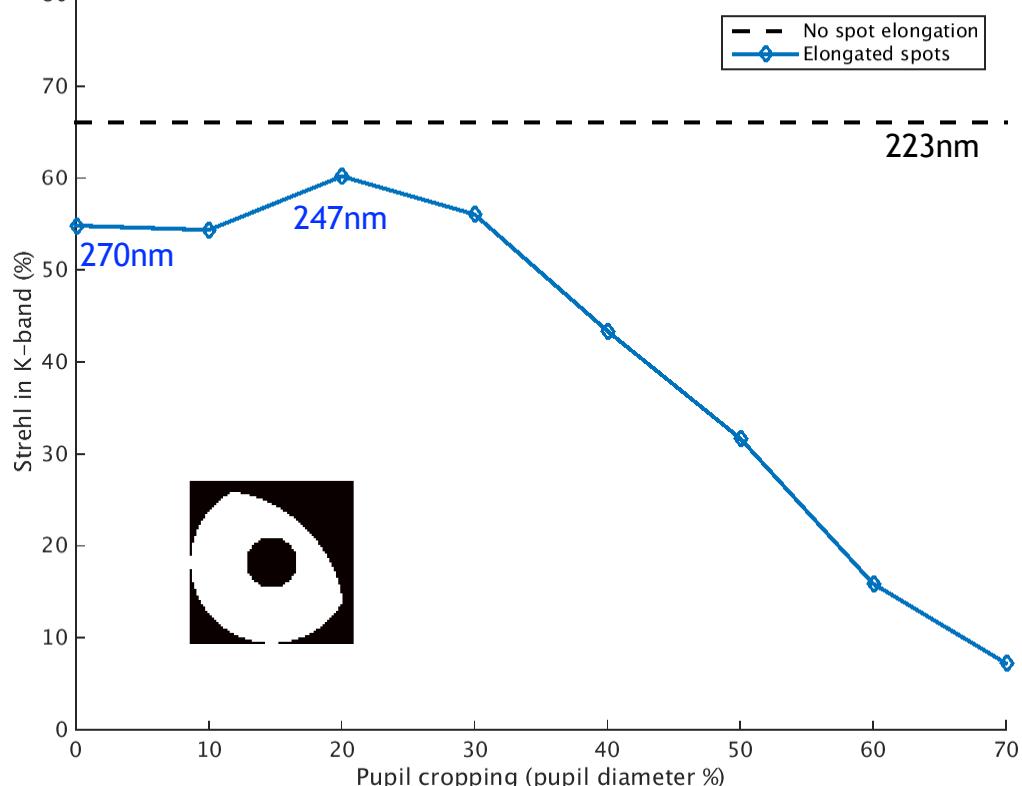


Conventional SH

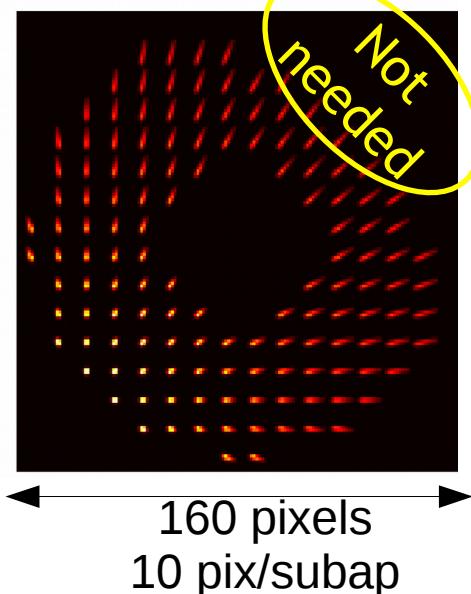


Corner Shack-Hartmann

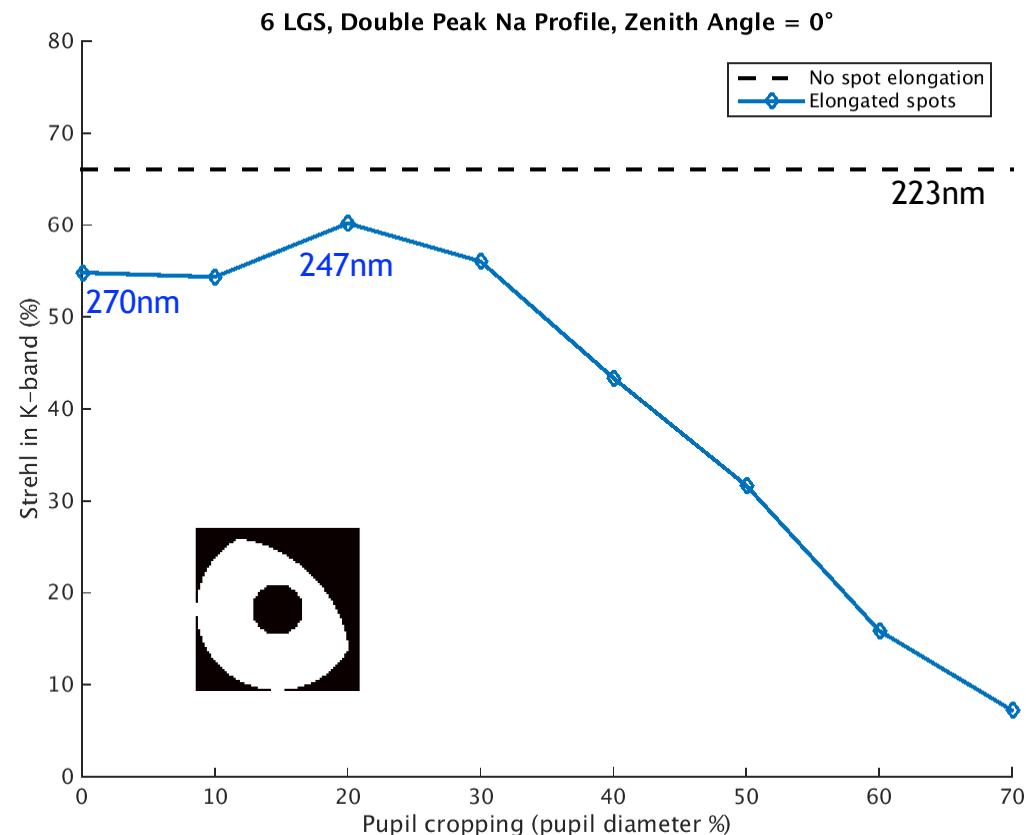
6 LGS, Double Peak Na Profile, Zenith Angle = 0°



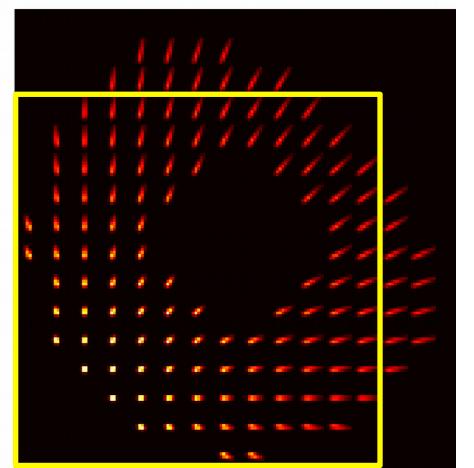
Conventional SH



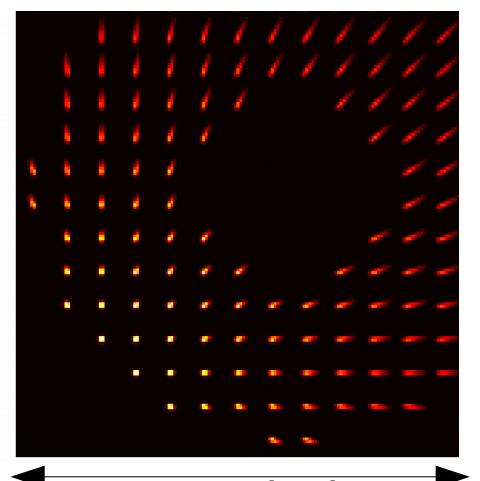
Corner Shack-Hartmann



Conventional SH

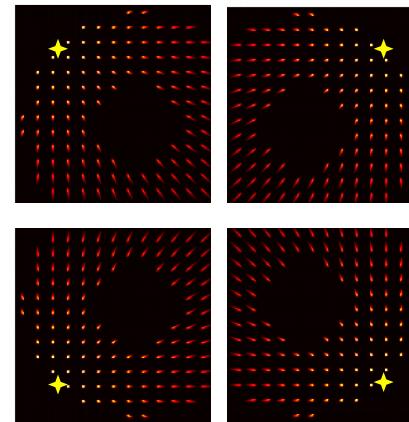
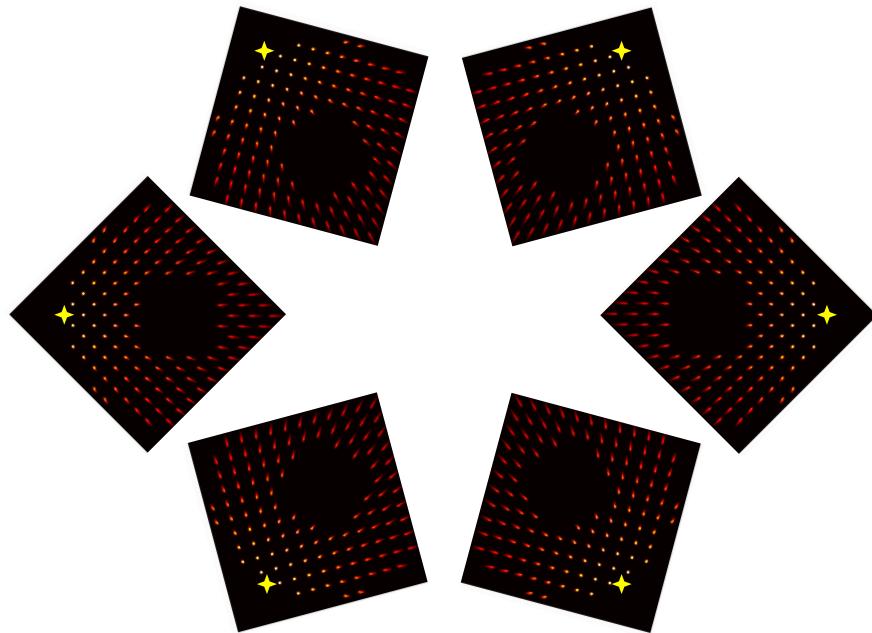


160 pixels
10 pix/subap
Corner SH

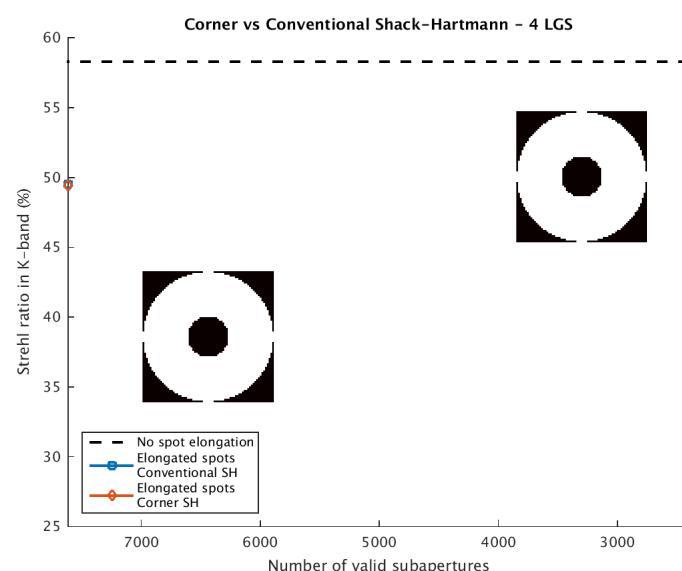
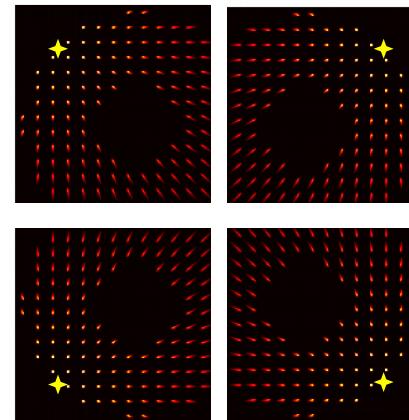
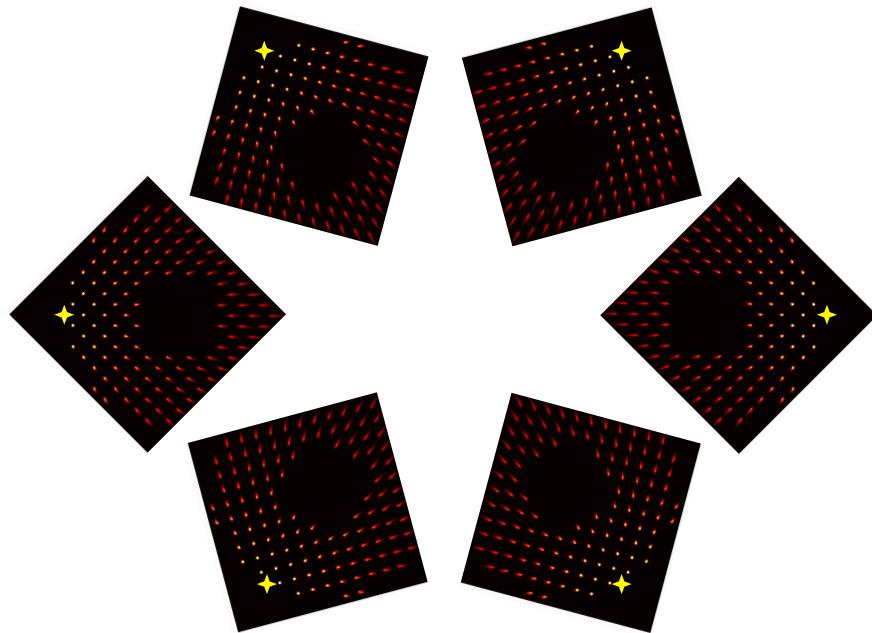


160 pixels
12 pix/subap

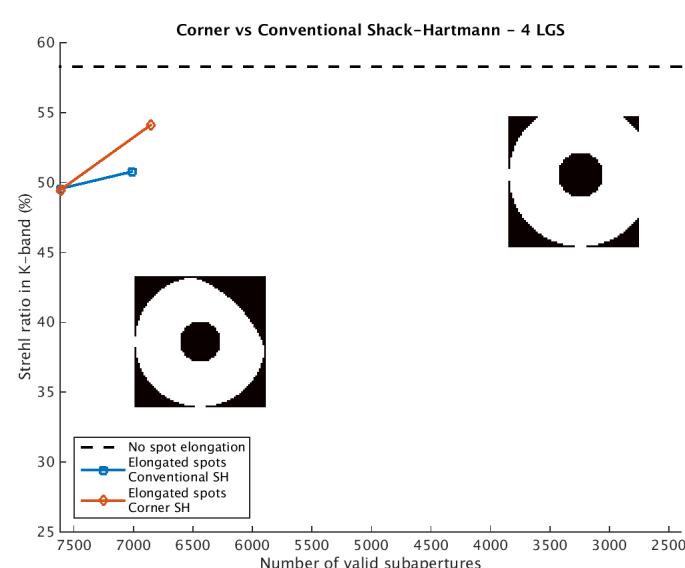
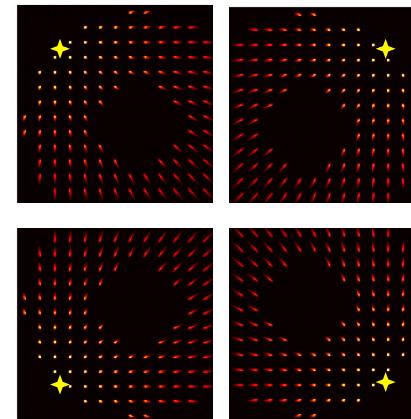
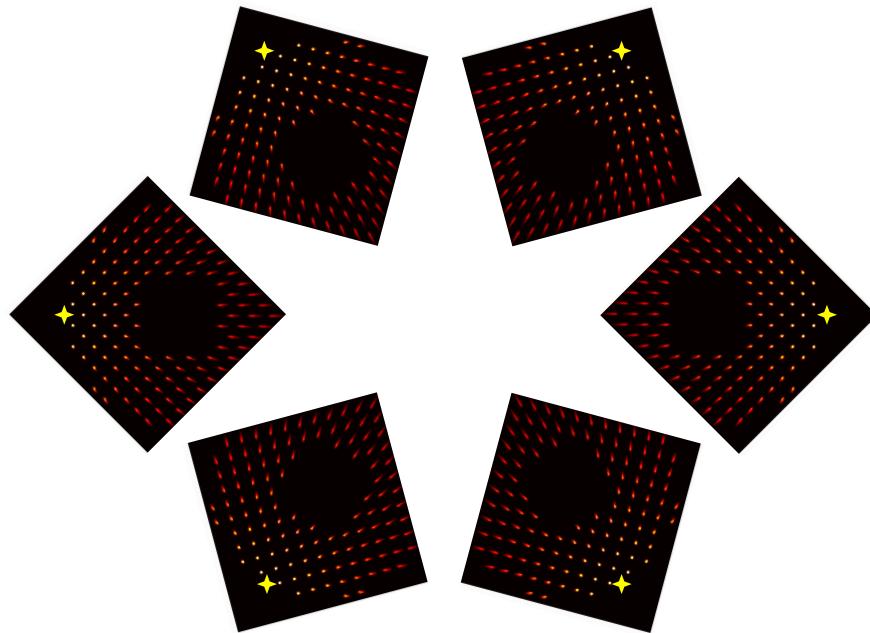
Corner Shack-Hartmann



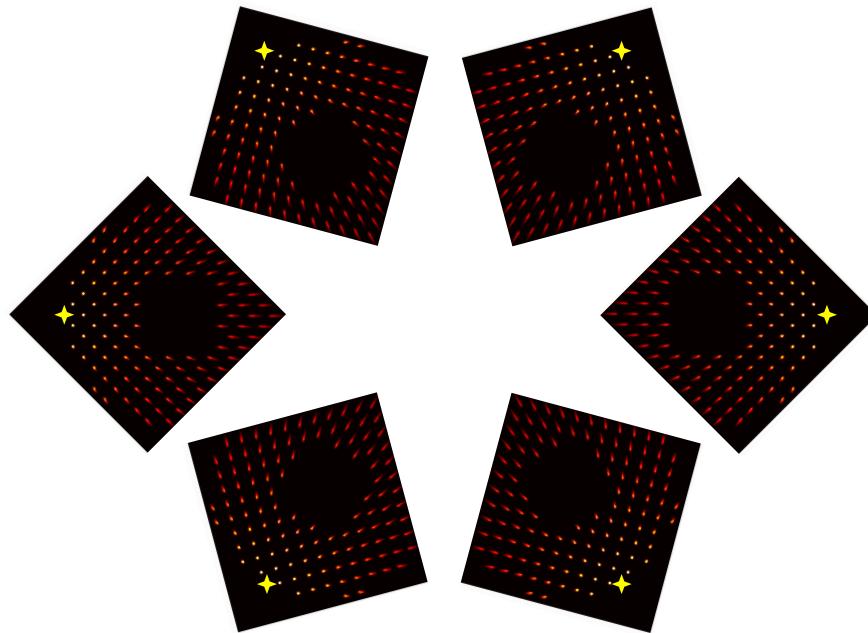
Corner Shack-Hartmann



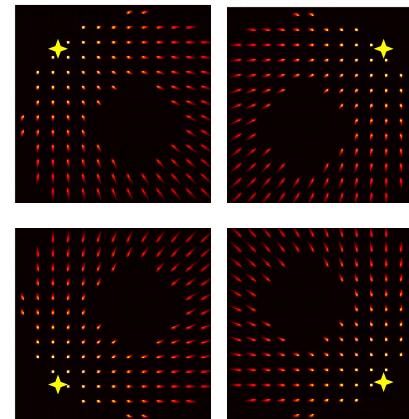
Corner Shack-Hartmann



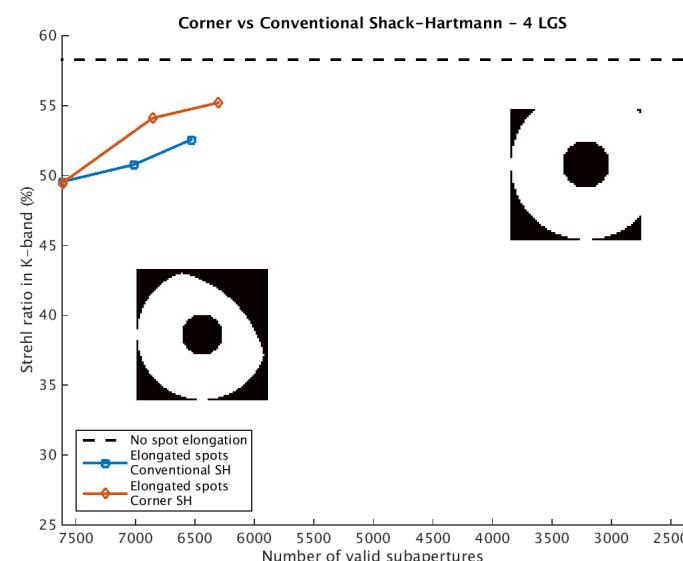
Corner Shack-Hartmann



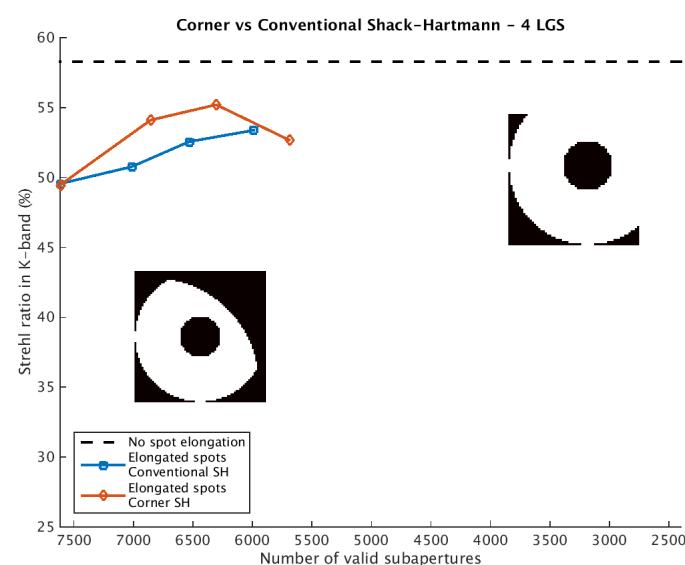
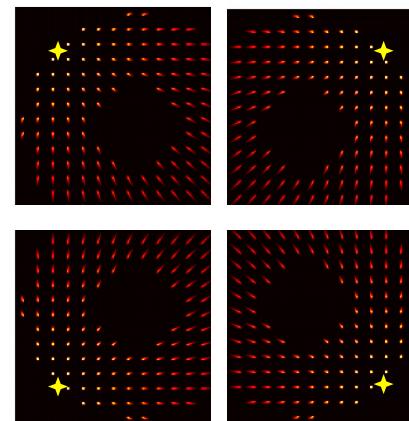
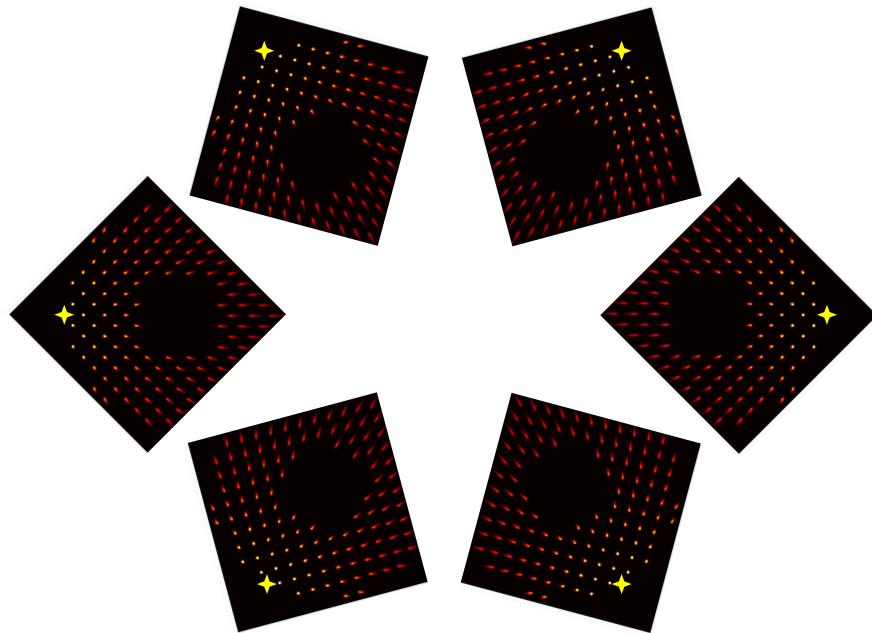
6 LGS configuration



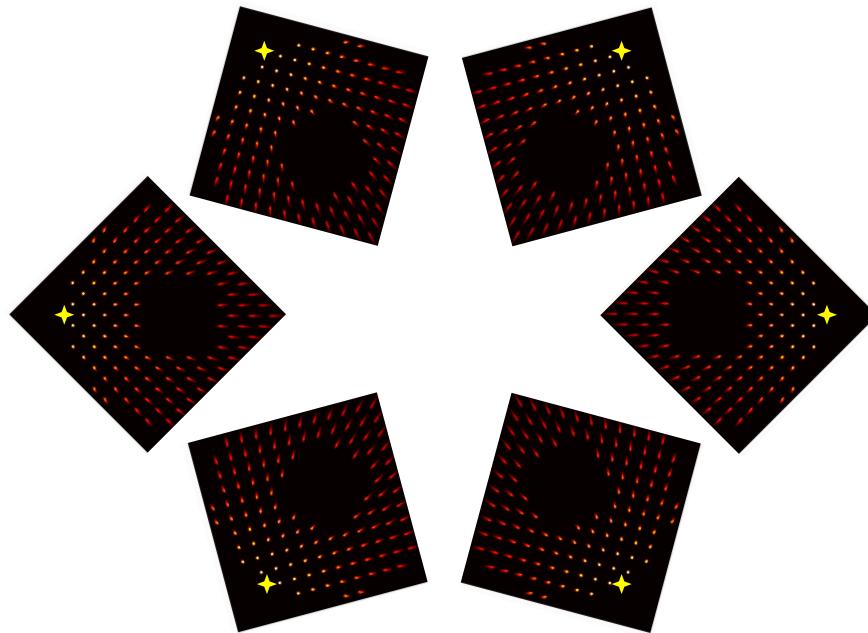
4 LGS configuration



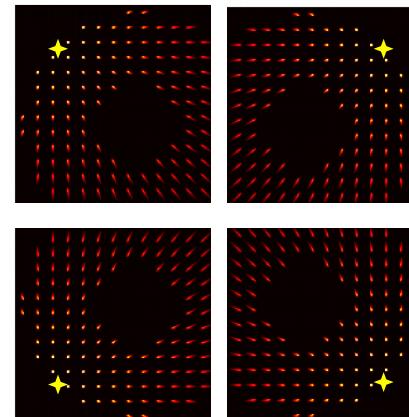
Corner Shack-Hartmann



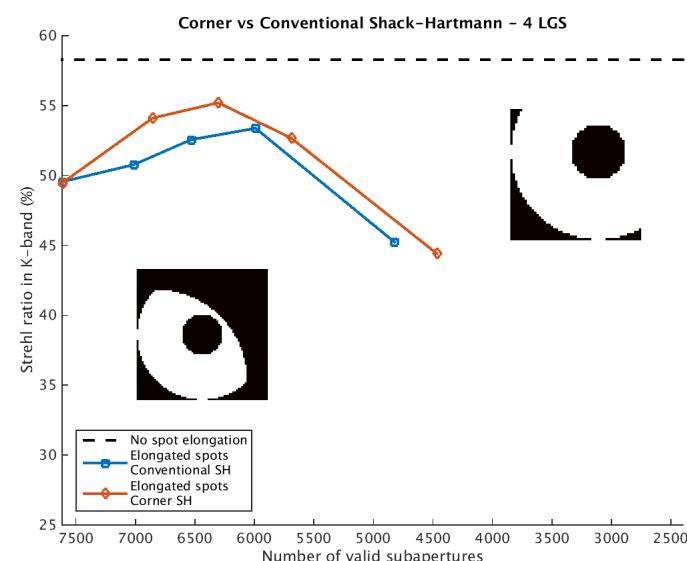
Corner Shack-Hartmann



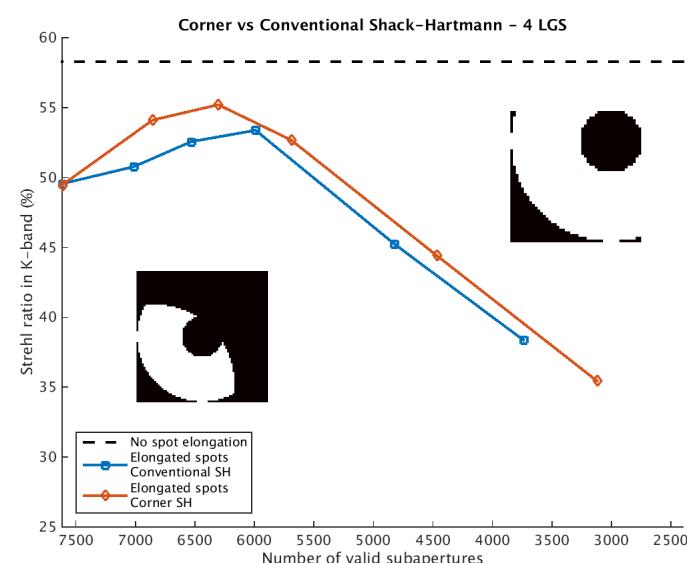
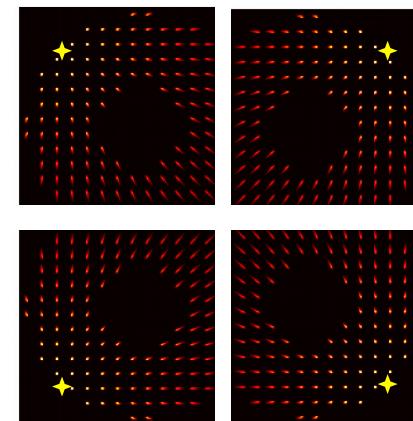
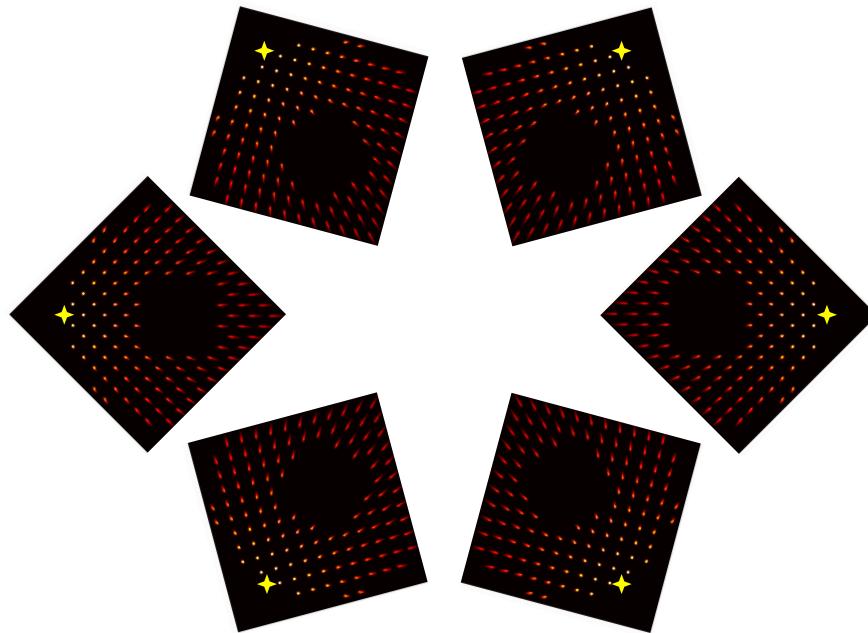
6 LGS configuration



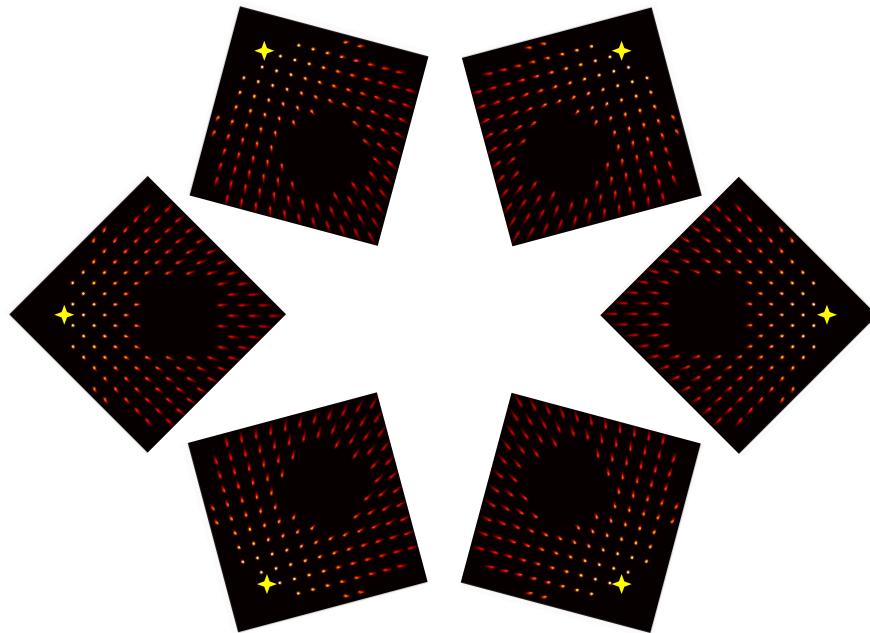
4 LGS configuration



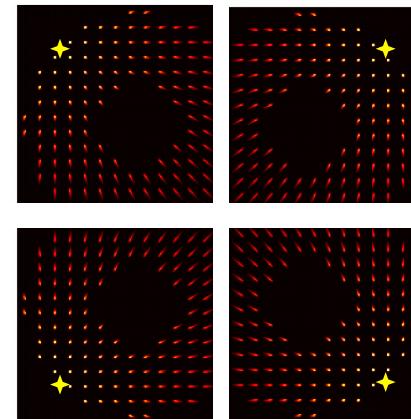
Corner Shack-Hartmann



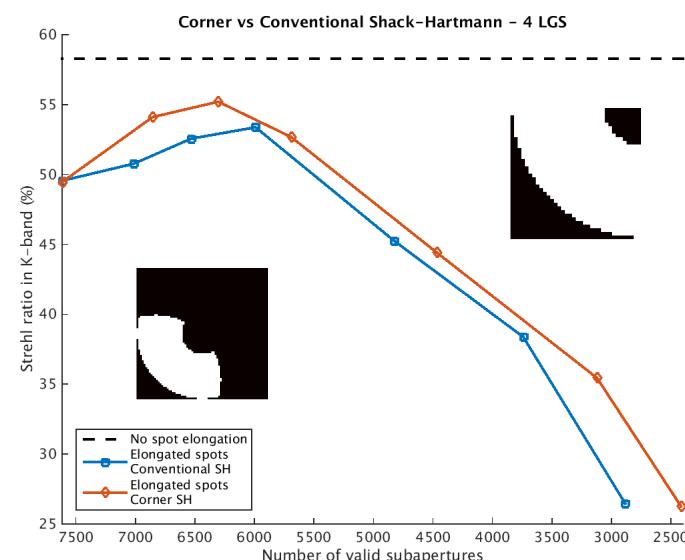
Corner Shack-Hartmann



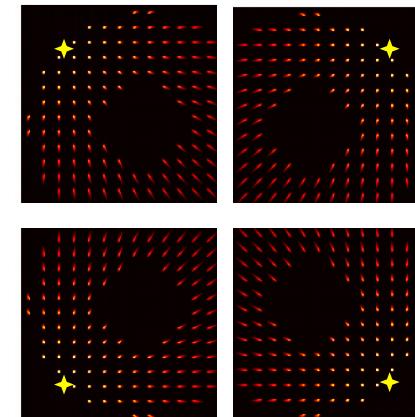
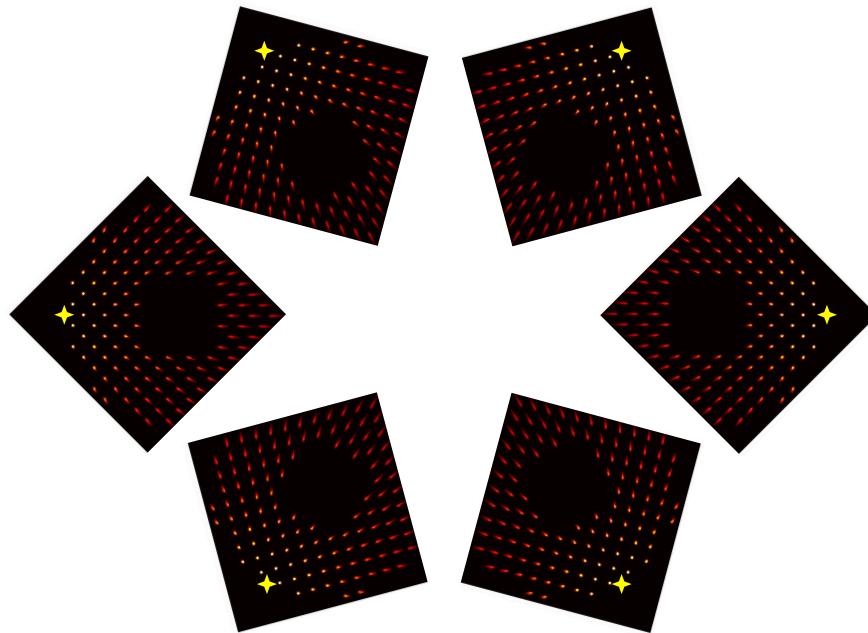
6 LGS configuration



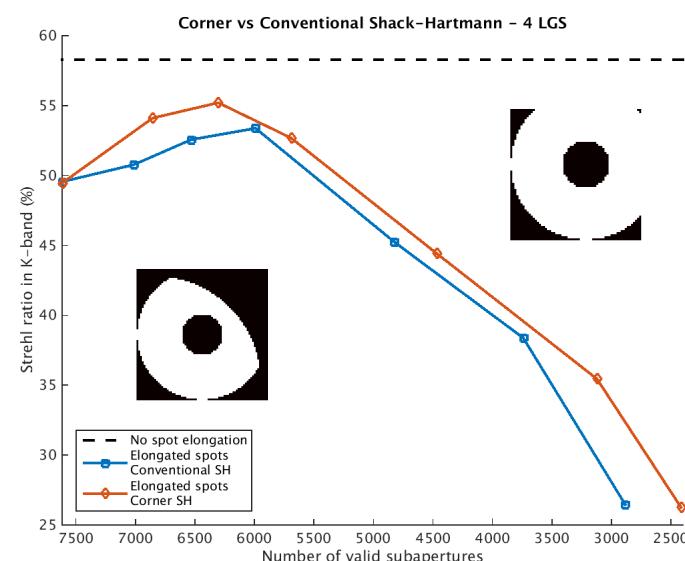
4 LGS configuration



Corner Shack-Hartmann



Corner SH always better
than conventional SH
for a given # of subapertures



Conclusion

- > Completely discarding 10-40% of the truncated subapertures actually increases Strehl Ratio (depending on Sodium profiles)
- > Combining subaperture differential weighting and cropping allow us to reach quasi-ideal performance on most sodium profiles
- > Help release constraints on Truth Sensor design: 8(x8) is enough !
- > Corner Shack-Hartmann to optimize pixel usage

On-going work

- Thorough analysis of more parameters (truncation, d/r0, central obscuration, asterism diameter, #LGS...)
- Advanced centroiding algorithms/strategies
- Actual ELT M4 geometry (see N. Schwartz talk tomorrow 12:30)
- Temporal aspects, sodium profil evolution during the night

Gracias



European Research Council

Established by the European Commission



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